



# KIRUNA GEOPHYSICAL DATA

DATA SUMMARY 06/01-03  
JANUARY-MARCH 2006

Collected at  
SWEDISH INSTITUTE OF SPACE PHYSICS  
KIRUNA, SWEDEN





# Kiruna Geophysical Data

Data Summary 06/01-03  
January – March 2006

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## INTRODUCTION

”Kiruna Geophysical Data” is a collection of results of ground-based more or less continuous measurements of upper atmosphere variables carried out at the Swedish Institute of Space Physics. Our objective with this data summary is to present current geophysical data, related to polar upper atmospheric conditions, to those who require it for their research work.

Coordinates of the Swedish Institute of Space Physics are:

Geogr.		Geomagn. Dipole		Corr. Geomagn.	
Lat.	Long.	Lat.	Long.	Lat.	Long.
67.8°N	20.4°E	65.2°N	116.0°E	64.8°N	104.2°E

The following ground-based measurements of upper atmospheric parameters have been conducted at the Swedish Institute of Space Physics in the period for which data is presented in this report.

Geomagnetism		Optical aurora	Ionospheric ionization
1.	Geomagnetic digital recording	1. All-sky camera	1. Riometer 30.0 MHz
			2. Riometer 38.2 MHz

The Ionospheric Observatory at Lycksele (64.7°N, 18.8°E) is also operated by the Swedish Institute of Space Physics. Magnetometer, riometer, and ionosonde measurements are conducted there.

Universal time is used in all tables.

Ionosonde data obtained at the Swedish Institute of Space Physics in Kiruna have been published together with ionosonde data from Lycksele and Uppsala in a monthly report ”Ionospheric Data Sweden” between January 1990 and March 2000. Copies are available from the Swedish Institute of Space Physics.

Scaled ionosonde parameters from 1965 – 1998 are also available from  
[http://www.wdc.rl.ac.uk/wdcc1/ionosondes/data\\_archive.html](http://www.wdc.rl.ac.uk/wdcc1/ionosondes/data_archive.html)

The most recent data from all instruments are accessible via internet at  
<http://www.irf.se/Data>

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## GEOMAGNETIC DATA

### Magnetograms

The magnetograms reproduced here are plots of data recorded by a three component digital fluxgate magnetometer with a sampling frequency of 1 sample/second.

The scaling is indicated by the tick mark values on the plots. Note that for large disturbances the normal scaling has been reduced. The positive direction of the components are: X north, Y east and Z down, in geographic coordinates.

### K-indices

The K-indices give a measure of the amount of disturbance of the geomagnetic field.

A reading of maximum deflection has been made during each 3-hour interval (starting from 00 to 03 UT) of two components X and Y. The Sq + L variation has been compensated for. A detailed description of the determination of K-indices has been given by J. Bartels: The technique of scaling indices K and Q geomagnetic activity. IGY Annals 4, London 1957.

The following scale has been used:

K-indices	Deflection in nT		
0	0	-	15
1	15	-	30
2	30	-	60
3	60	-	120
4	120	-	210
5	210	-	360
6	360	-	600
7	600	-	990
8	990	-	1500
9	1500		and more



## **ALL-SKY CAMERA DATA**

Digital all-sky camera images are typically recorded in jpeg format at a rate of one frame per minute. The filename of each image contains information about UTC and exposure time. Longer exposure times are used at regular intervals (typically eight minutes). Data might be recorded at higher frame rates during selected periods. More information and image data are available from: <http://www.irf.se/allsky>

Information about available records of all-sky camera data is presented as colour keograms. Keograms are time sequences of poleward to equatorward slices through the centre of all-sky images. The top of the image points poleward, the bottom equatorward. Time is on the horizontal axis (0:00-24:00 UTC). Each page contains keograms for seven days. White areas indicate missing data (e.g. daylight conditions).

## **IONOSPHERIC ABSORPTION DATA**

Ionospheric absorption at 30.0 MHz and 38.2 MHz is recorded by two La Jolla riometers. The riometer records of cosmic noise power are reproduced in this report. The zero power level is at bottom of each panel.

The absorption (in decibels) can be computed from the formula

$$A(\text{dB}) = 10 \log_{10} \frac{P_o}{P}$$

where  $P$  is the received cosmic noise power,  $P_o$  the cosmic noise power expected under ionospherically quiet conditions at the same sidereal time.



## GEOMAGNETIC K-INDICES, KIRUNA

January 2006

Lower limit of K=9 is 1500 nT

Time in UT

Day	00 -	03 -	06 -	09 -	12 -	15 -	18 -	21 -	24	Sum
1	5	1	0	1	2	2	2	3		16
2	0	1	1	2	1	2	2	4		13
3	1	0	0	0	1	2	2	1		7
4	0	0	0	0	0	0	2	0		2
5	0	0	0	1	1	1	2	3		8
6	3	2	1	1	3	4	2	3		19
7	3	1	1	1	1	3	0	2		12
8	0	0	0	1	0	1	0	2		4
9	0	0	0	0	0	0	0	0		0
10	0	0	0	0	0	0	1	1		2
11	0	0	0	1	1	1	2	1		6
12	2	1	1	0	0	0	2	2		8
13	1	0	0	1	0	2	2	1		7
14	0	1	0	0	0	0	3	1		5
15	2	1	0	0	0	2	3	1		9
16	0	1	2	3	3	4	5	6		24
17	5	3	2	2	2	1	2	2		19
18	2	2	2	2	1	2	3	4		18
19	2	0	1	1	1	2	2	5		14
20	2	0	0	1	1	3	3	2		12
21	2	2	1	1	2	0	0	0		8
22	0	0	1	1	1	1	0	0		4
23	1	3	4	3	3	3	3	6		26
24	5	2	2	2	2	0	0	0		13
25	0	1	1	1	1	1	6	7		18
26	7	6	2	3	2	7	7	6		40
27	5	2	2	2	3	2	3	4		23
28	3	2	1	2	1	1	5	3		18
29	1	1	1	0	1	0	0	1		5
30	1	0	0	0	0	0	1	0		2
31	0	0	0	0	1	2	0	1		4
Sum	53	33	26	33	35	49	65	72		

## GEOMAGNETIC K-INDICES, KIRUNA

February 2006

Lower limit of K=9 is 1500 nT

Time in UT

Day	00 -	03 -	06 -	09 -	12 -	15 -	18 -	21 -	24	Sum
1	0	0	0	1	1	1	0	3		6
2	2	0	0	1	0	3	4	4		14
3	1	1	0	1	0	1	6	3		13
4	5	2	1	1	1	1	0	1		12
5	1	0	1	1	1	0	0	0		4
6	1	2	1	2	4	3	6	6		25
7	3	0	1	1	1	1	3	4		14
8	4	0	1	1	0	0	1	0		7
9	0	0	1	1	1	0	2	2		7
10	0	1	0	1	0	0	3	4		9
11	2	2	1	1	1	3	1	1		12
12	2	3	1	1	1	1	0	0		9
13	0	0	0	0	0	1	2	2		5
14	0	0	0	0	0	0	0	0		0
15	0	1	2	2	2	1	7	6		21
16	5	3	3	2	1	1	4	4		23
17	2	2	2	2	1	1	0	3		13
18	1	0	0	0	1	0	0	0		2
19	2	2	2	2	2	2	6	4		22
20	4	2	2	4	6	5	6	5		34
21	4	4	2	3	5	5	4	6		33
22	6	3	3	3	3	4	3	3		28
23	2	2	1	2	1	1	1	3		13
24	3	2	2	2	2	0	0	0		11
25	2	1	1	1	0	0	0	0		5
26	1	1	1	1	2	4	3	6		19
27	3	1	1	0	1	0	1	4		11
28	0	0	0	2	2	1	4	4		13
Sum	56	35	30	39	40	40	67	78		

## GEOMAGNETIC K-INDICES, KIRUNA

March 2006

Lower limit of K=9 is 1500 nT

Time in UT

Day	00 -	03 -	06 -	09 -	12 -	15 -	18 -	21 -	24	Sum
1	4	2	1	2	3	3	0	1		16
2	1	0	1	0	1	0	1	3		7
3	2	0	1	1	0	1	0	1		6
4	1	1	1	1	0	0	0	2		6
5	1	1	1	1	0	0	0	2		6
6	0	1	0	2	2	4	6	6		21
7	5	5	1	3	2	1	4	3		24
8	1	1	1	1	2	0	3	4		13
9	1	0	0	1	1	1	4	2		10
10	1	1	2	2	2	3	2	8		21
11	6	2	2	3	2	2	2	4		23
12	0	0	1	1	1	1	2	1		7
13	0	1	1	1	0	0	1	2		6
14	0	0	1	2	1	1	0	2		7
15	3	1	1	2	2	3	4	1		17
16	1	2	1	1	2	3	2	2		14
17	0	0	0	1	1	0	1	1		4
18	0	1	3	4	4	5	7	7		31
19	7	6	4	4	4	4	6	6		41
20	5	5	4	4	3	3	7	5		36
21	3	4	2	2	4	7	5	6		33
22	5	4	2	2	2	2	5	5		27
23	3	3	2	1	1	1	0	2		13
24	1	1	1	2	1	2	4	5		17
25	4	2	1	2	1	2	5	1		18
26	0	1	1	2	2	4	5	5		20
27	6	4	1	1	2	3	1	6		24
28	1	2	1	1	2	2	2	2		13
29	1	1	1	1	1	1	4	5		15
30	2	1	1	1	1	1	5	4		16
31	1	1	1	2	2	3	4	2		16
Sum	66	54	41	54	52	63	92	106		



# **ALL-SKY CAMERA KEOGRAMS**

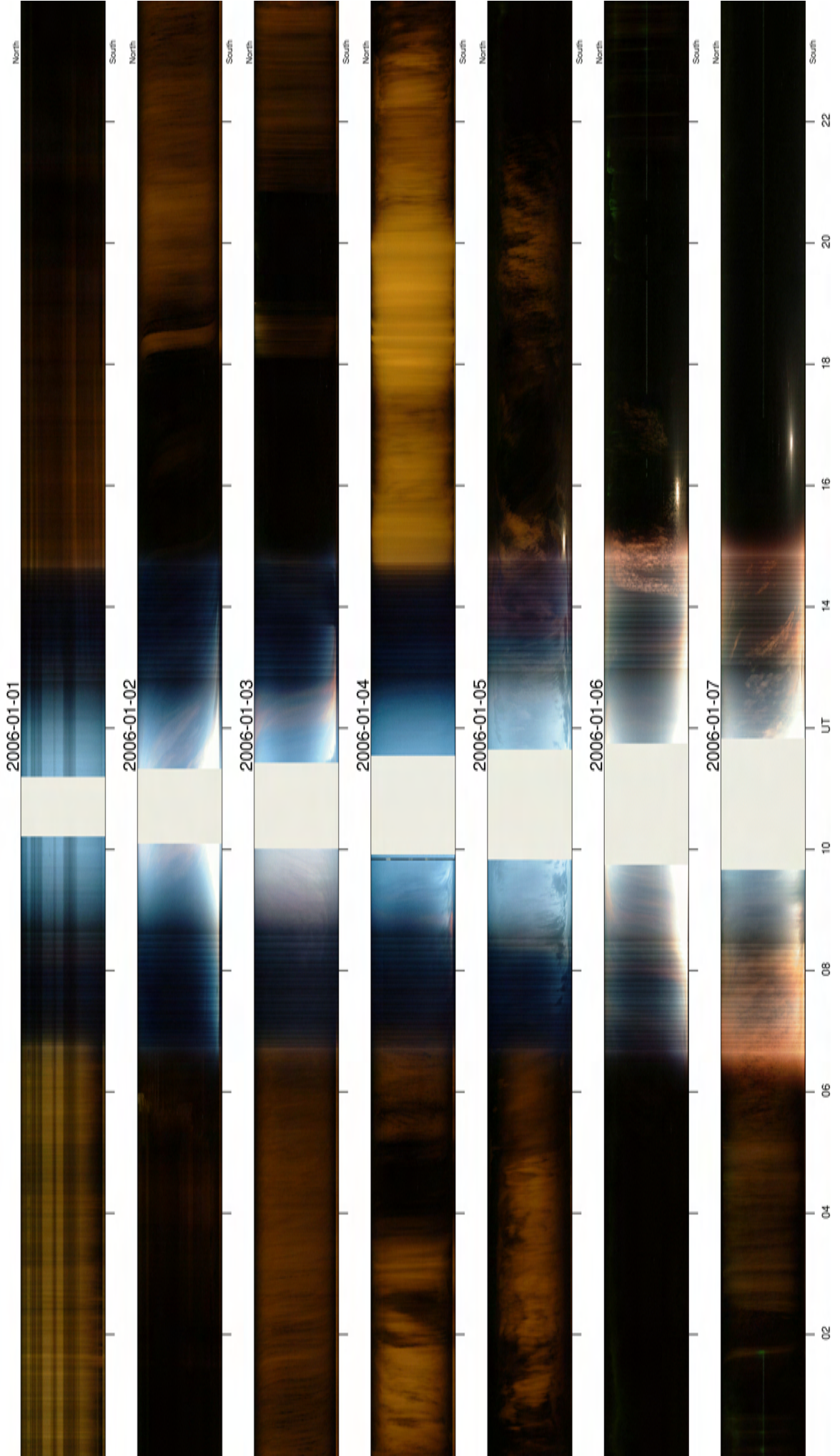
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<http://www.irf.se/allsky/kgd>

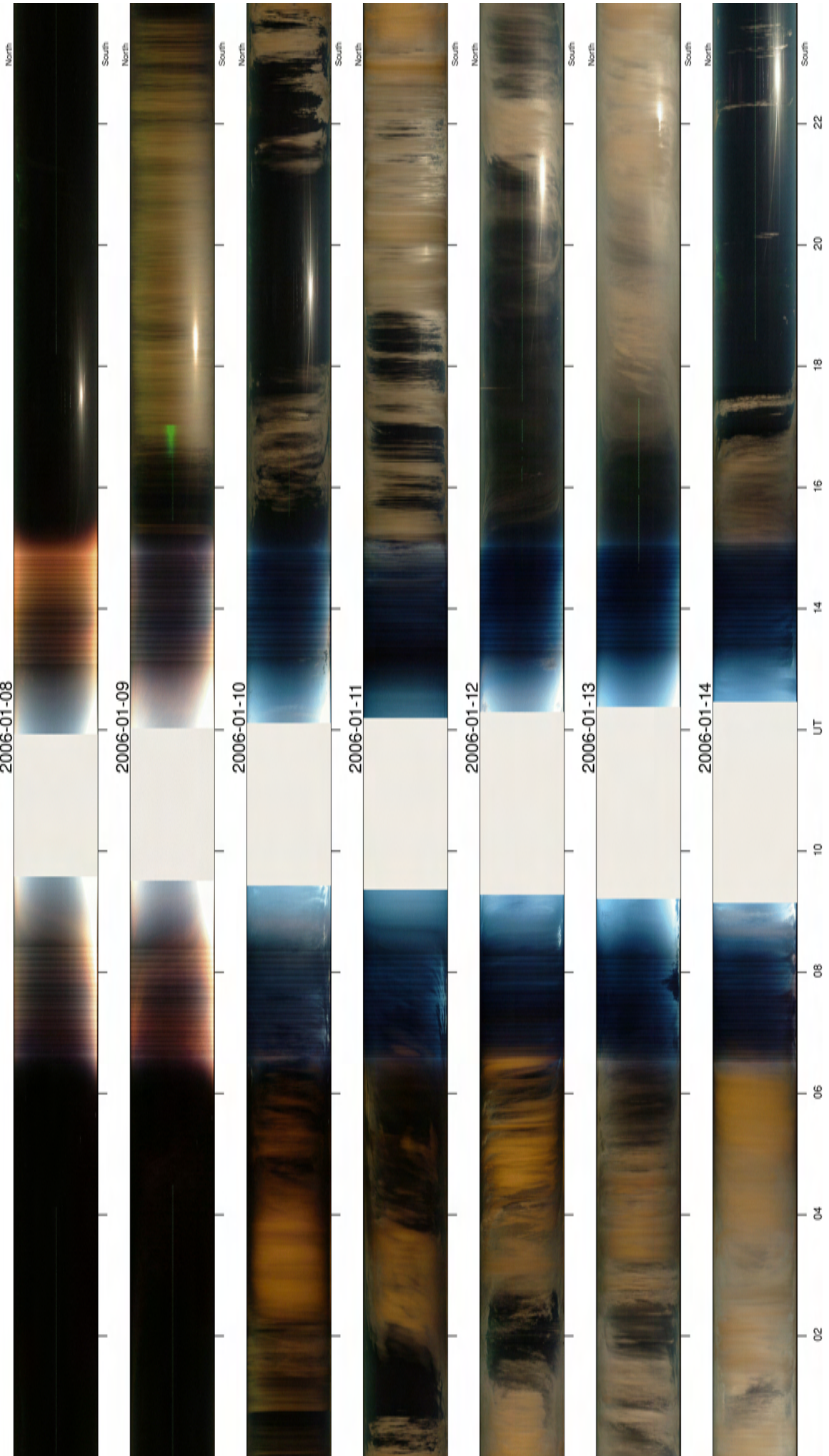




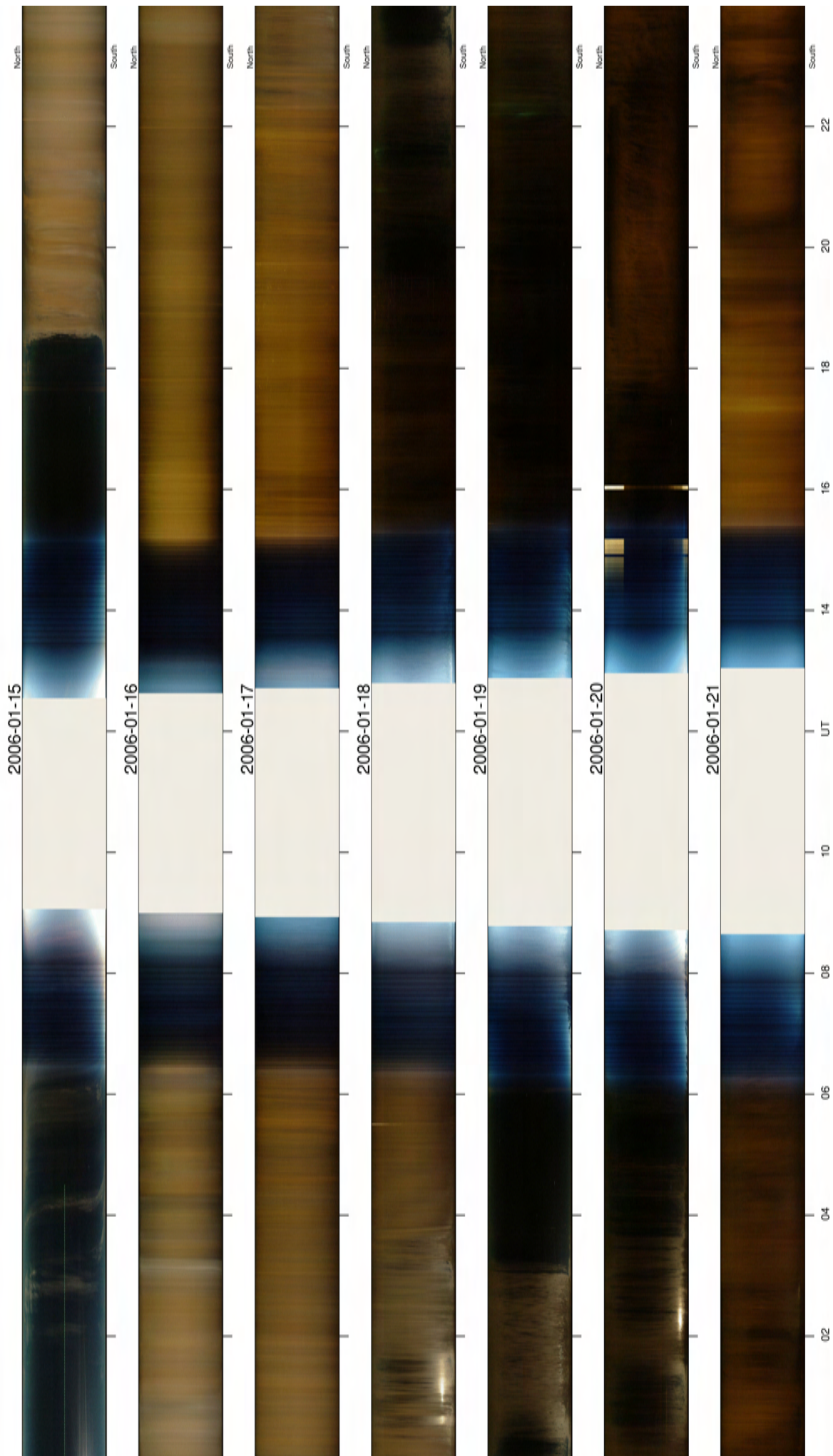
# Kiruna keogram



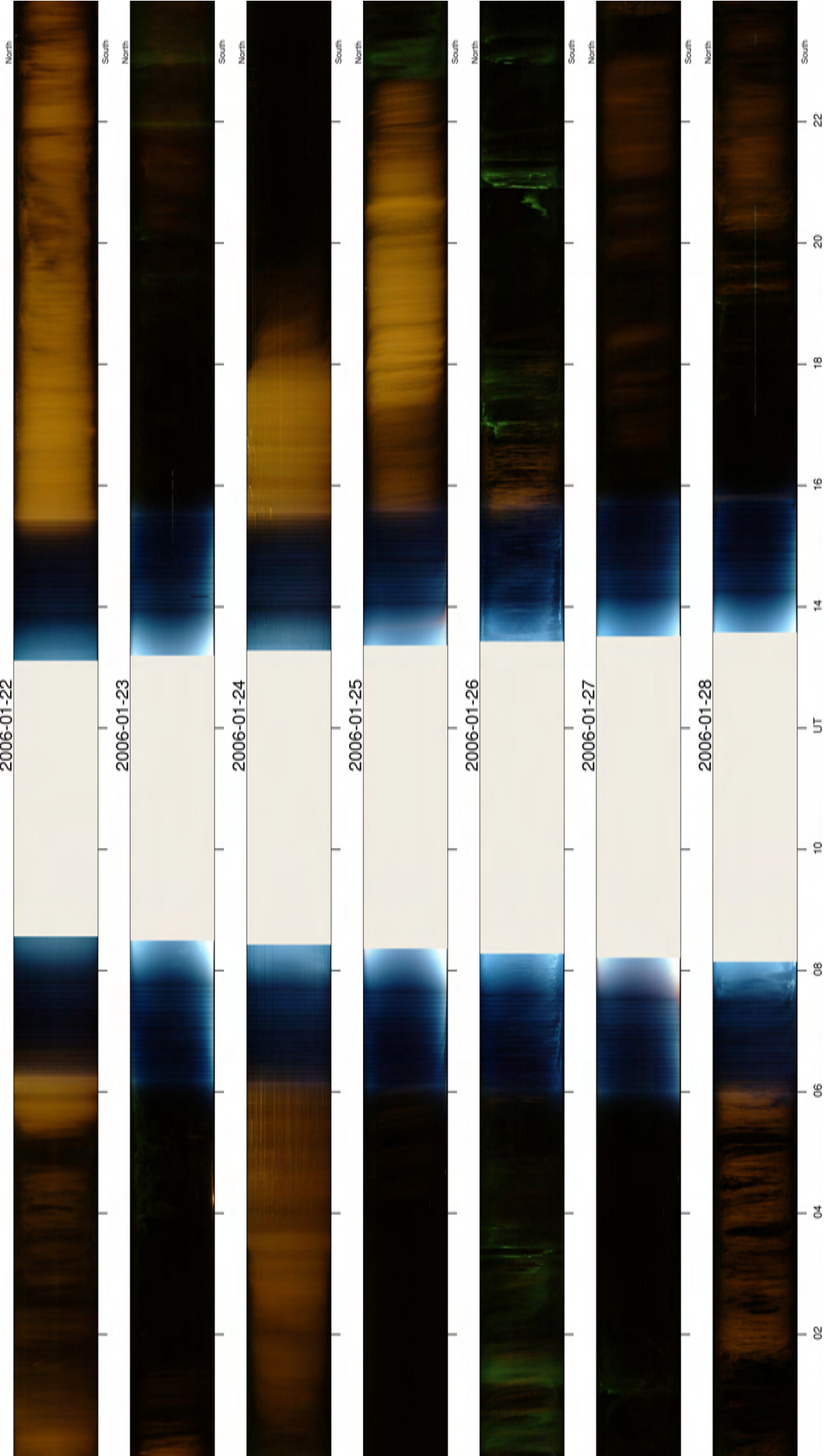
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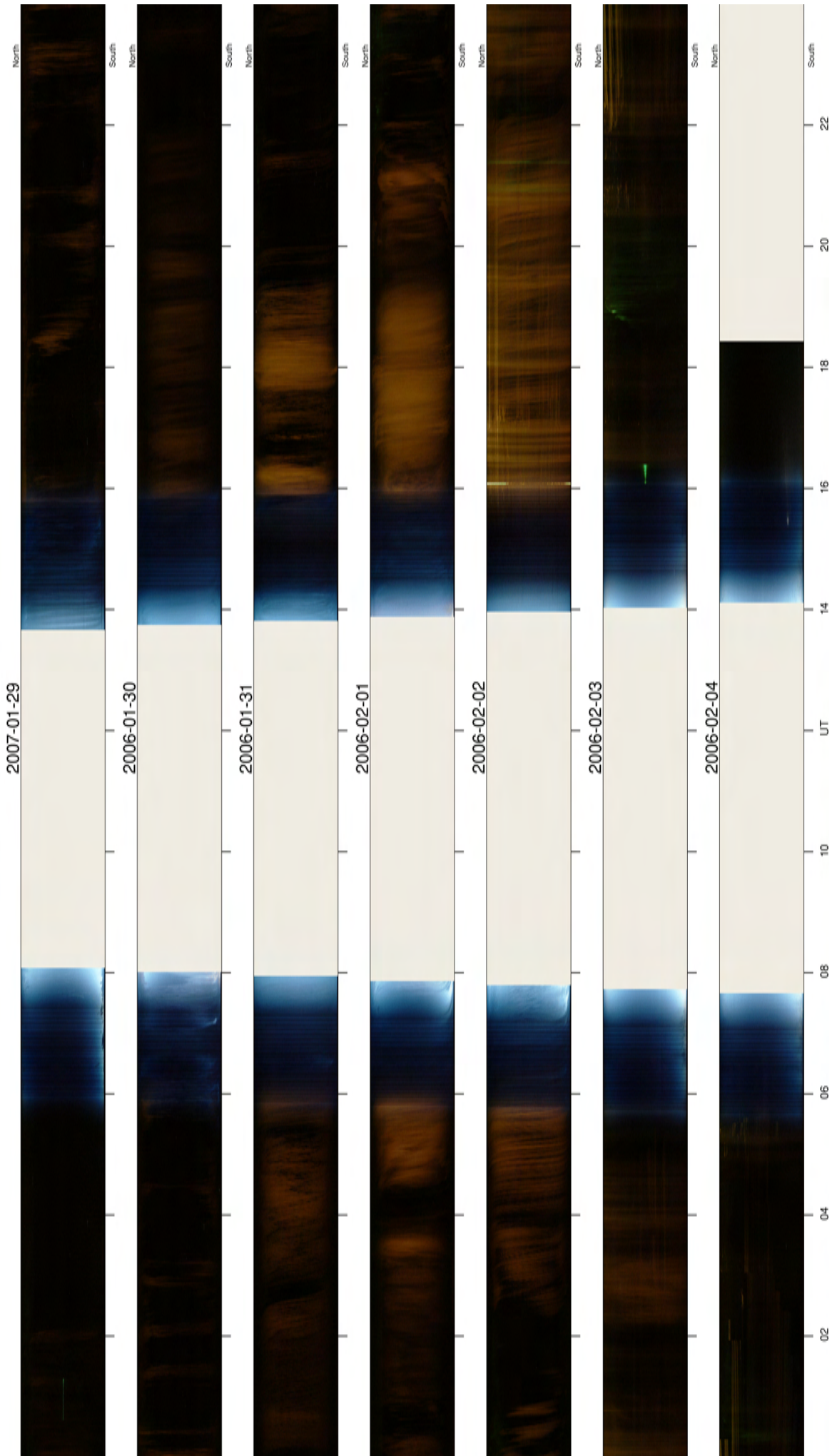
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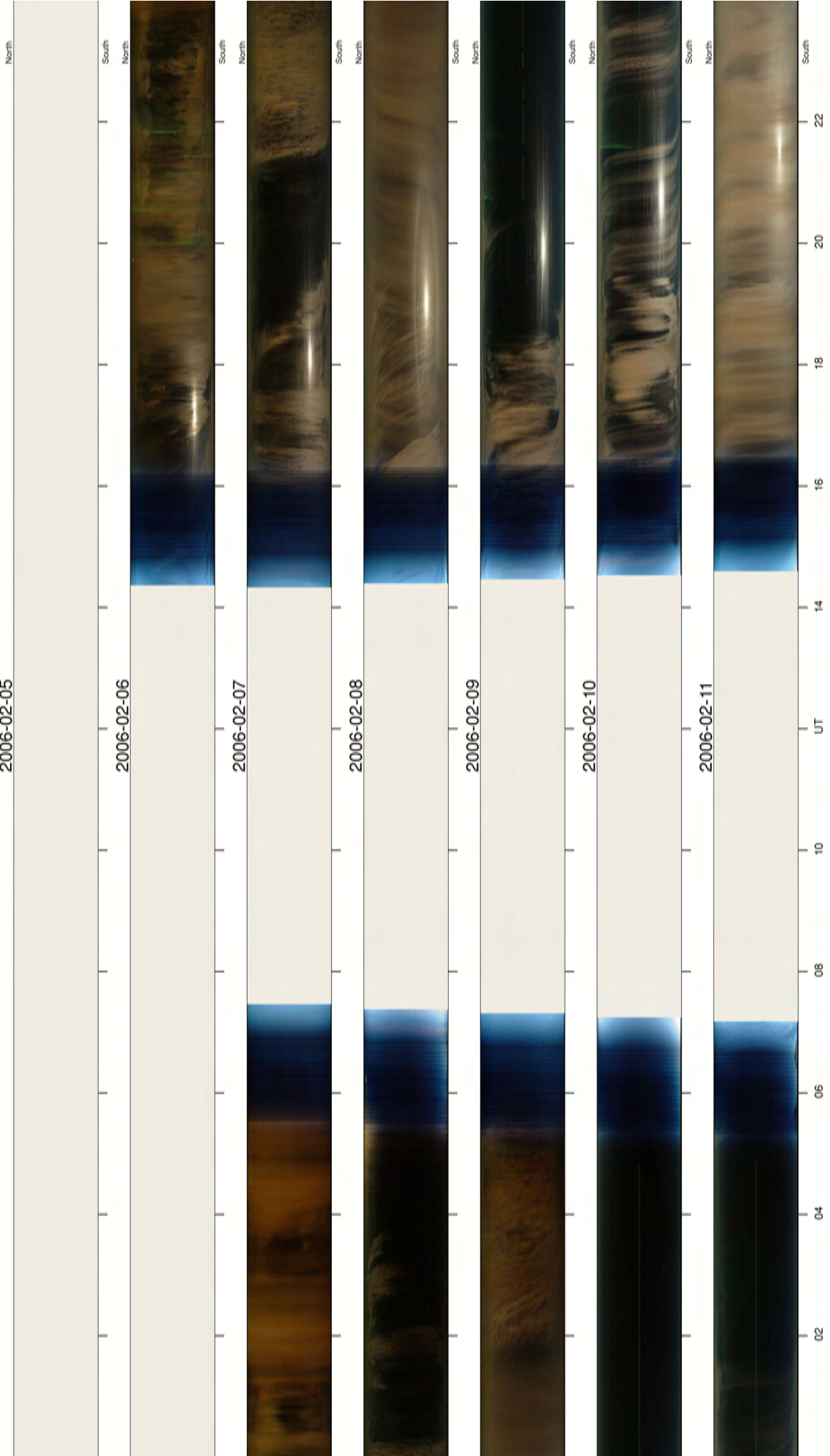
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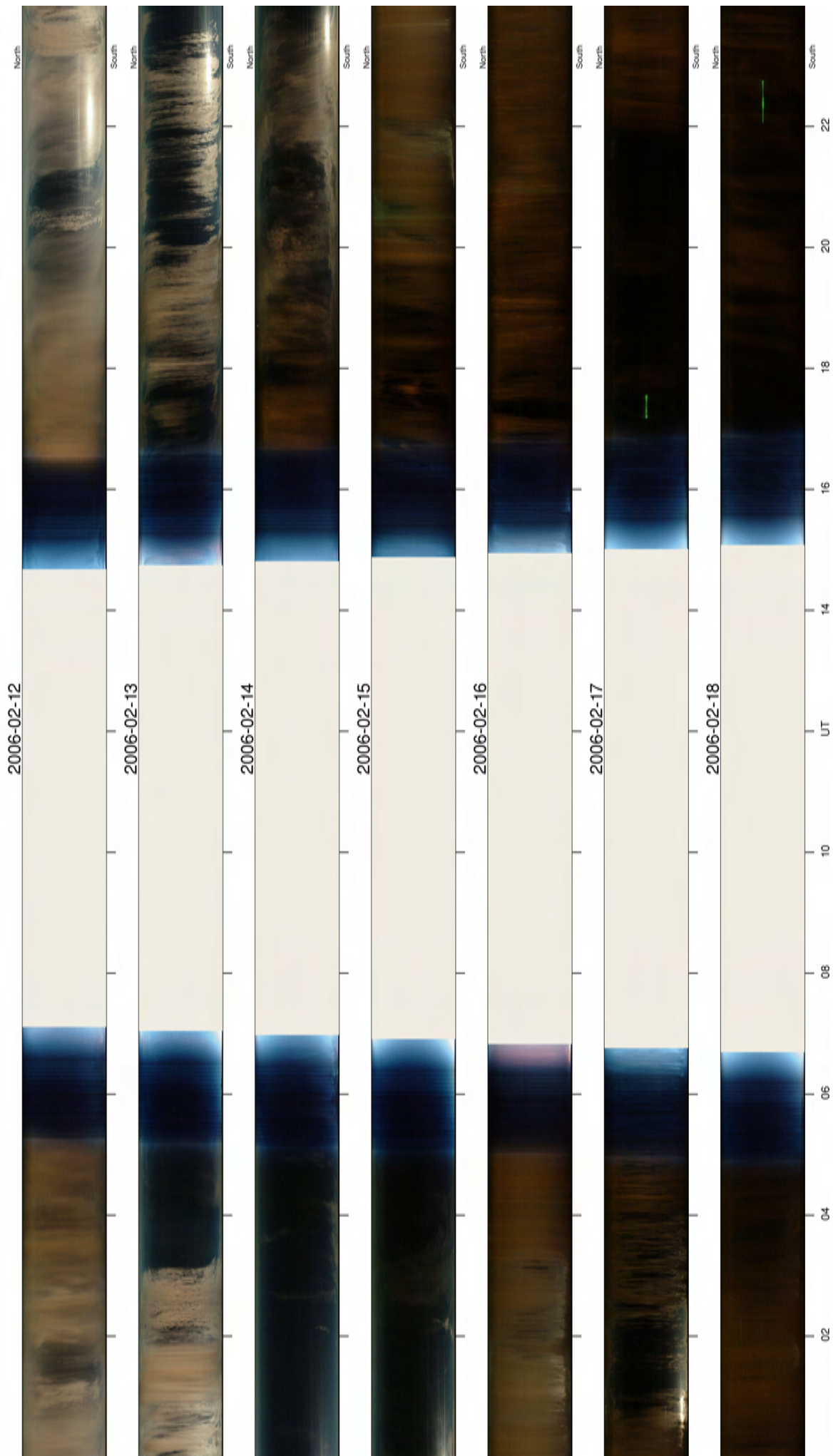
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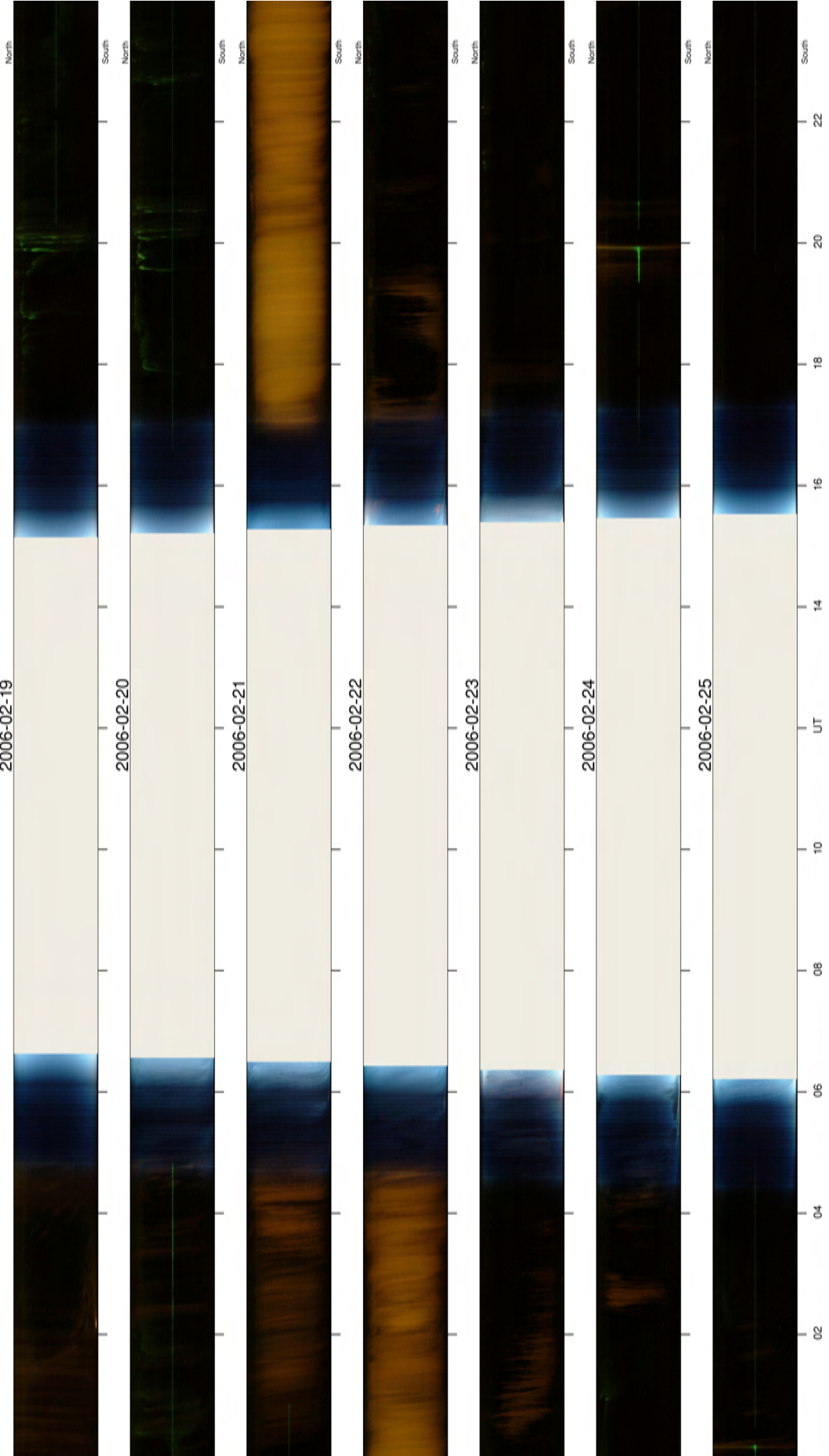
# Kiruna keogram



# Kiruna keogram

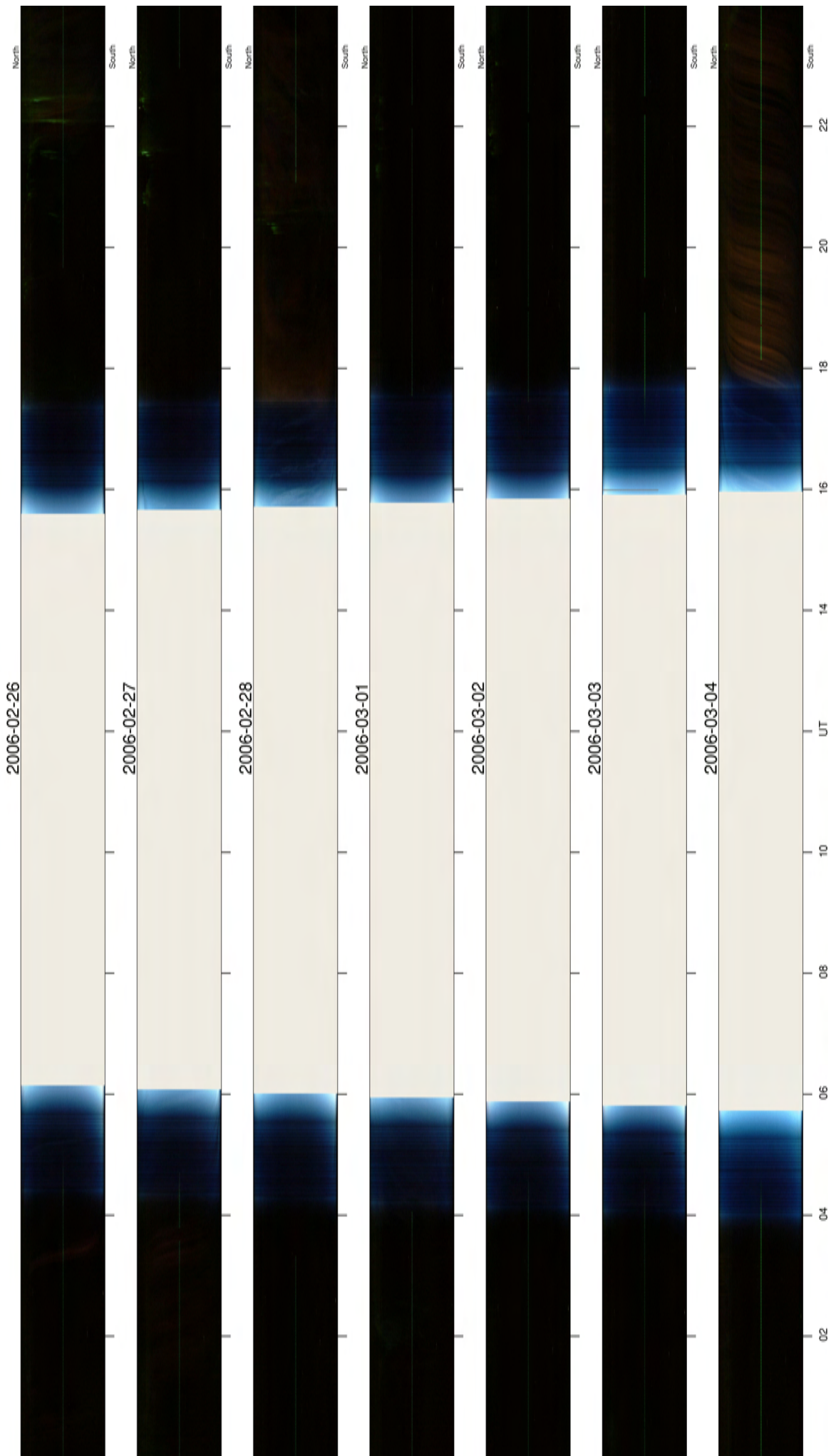


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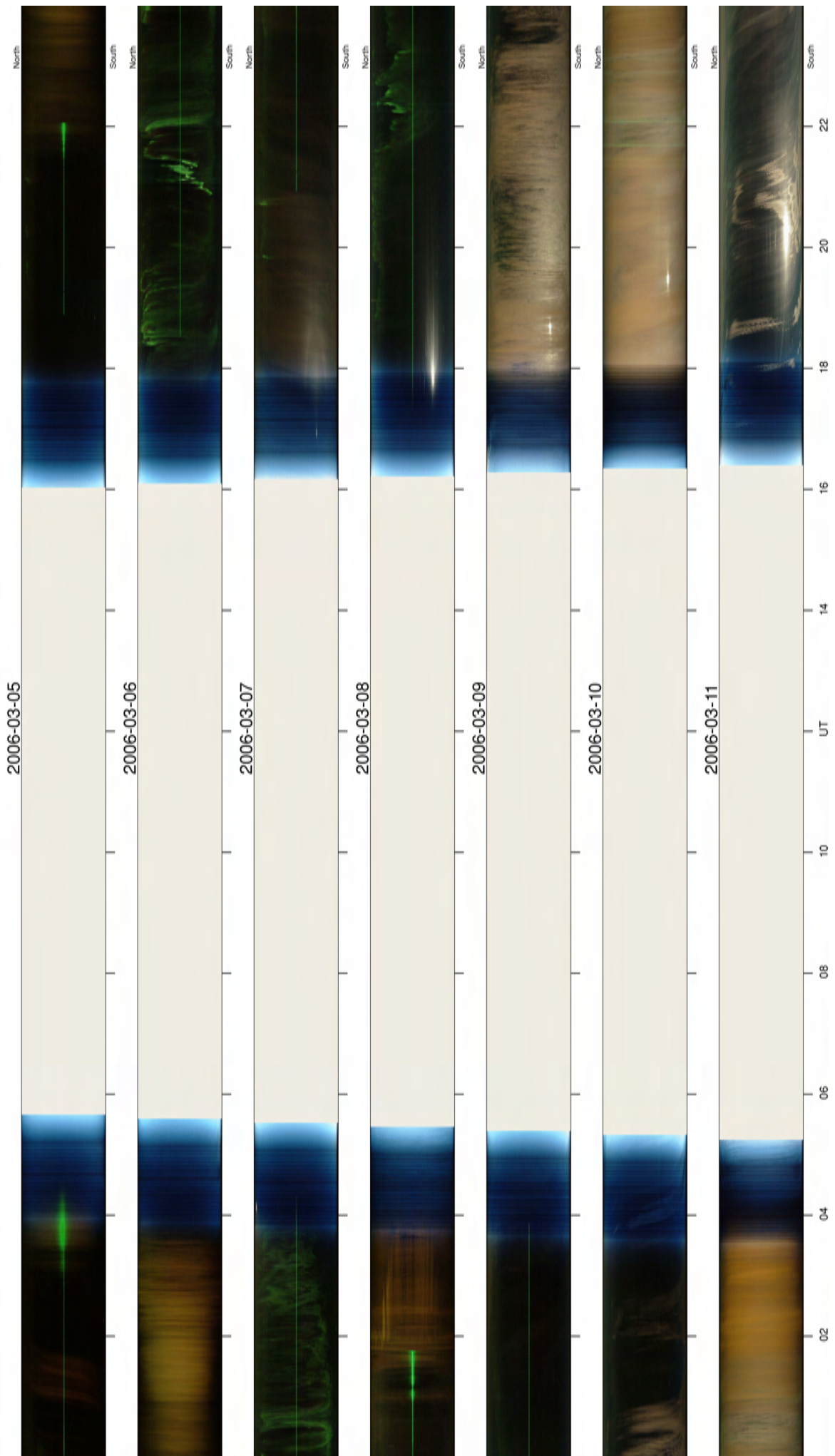




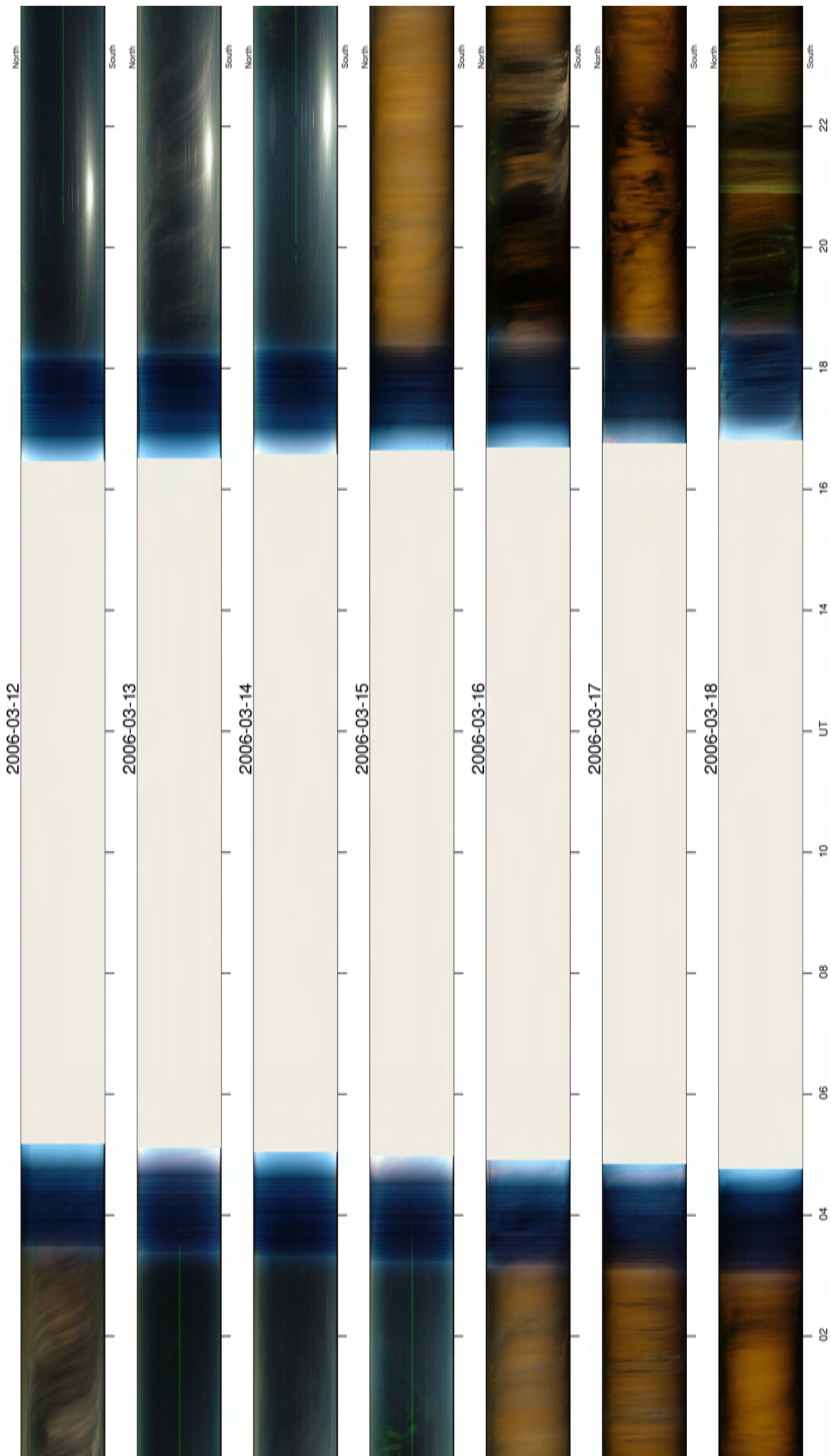
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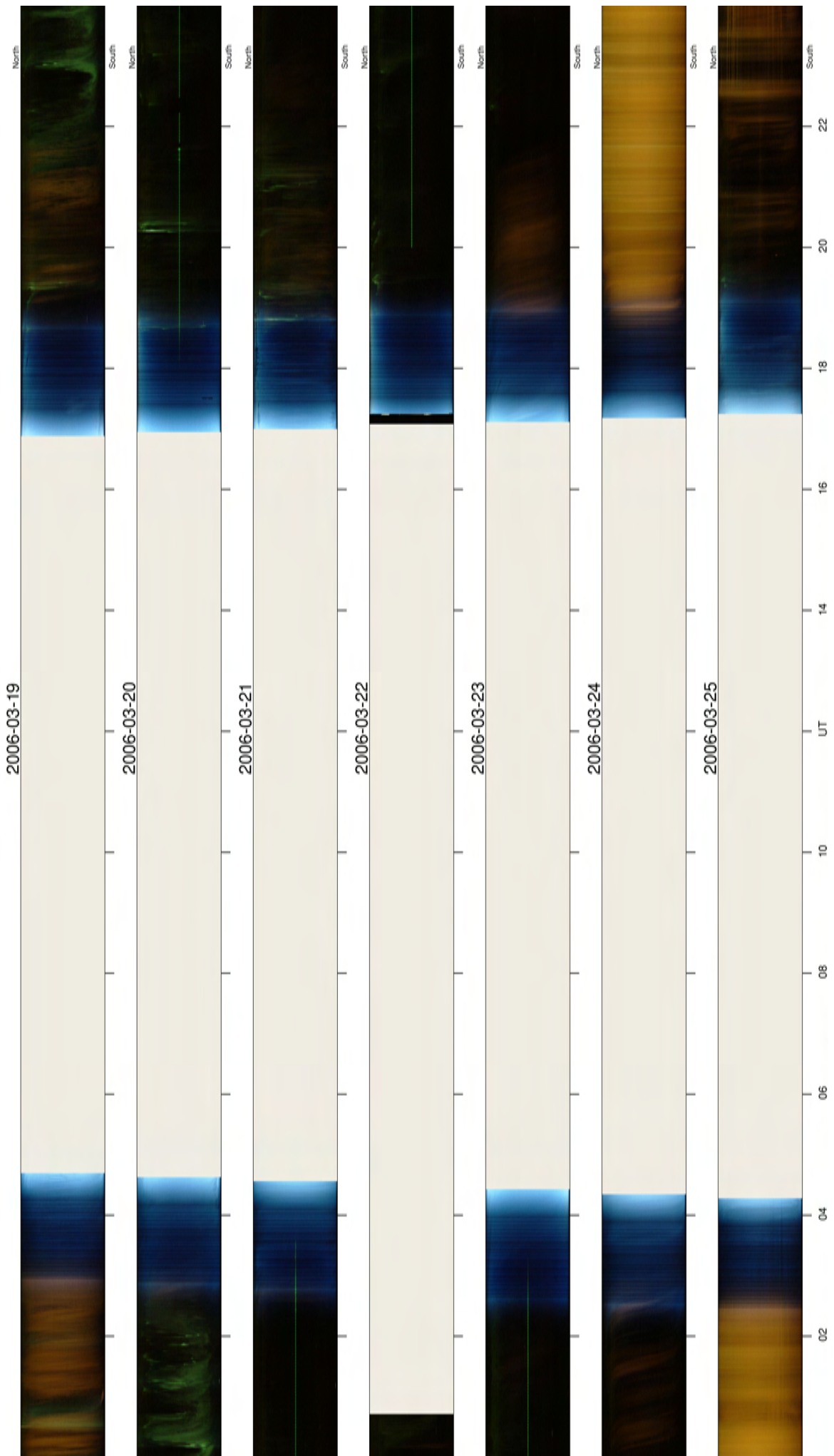
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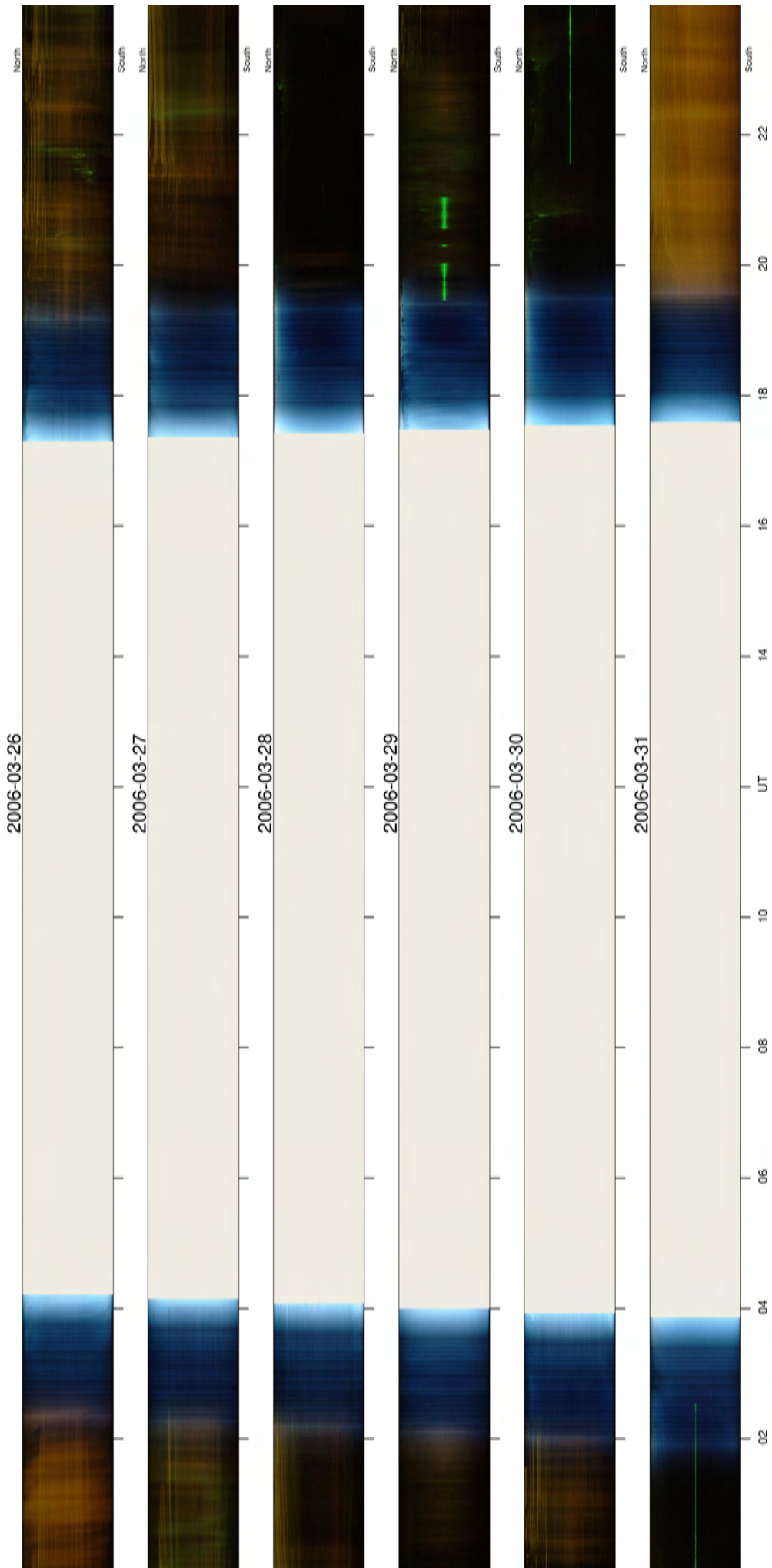
# Kiruna keogram



# Kiruna keogram



# Kiruna keogram





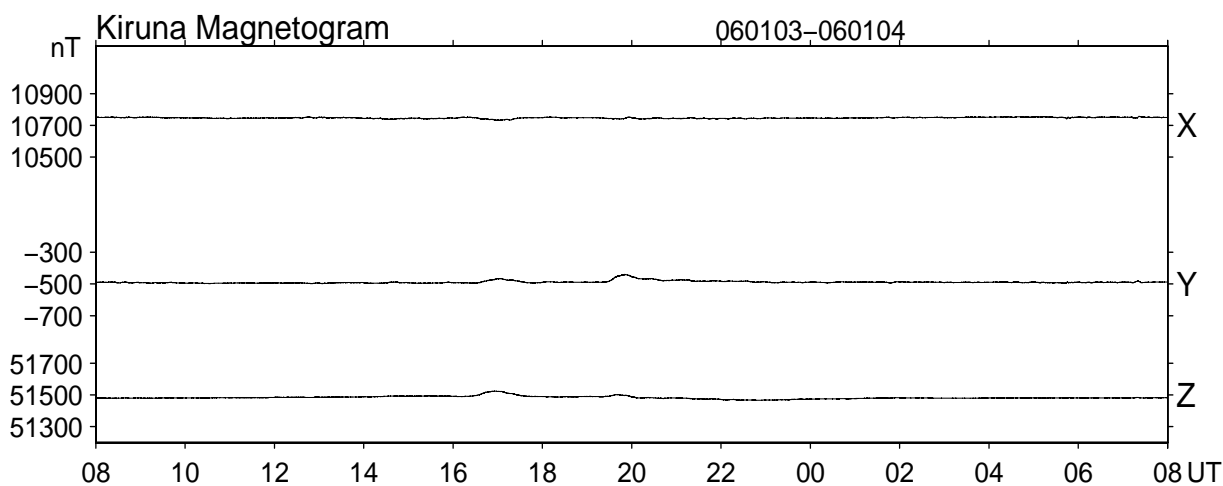
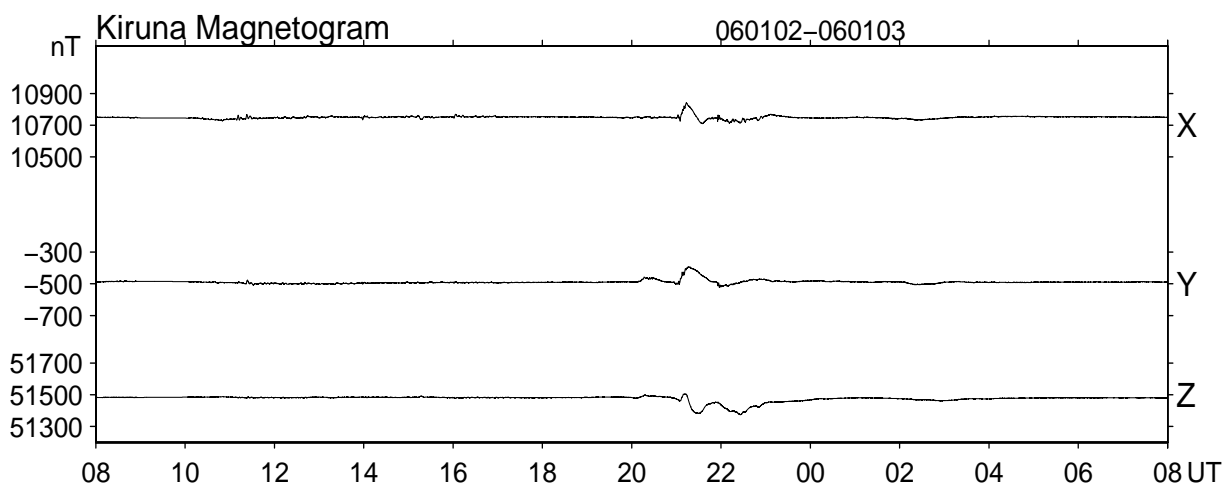
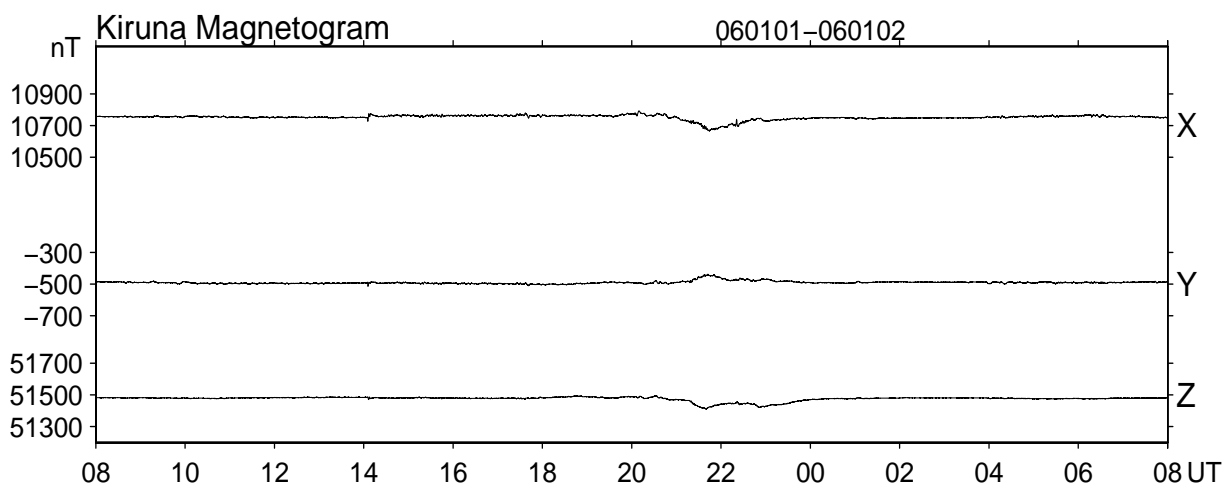
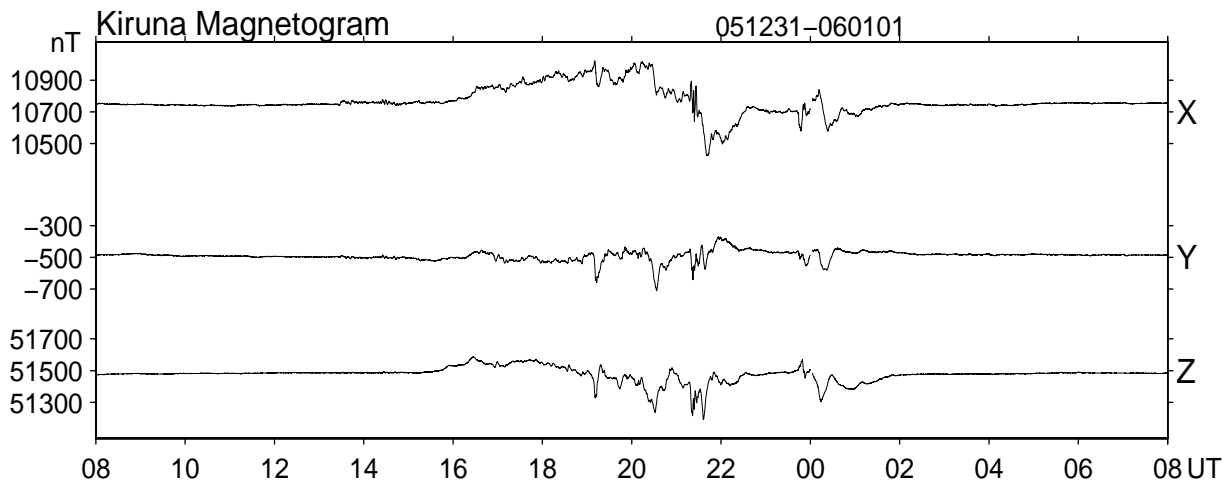
# **KIRUNA MAGNETOGRAM PLOTS**

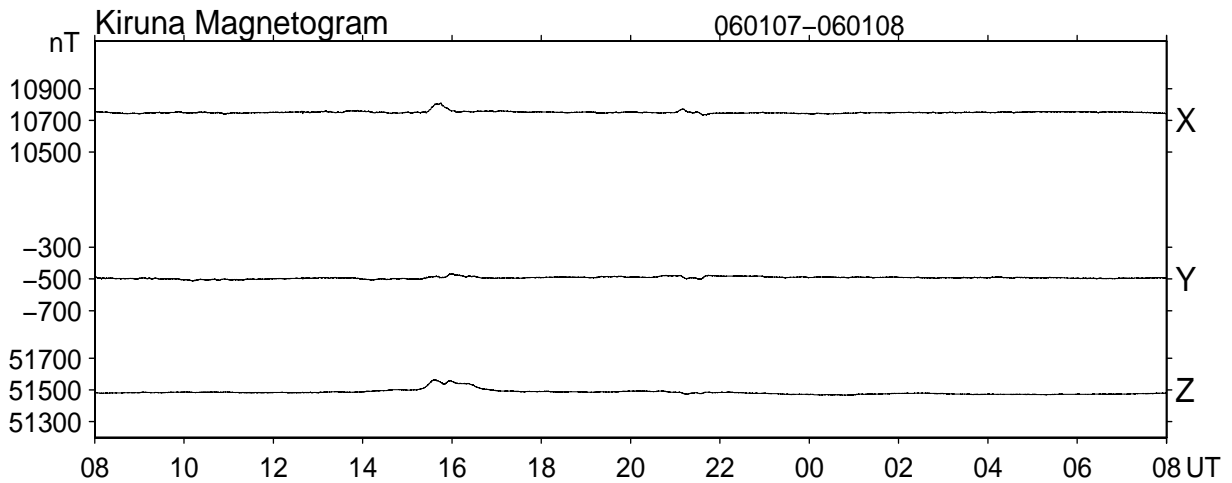
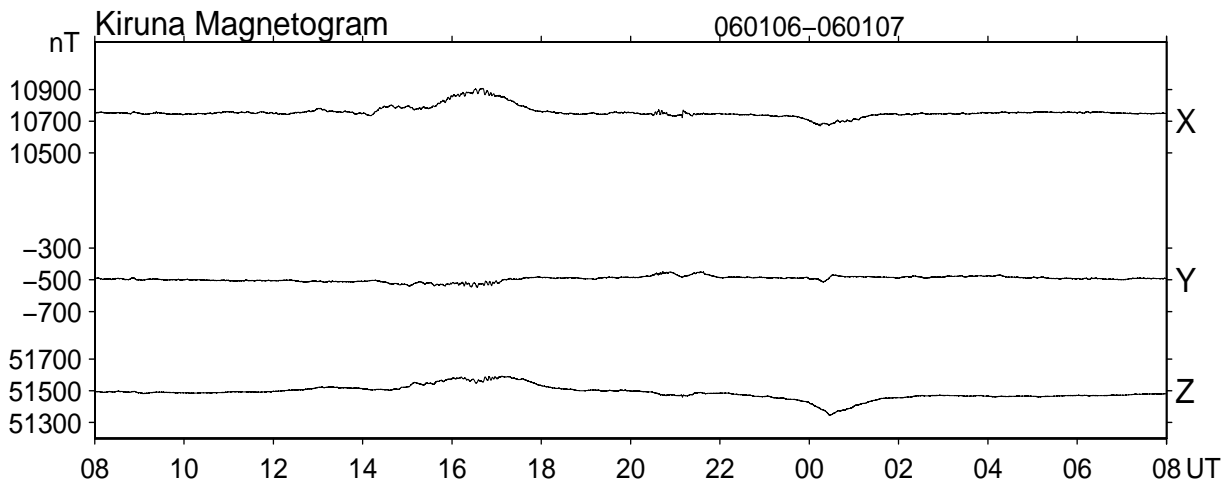
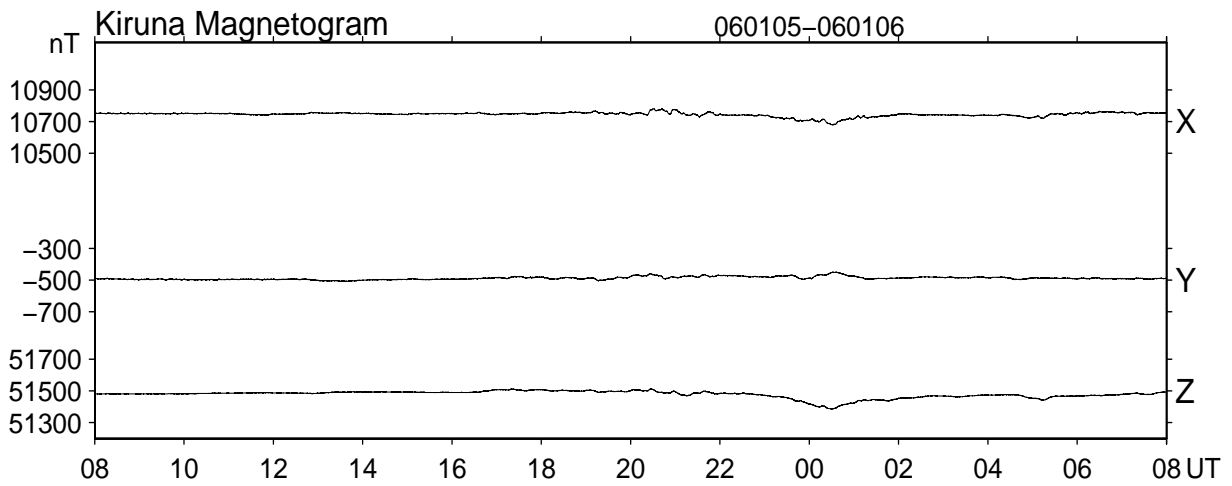
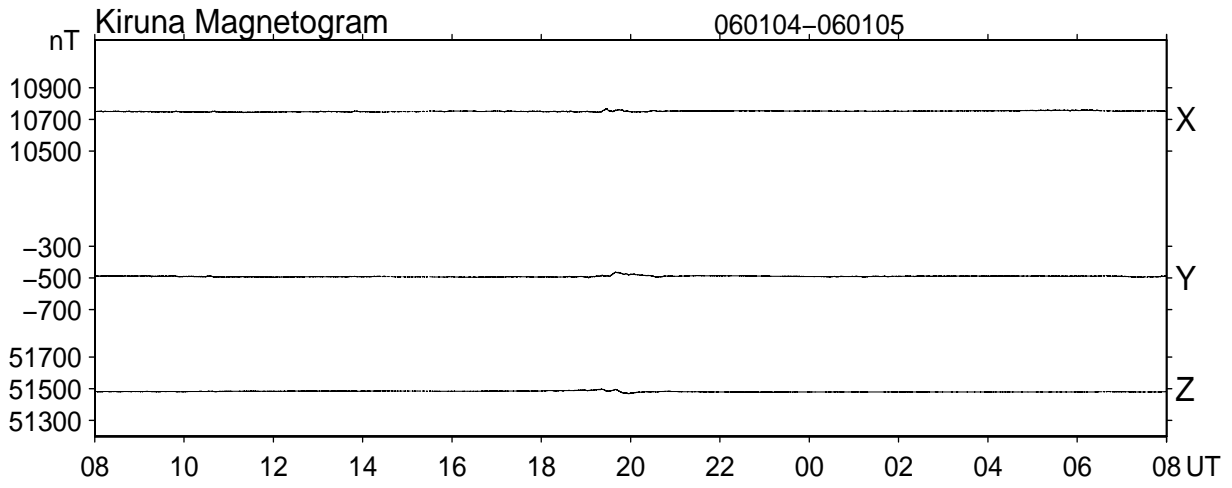
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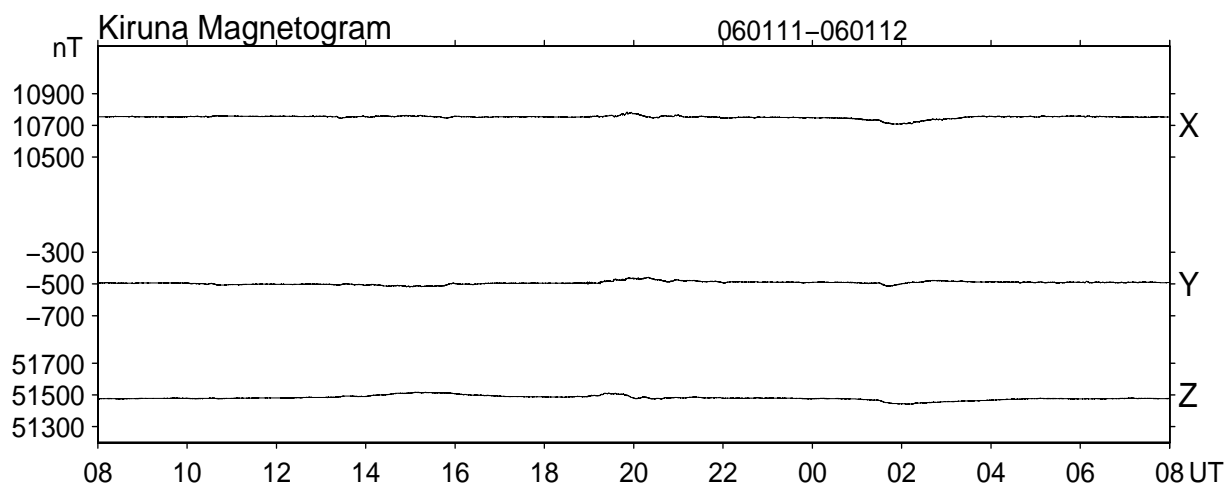
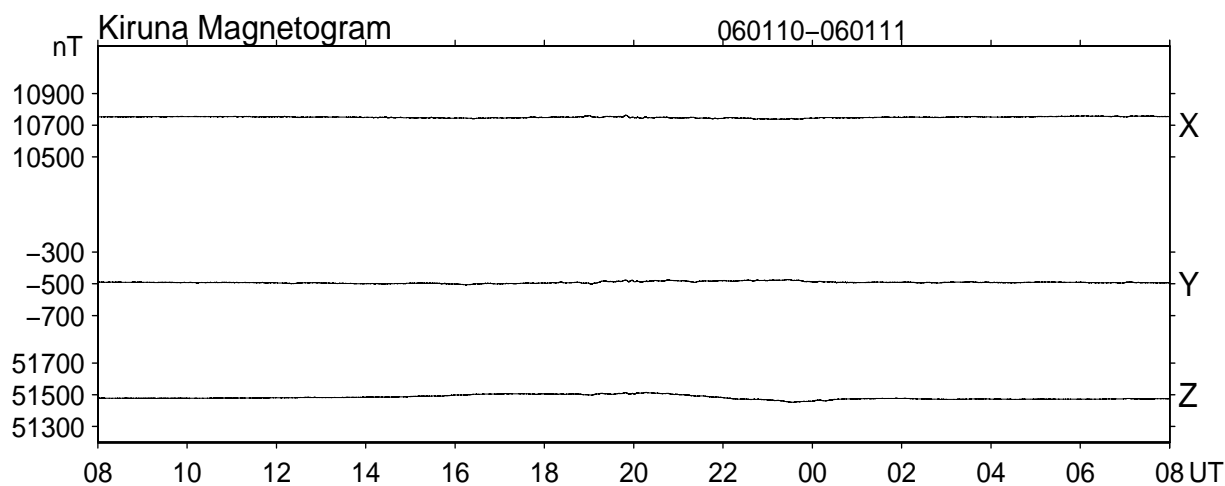
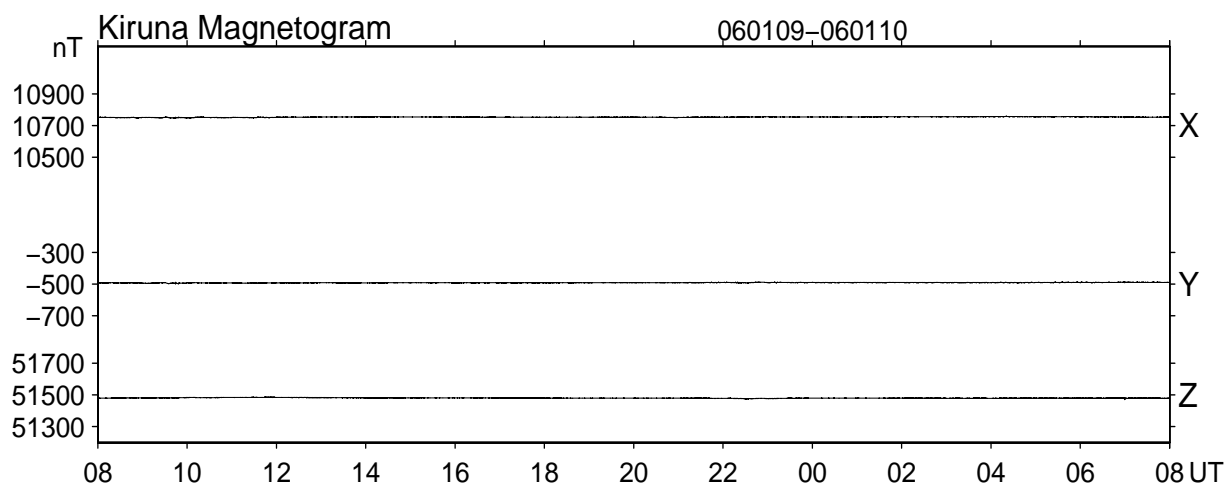
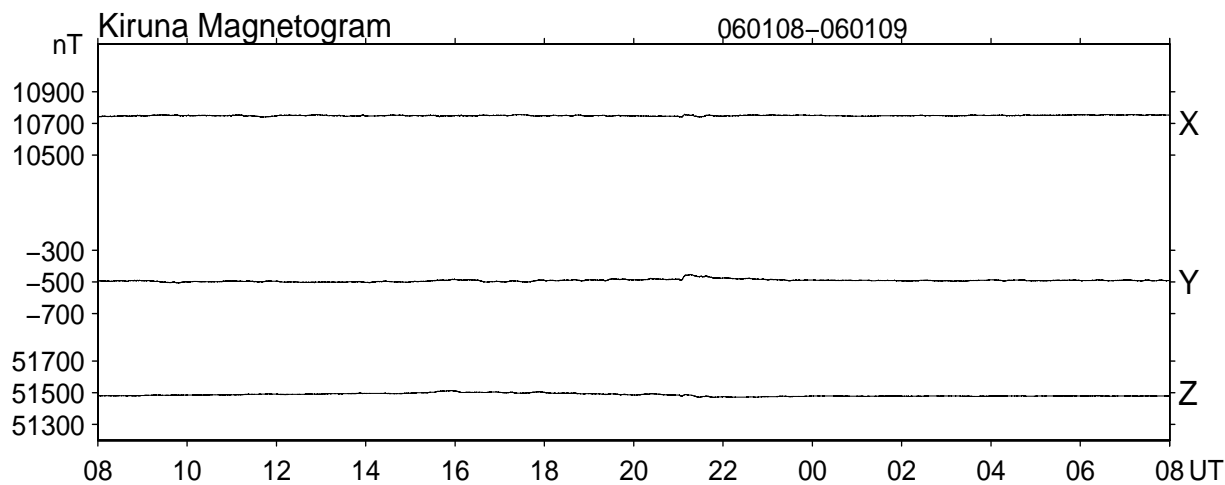
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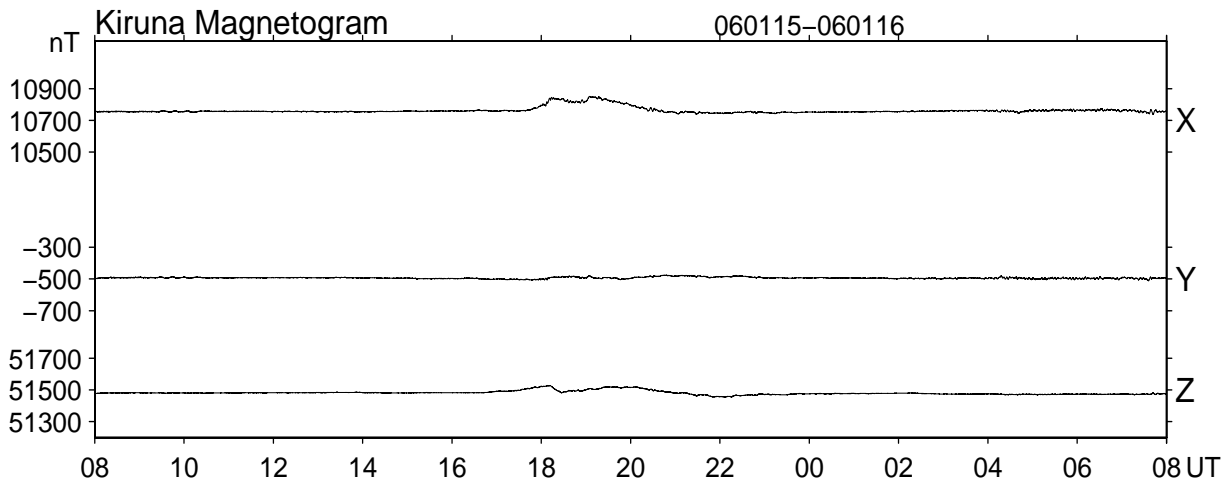
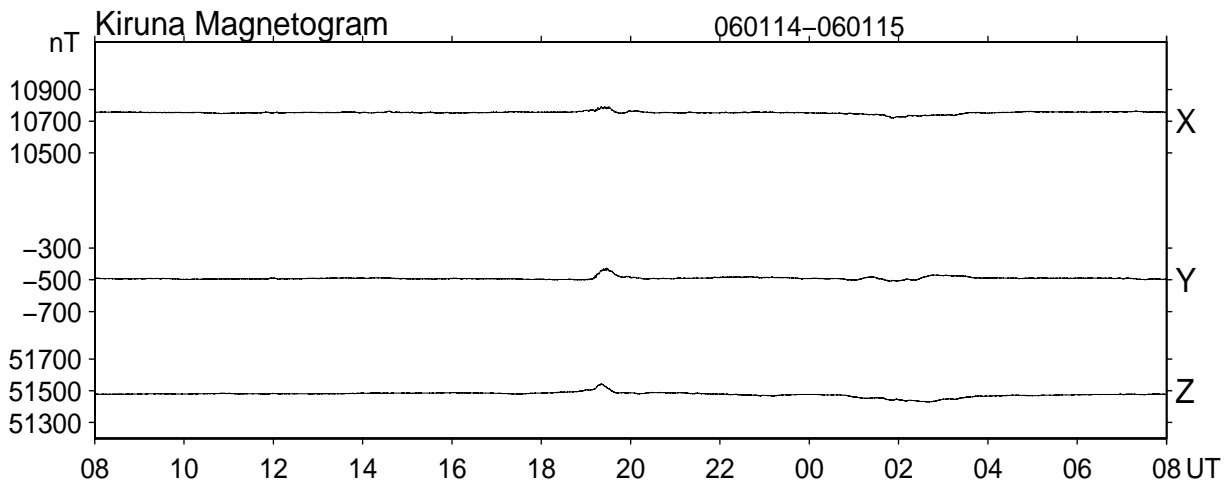
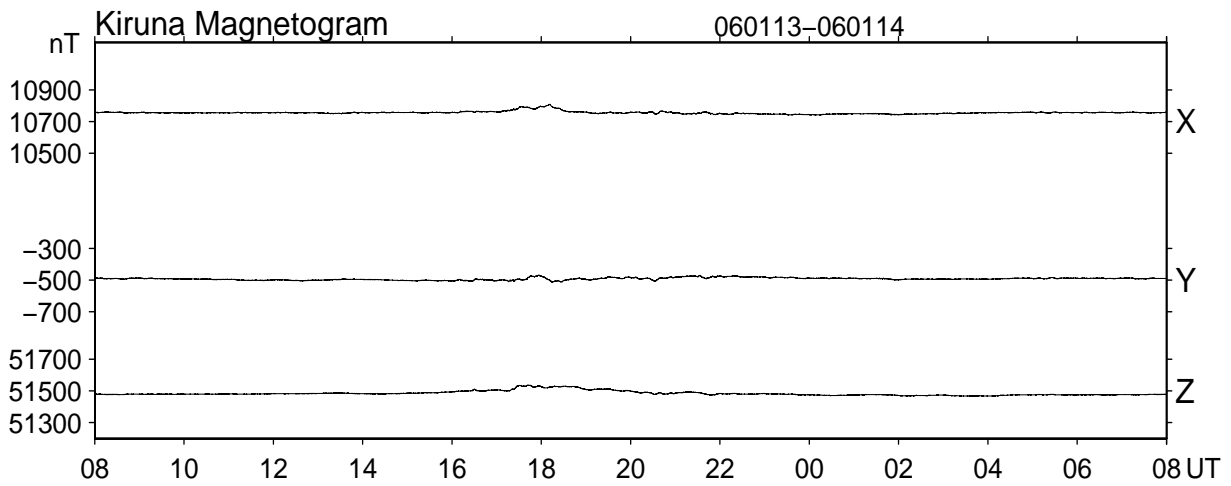
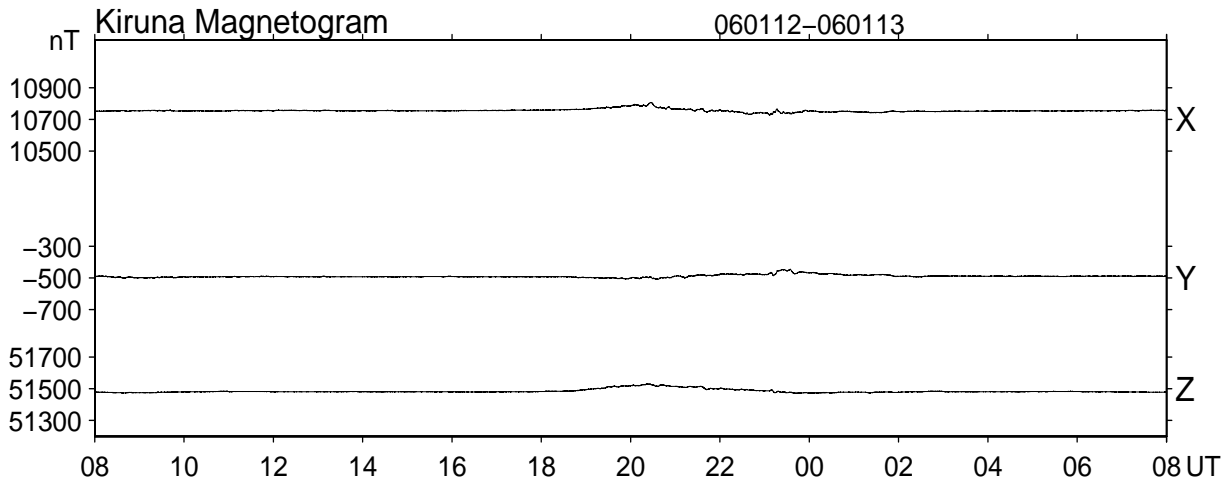


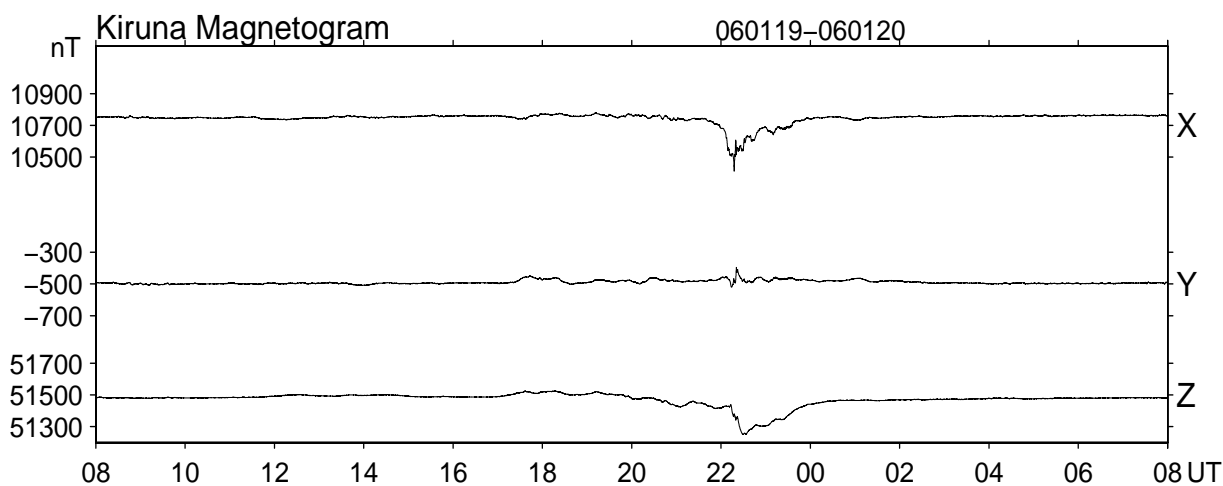
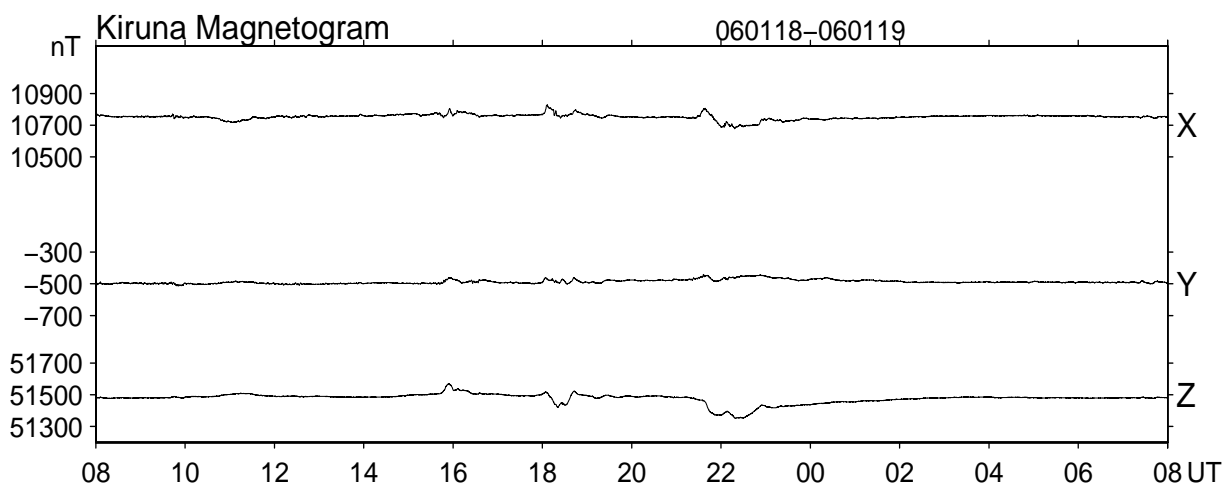
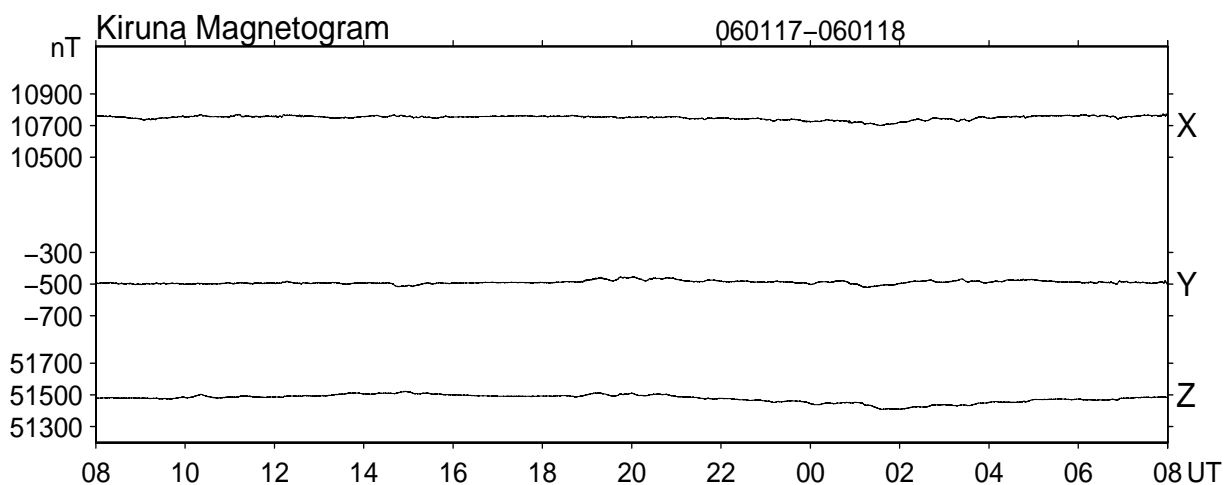
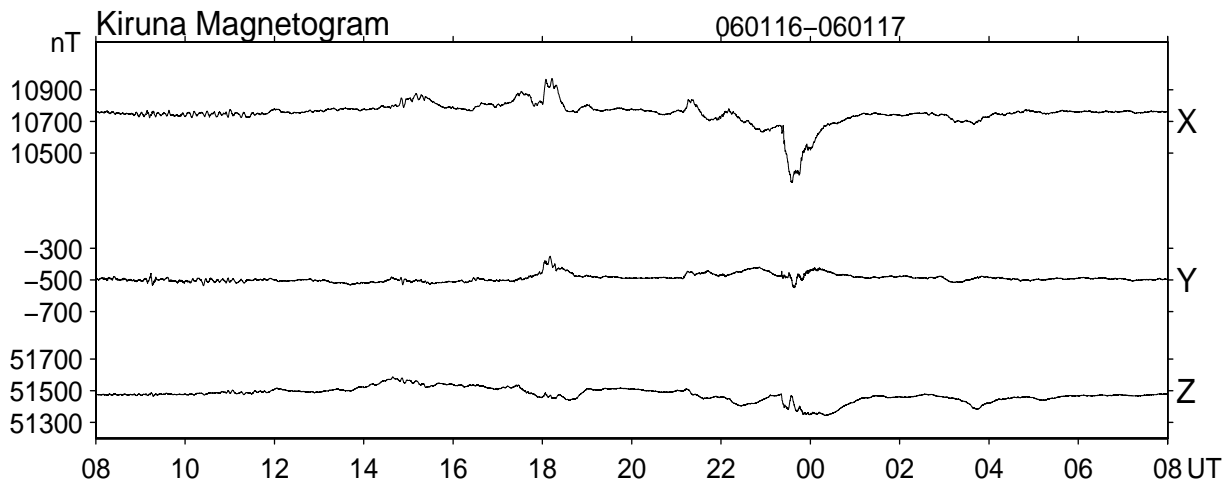


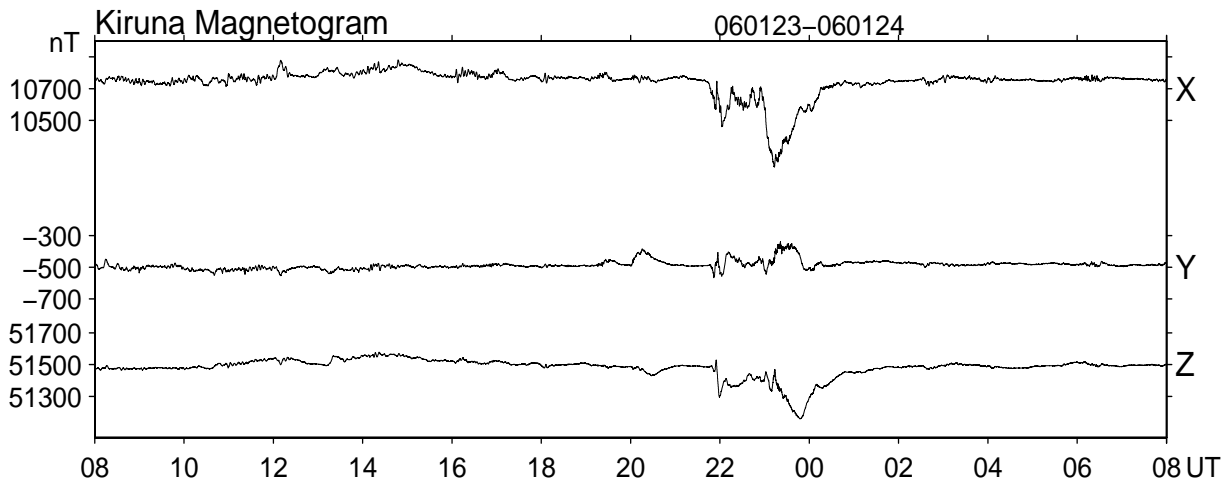
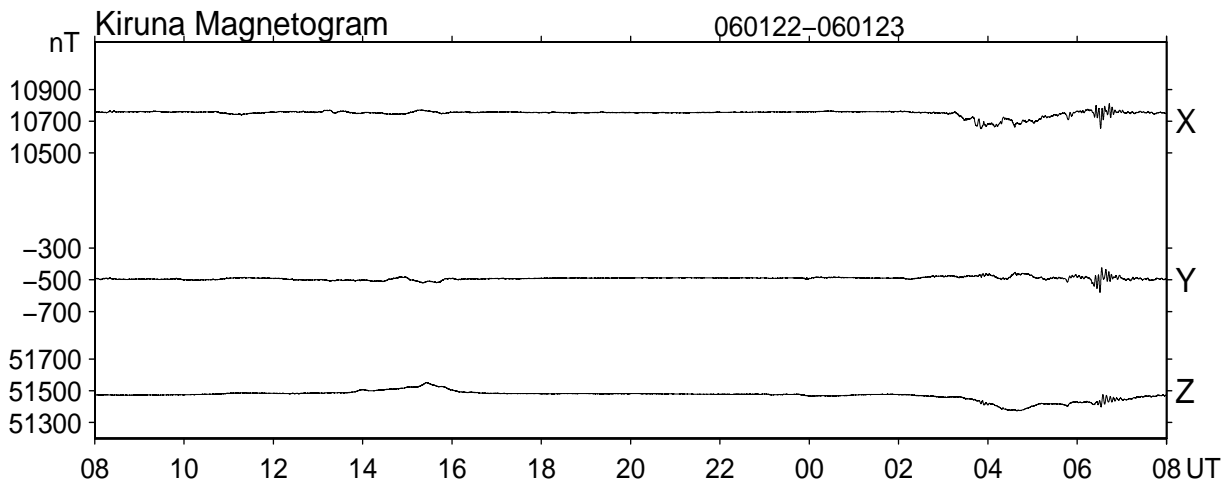
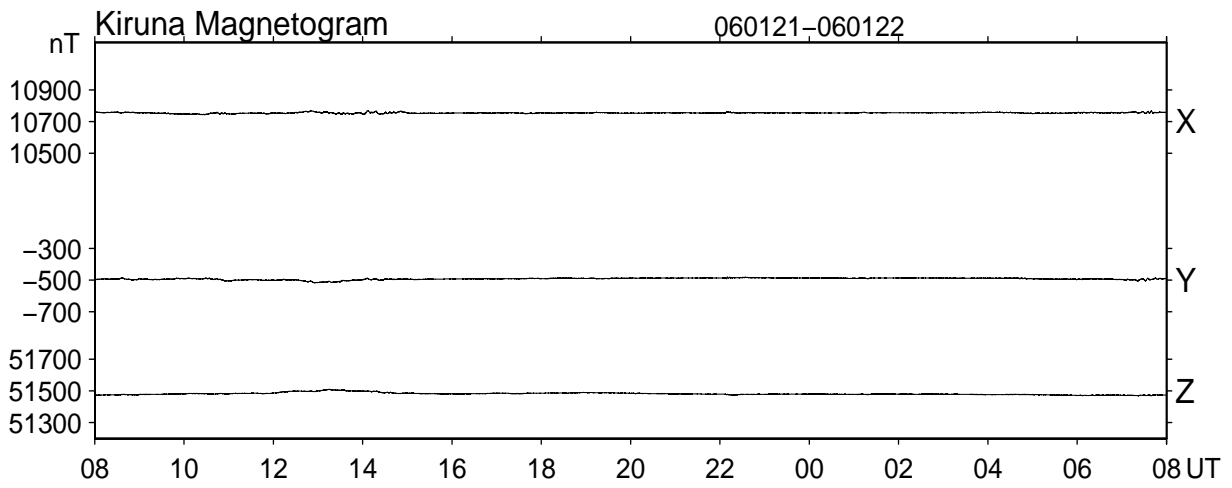
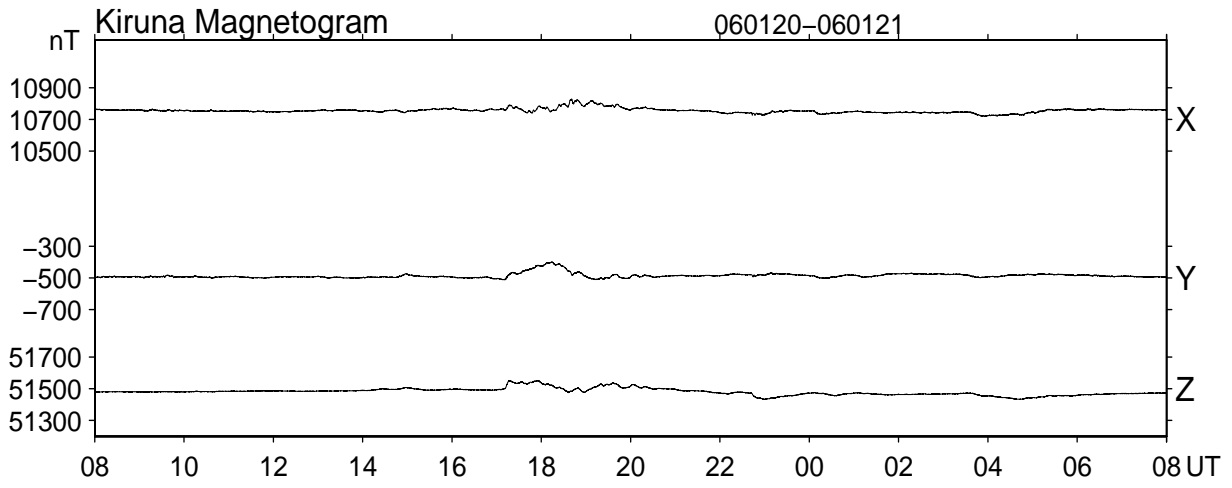


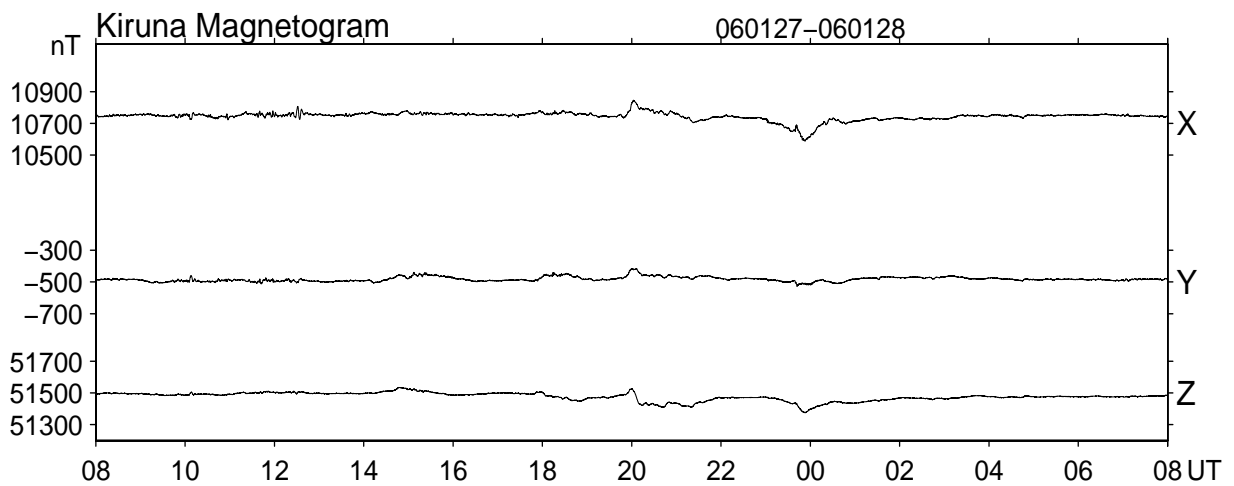
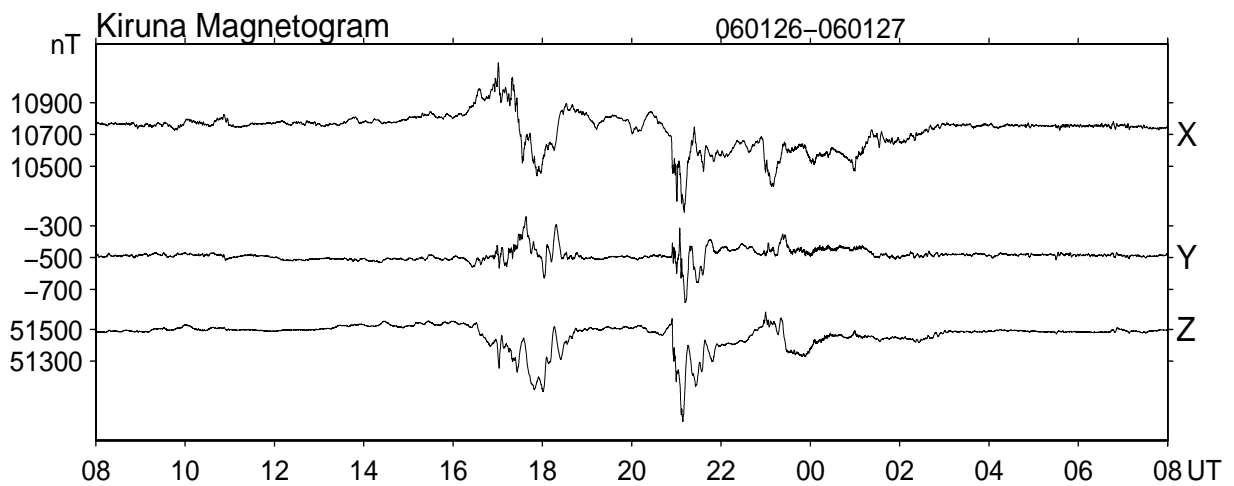
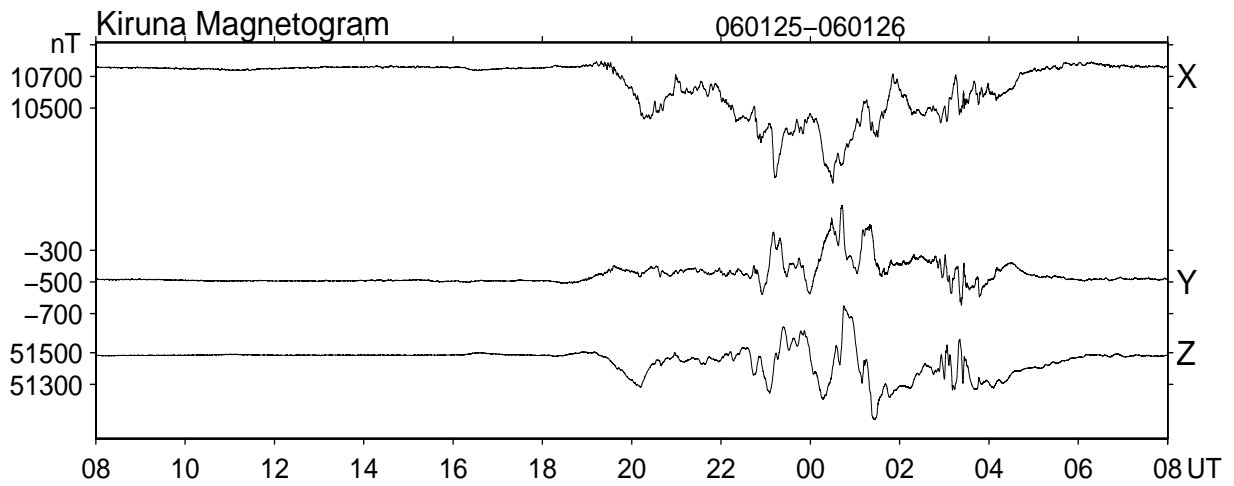
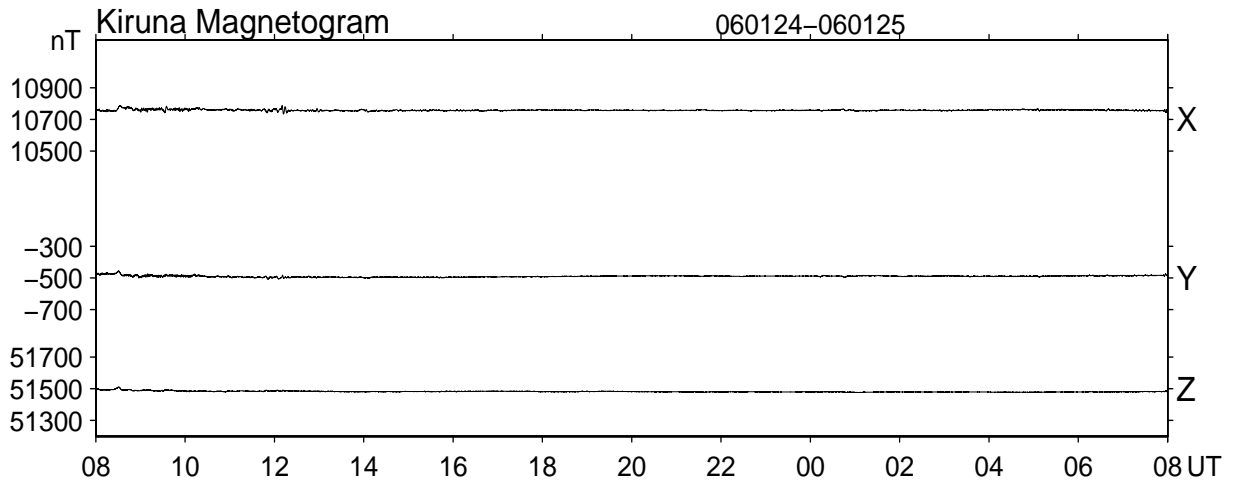


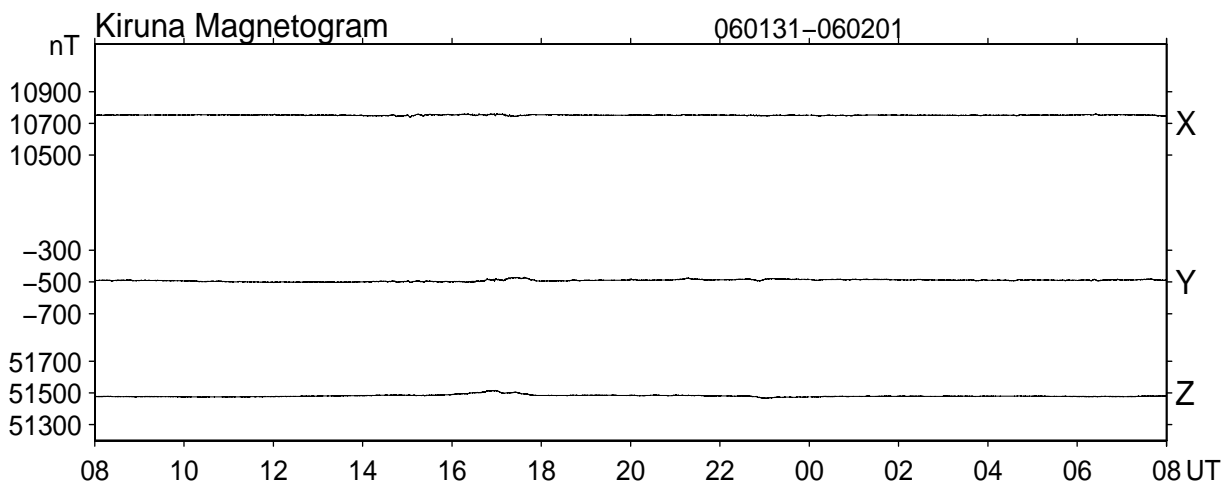
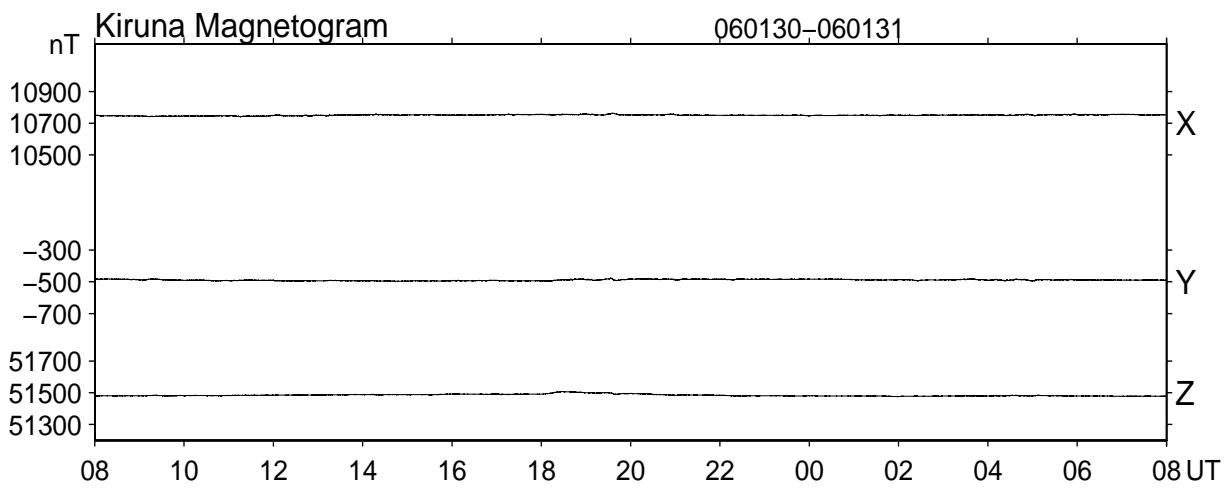
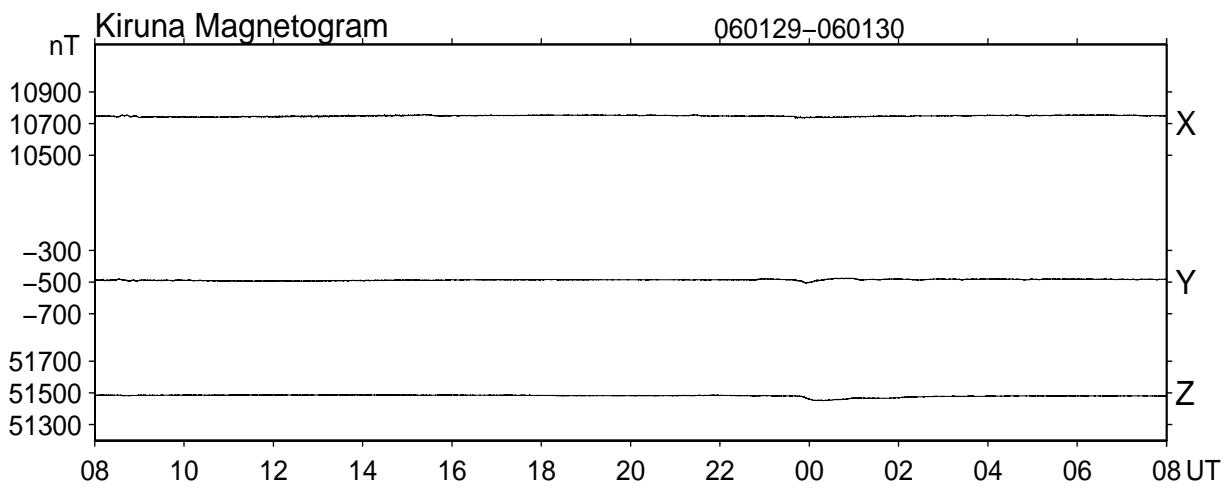
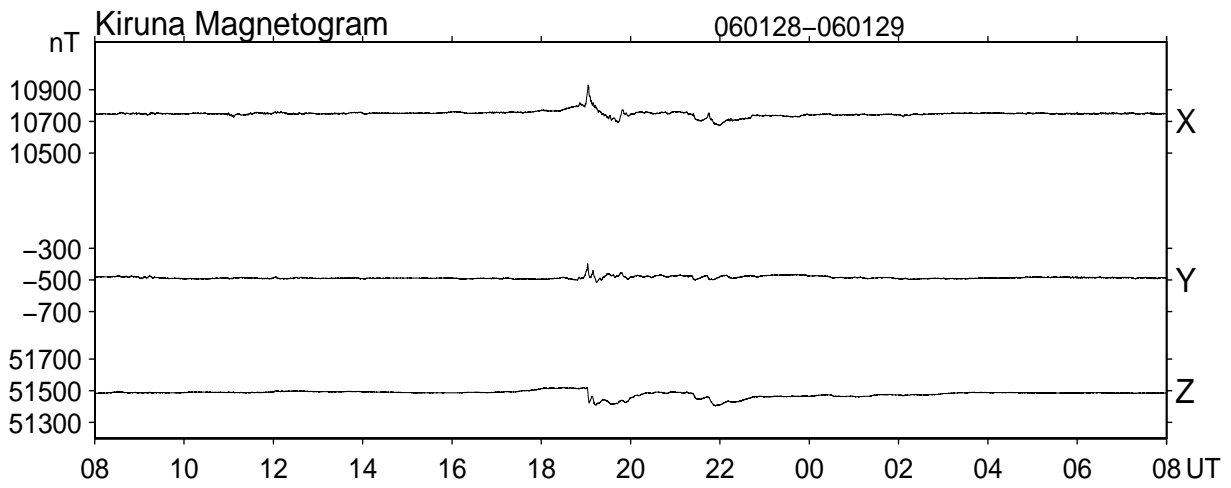




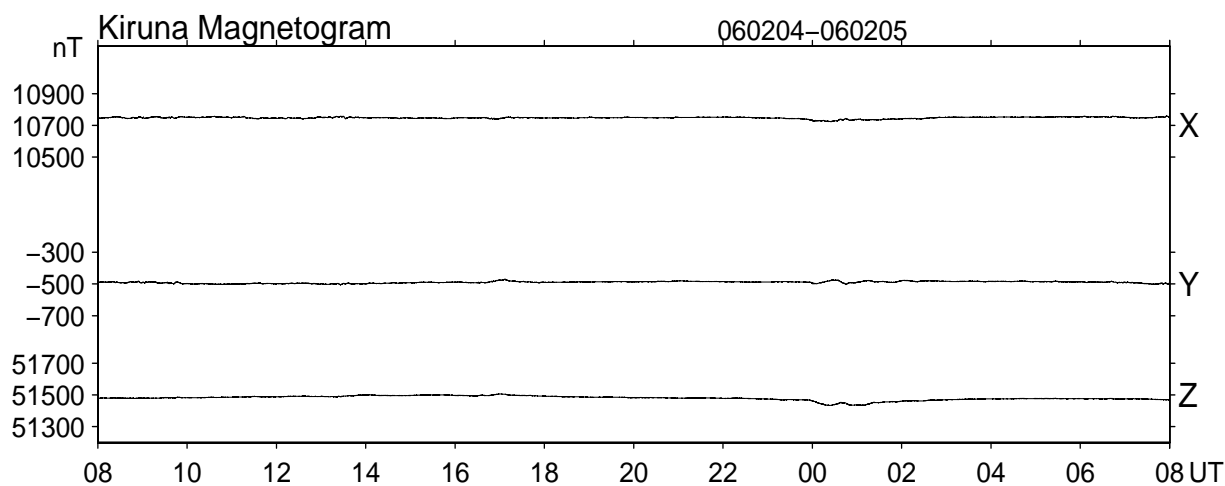
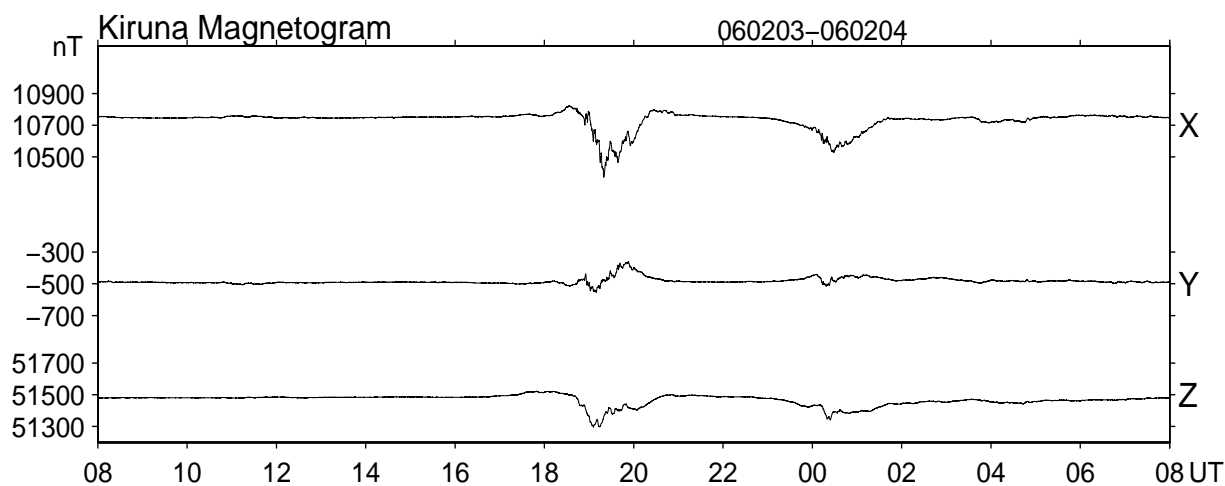
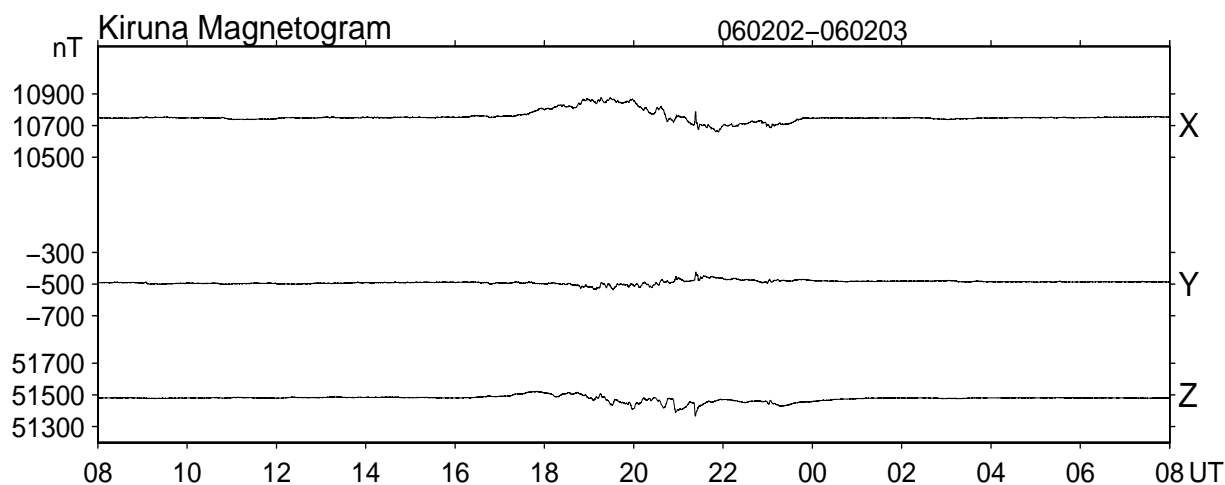
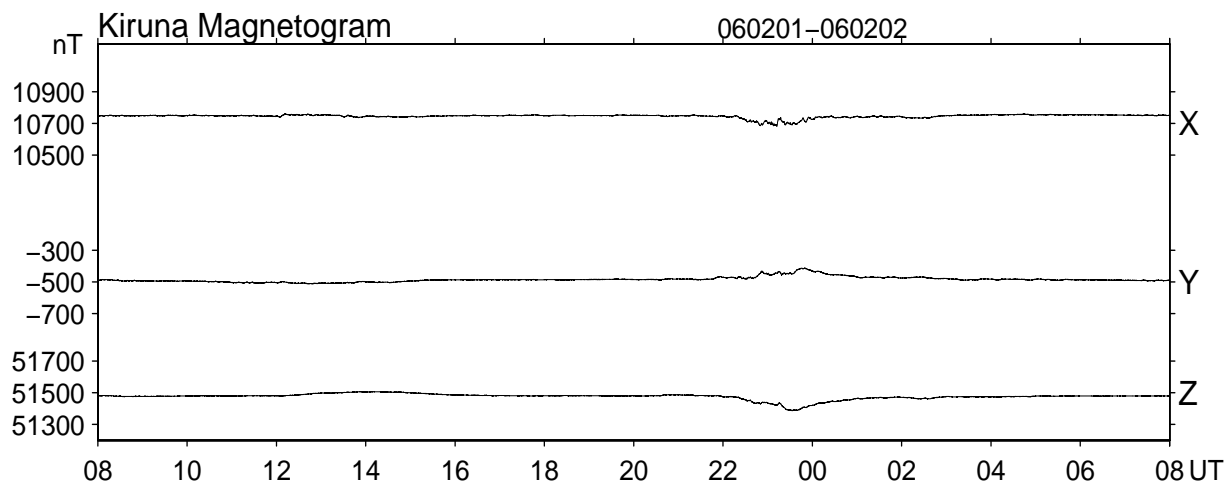


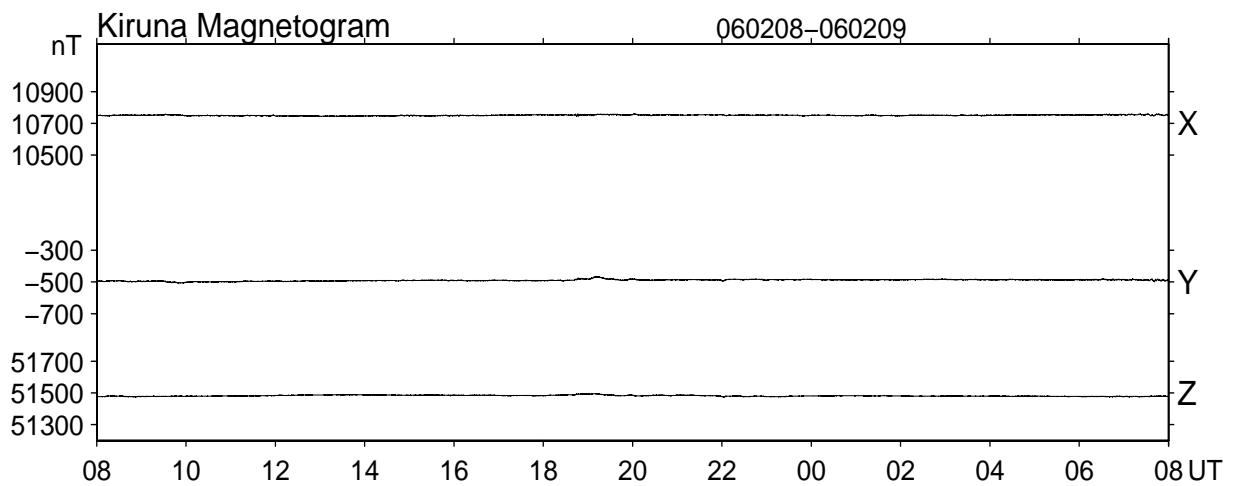
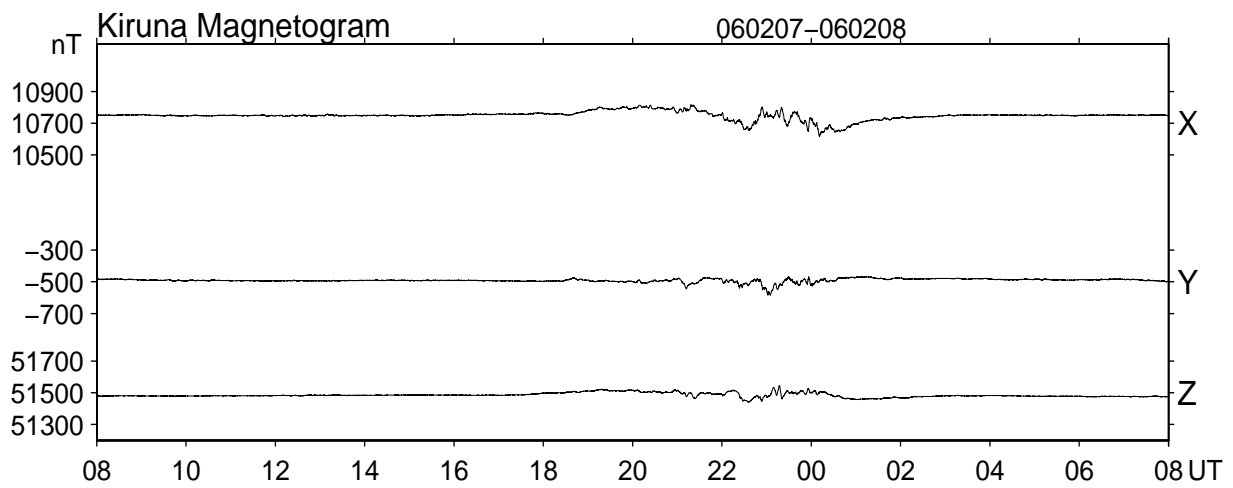
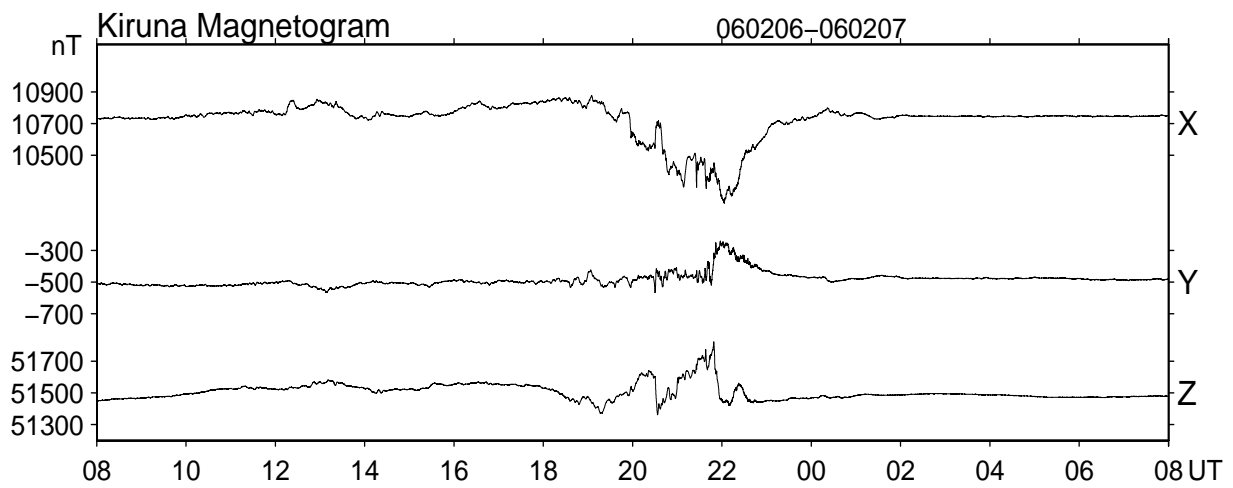
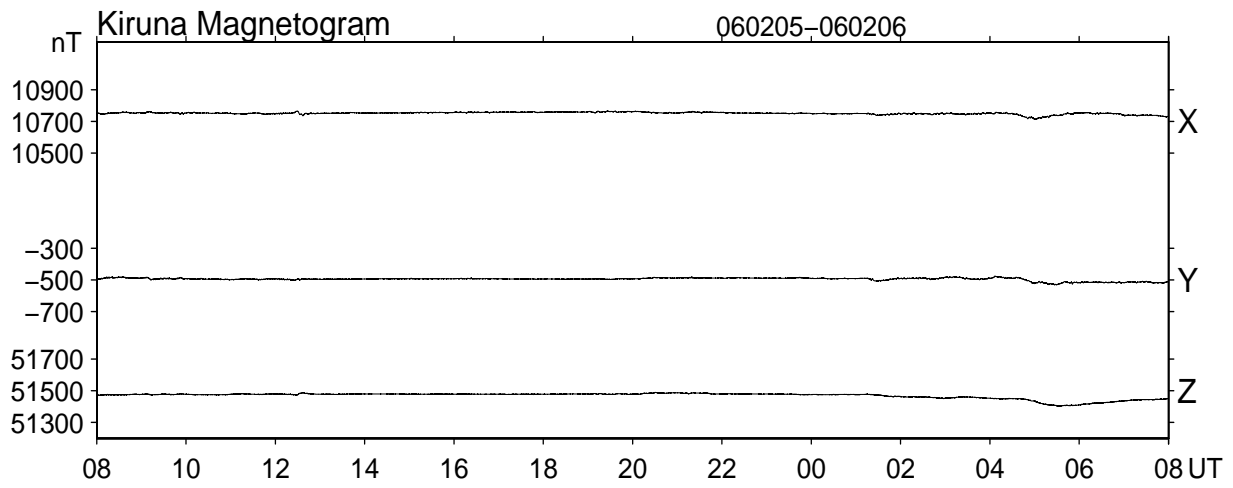


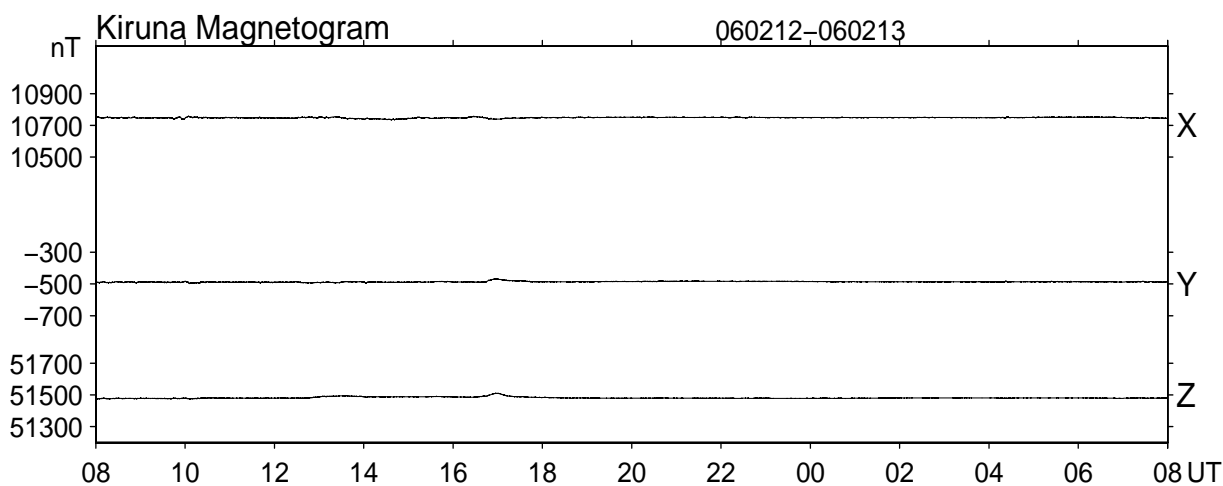
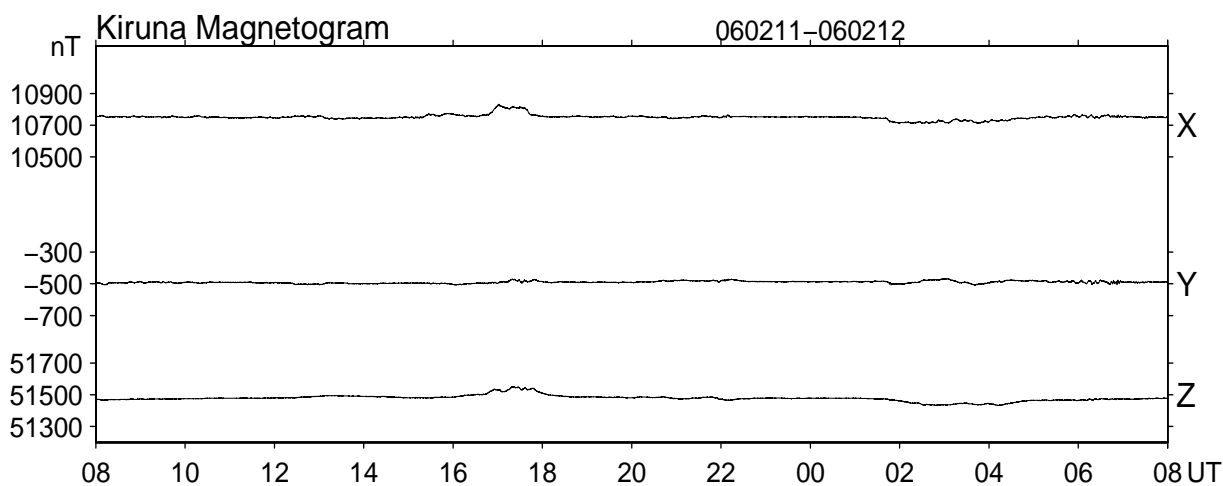
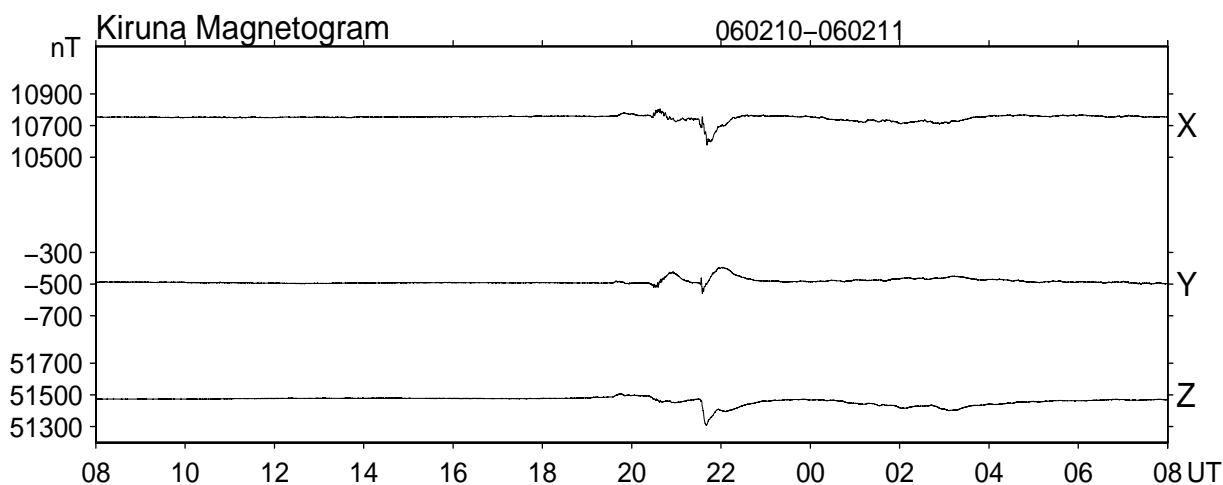
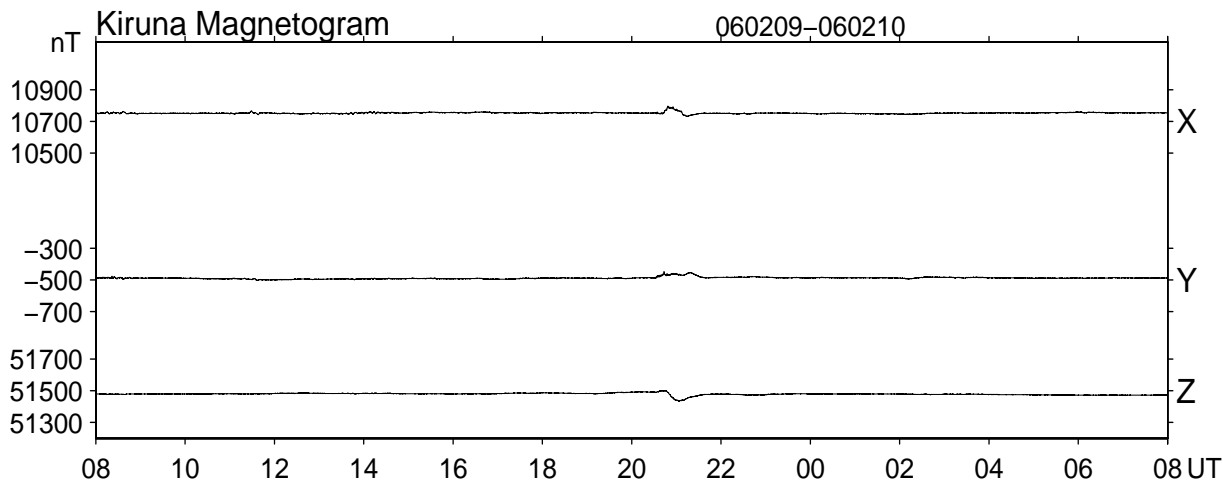


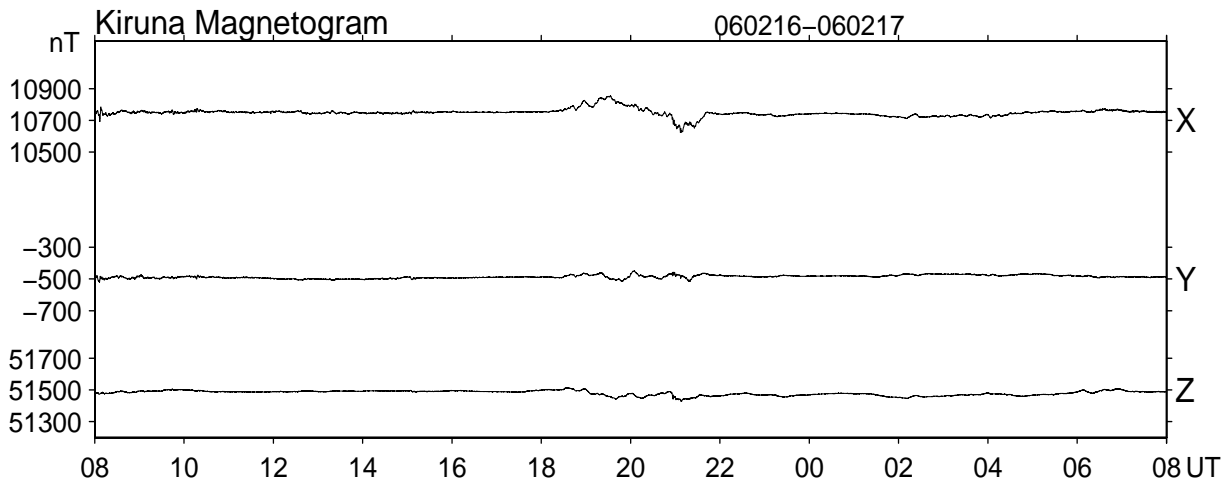
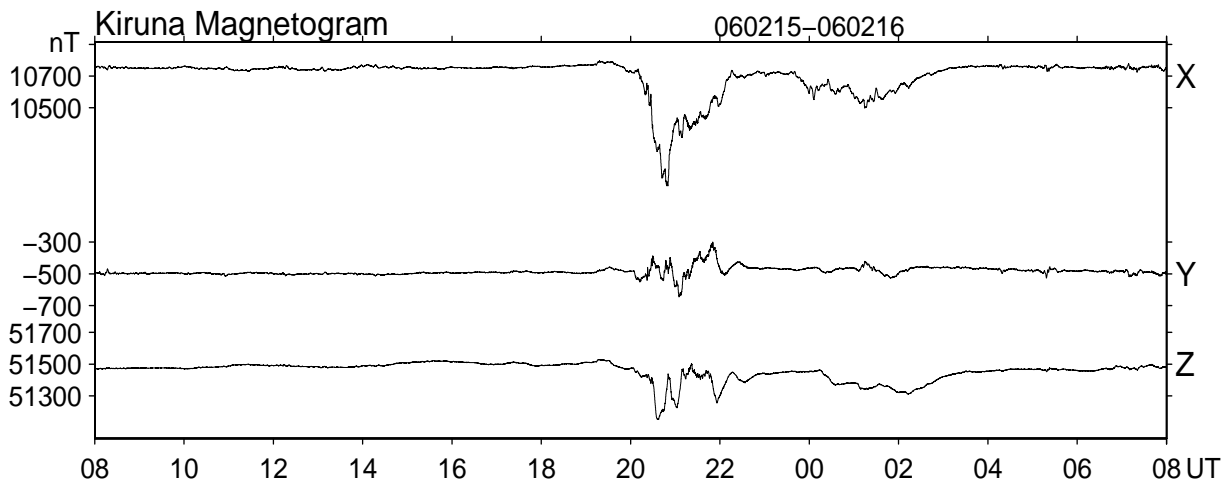
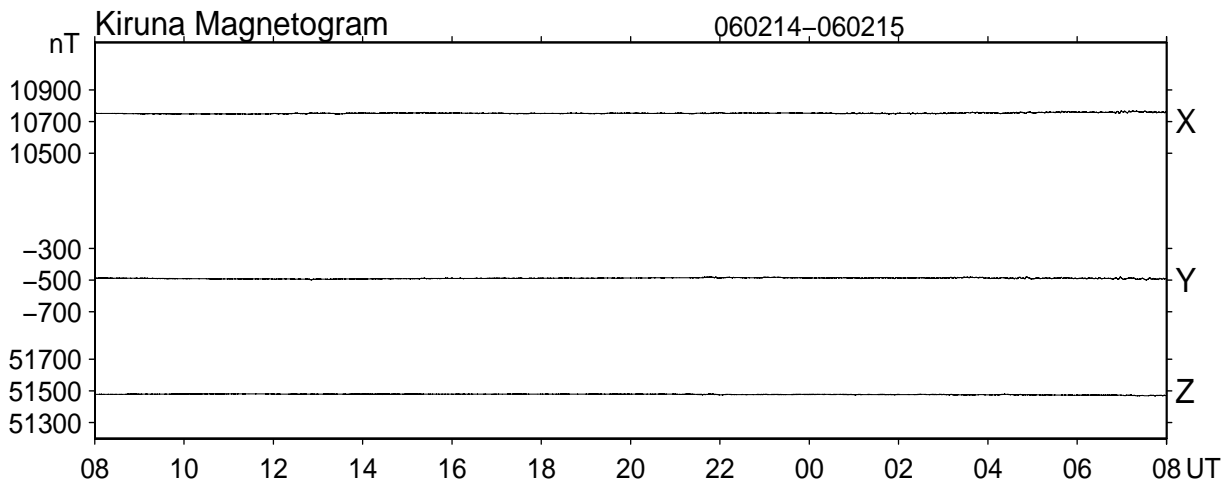
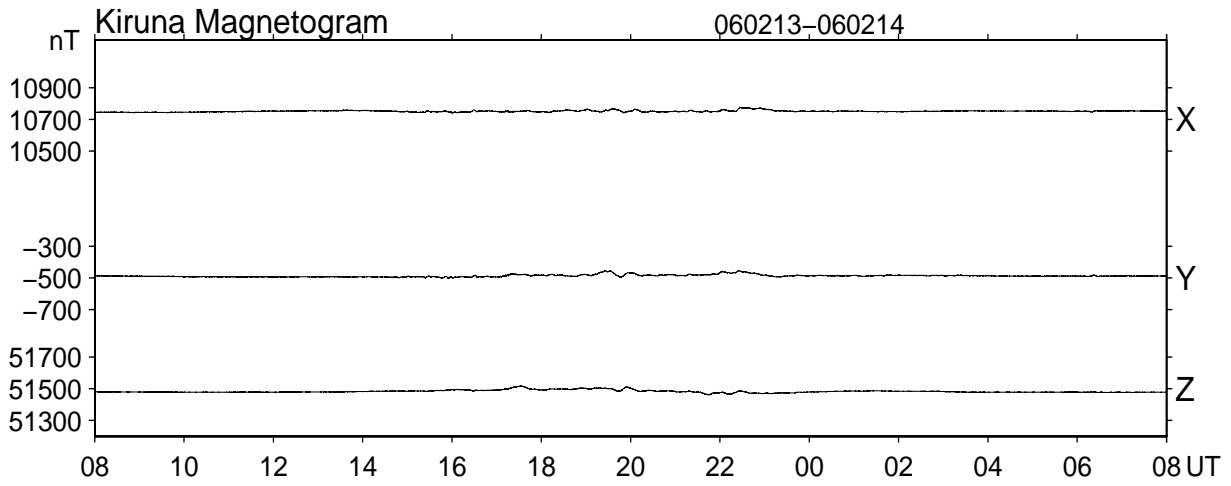


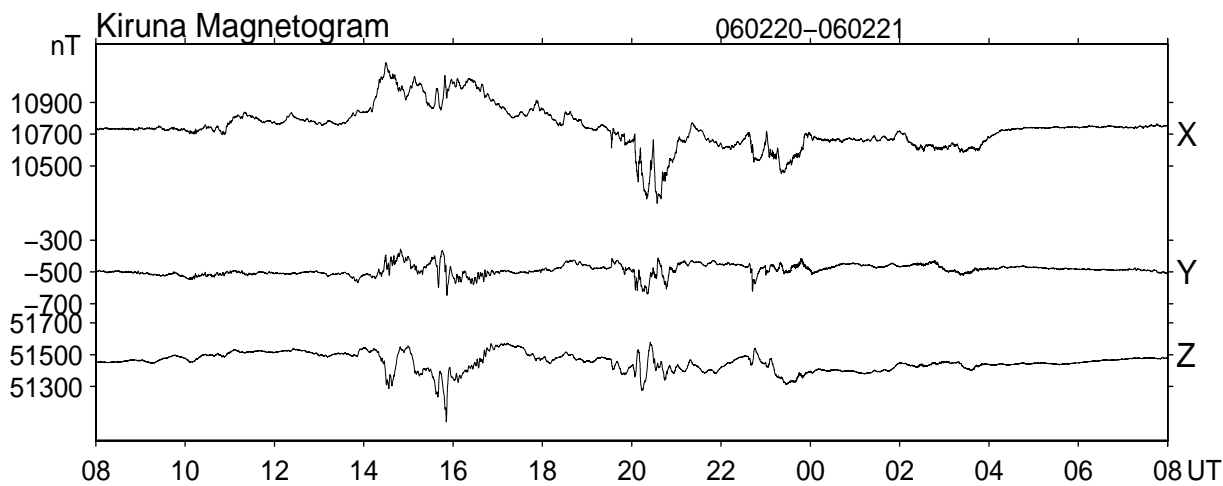
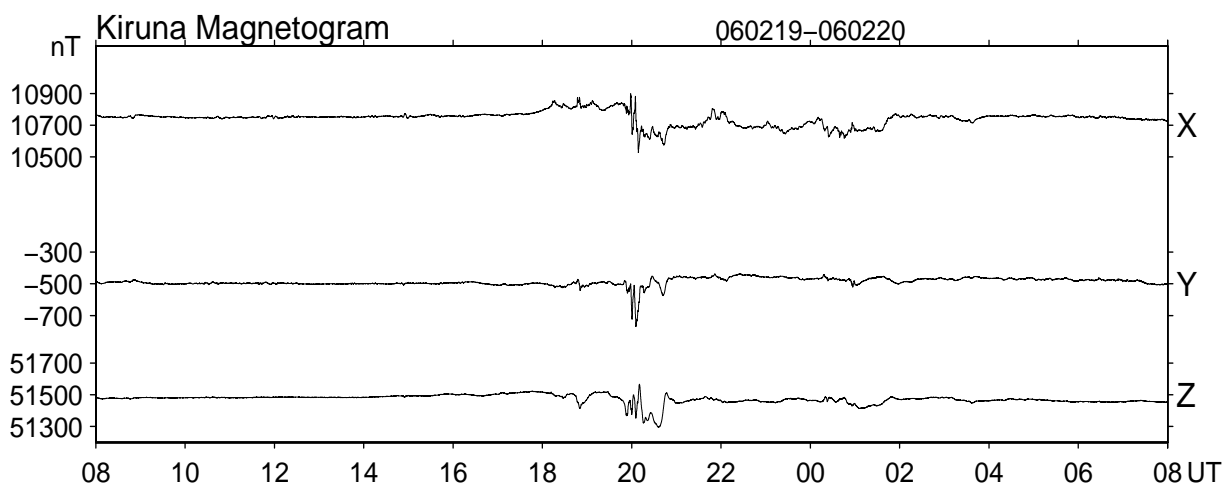
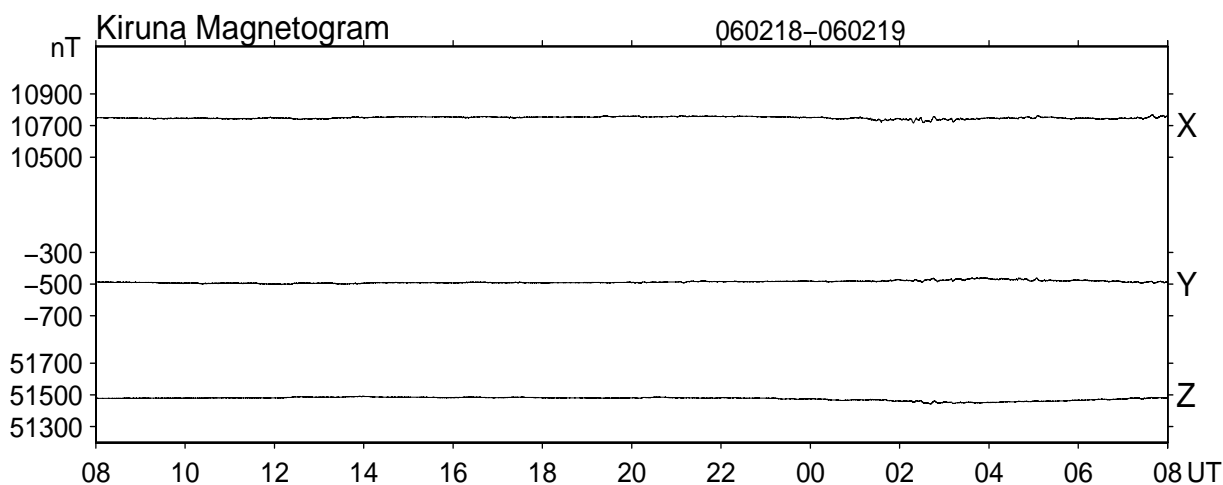
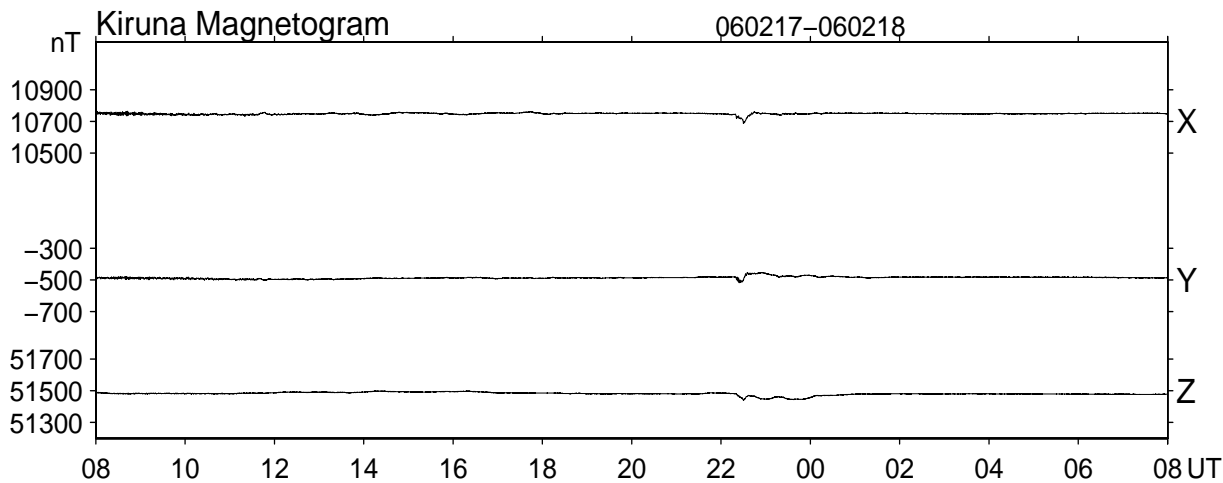


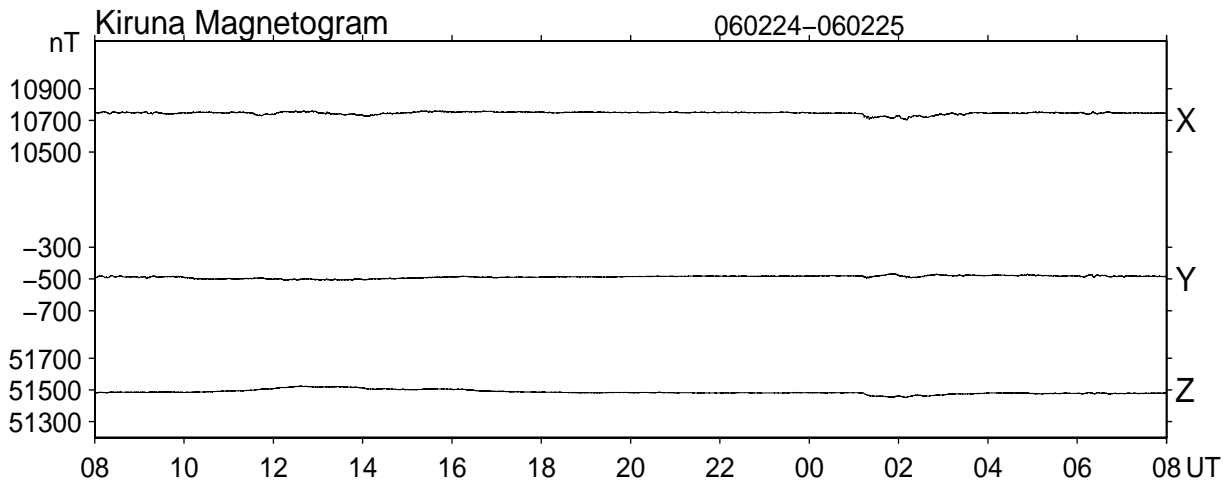
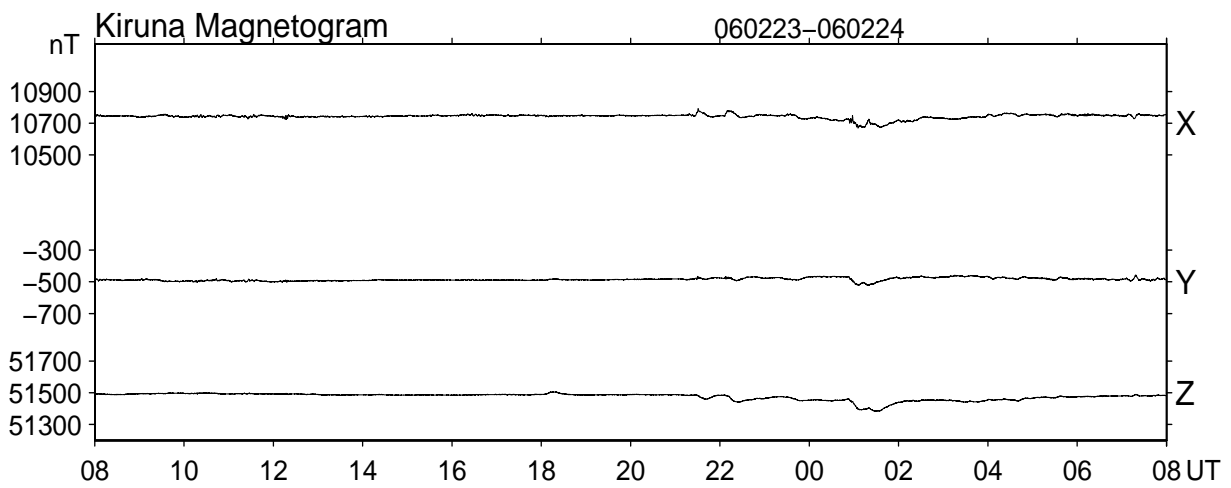
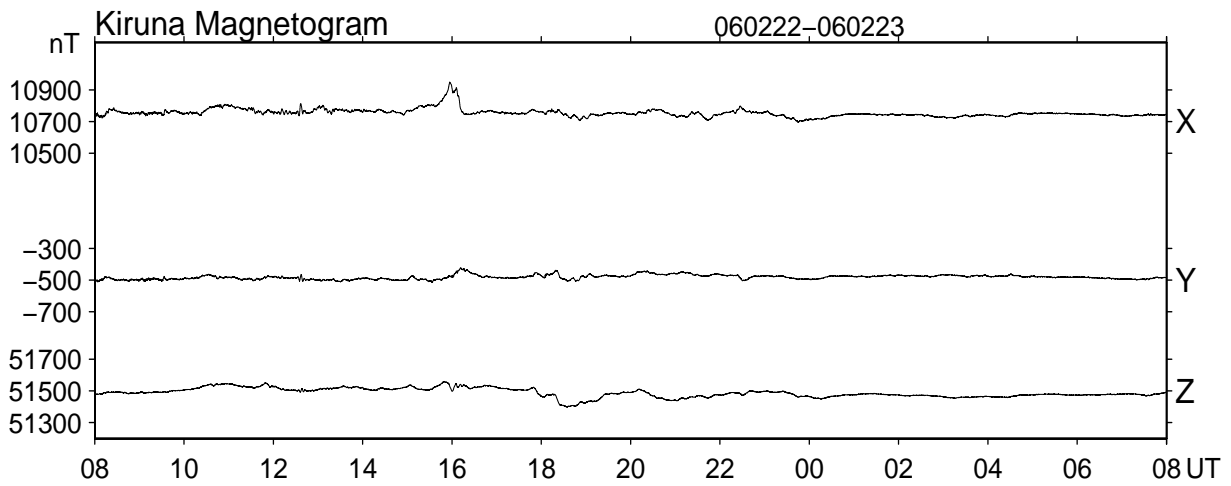
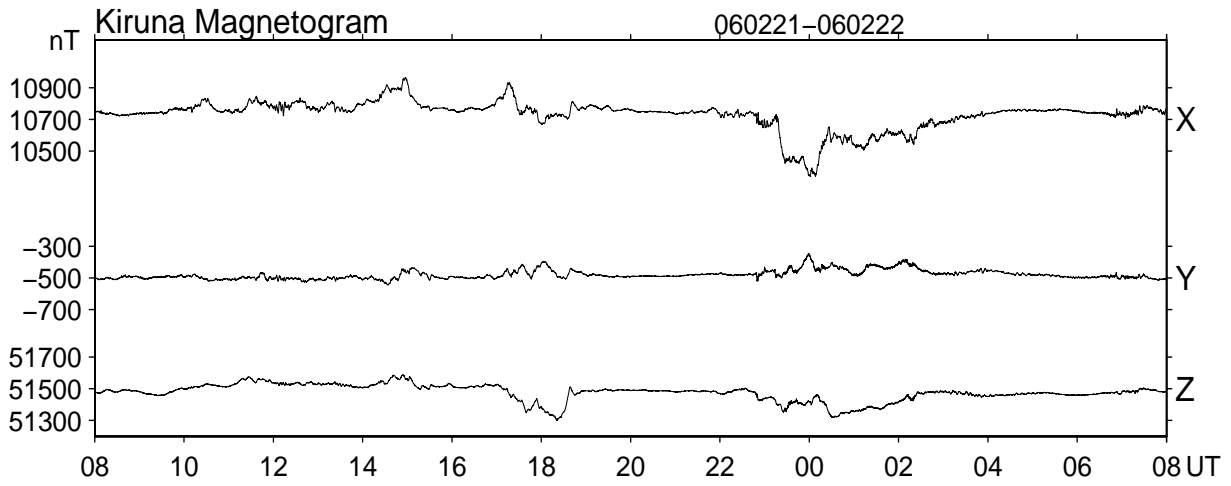


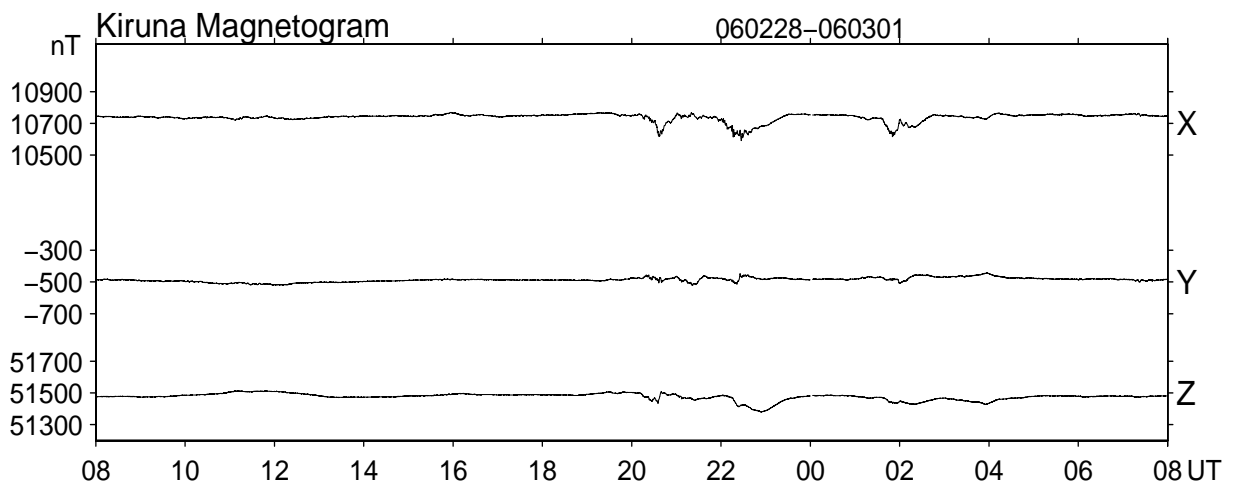
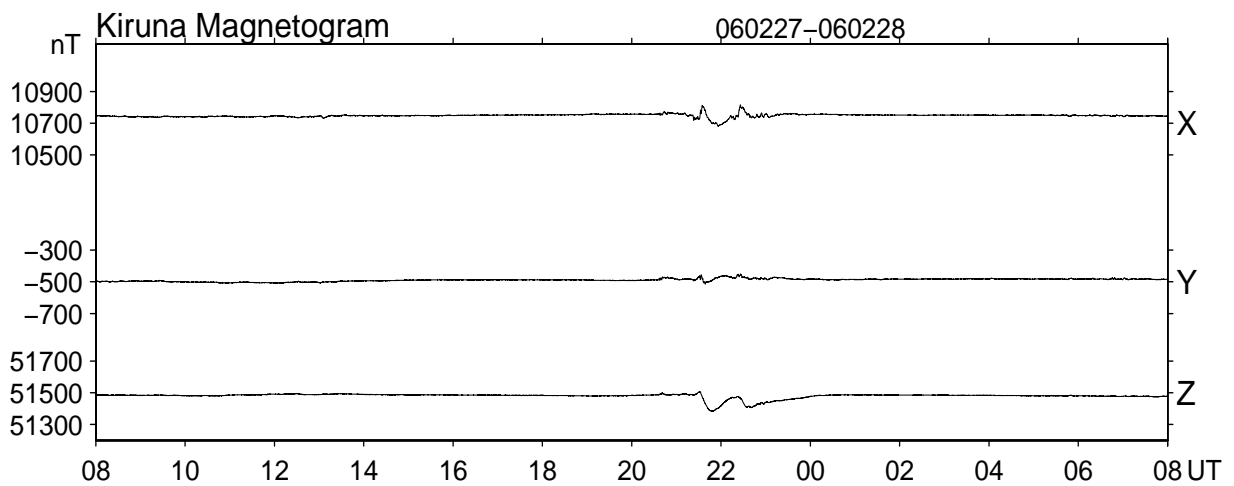
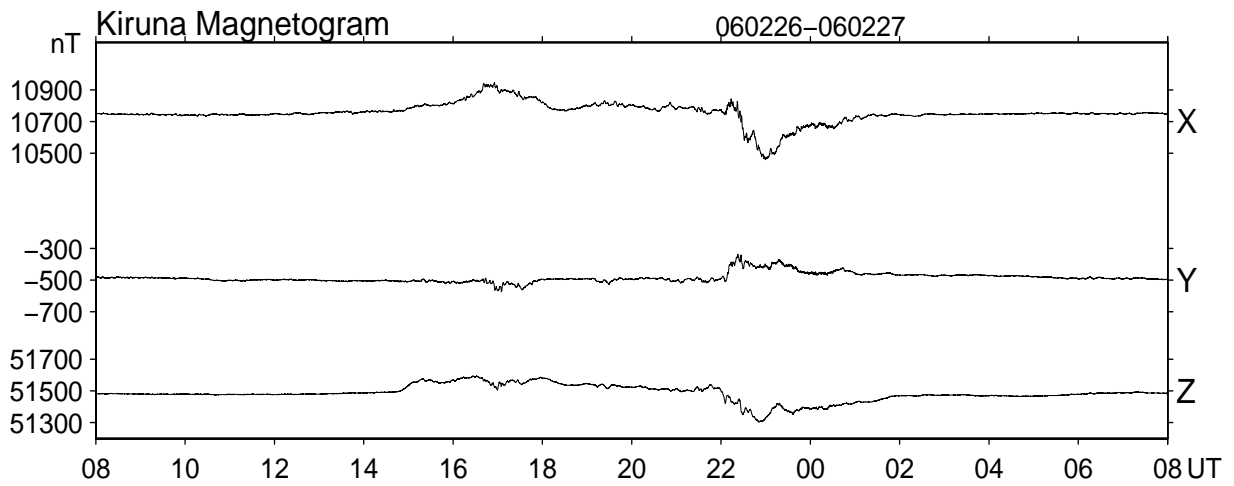
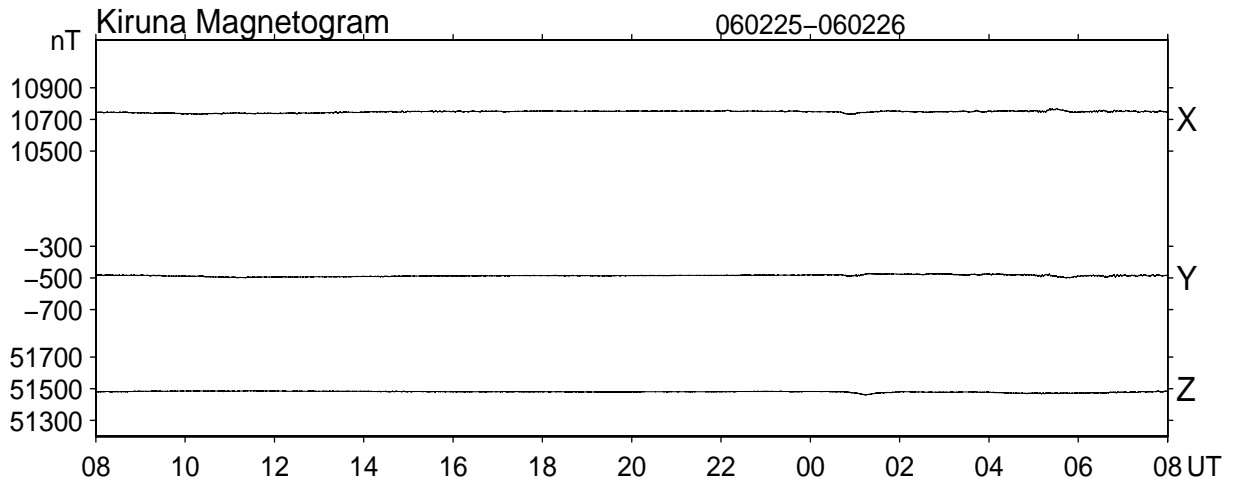


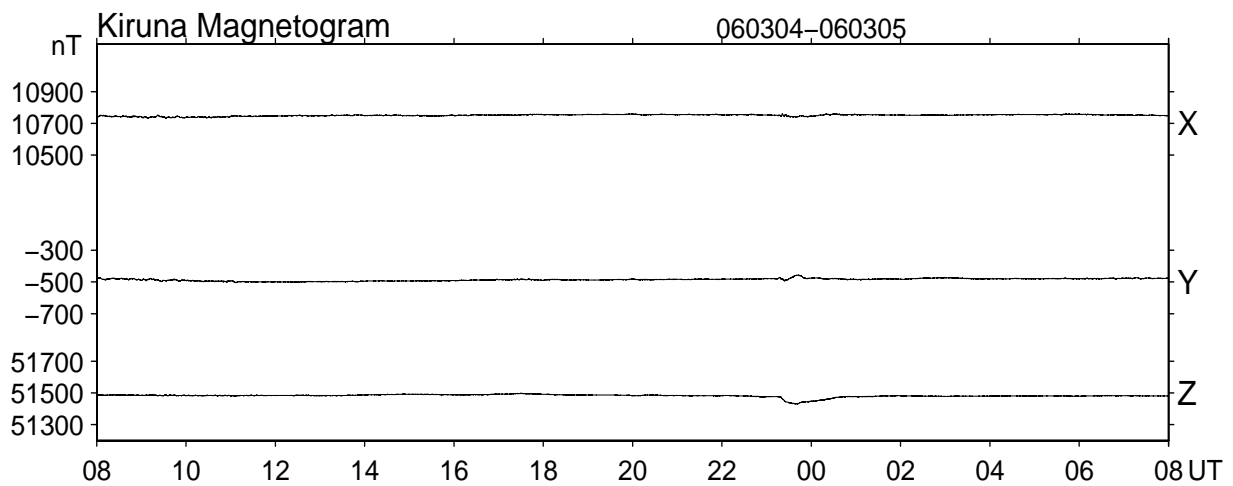
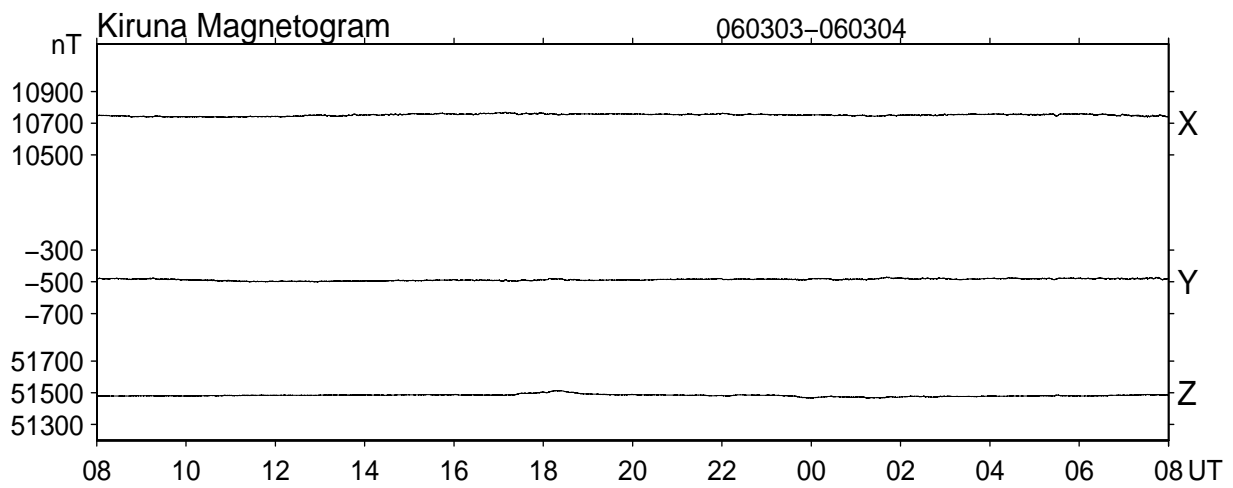
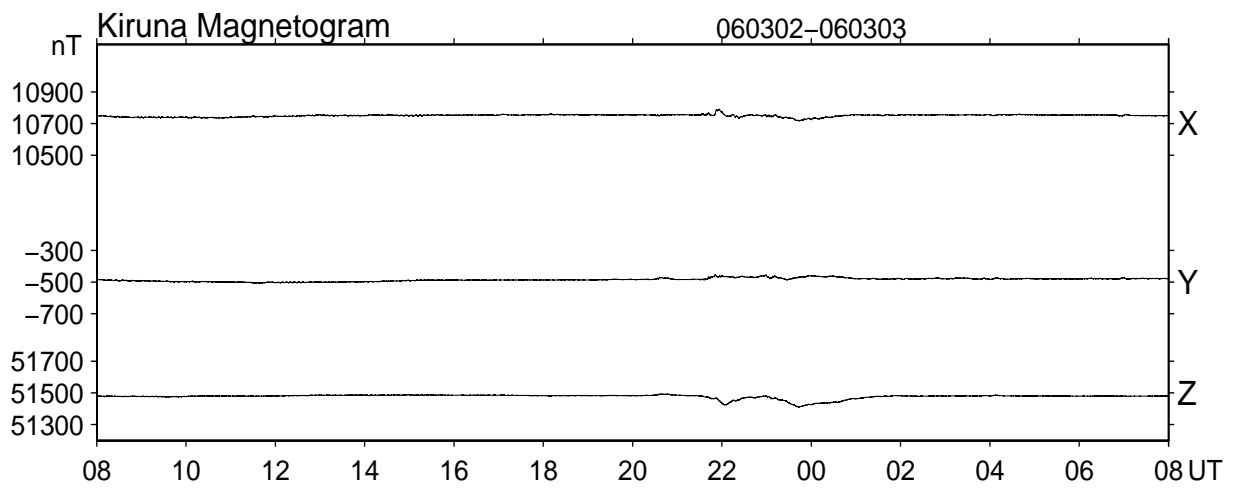
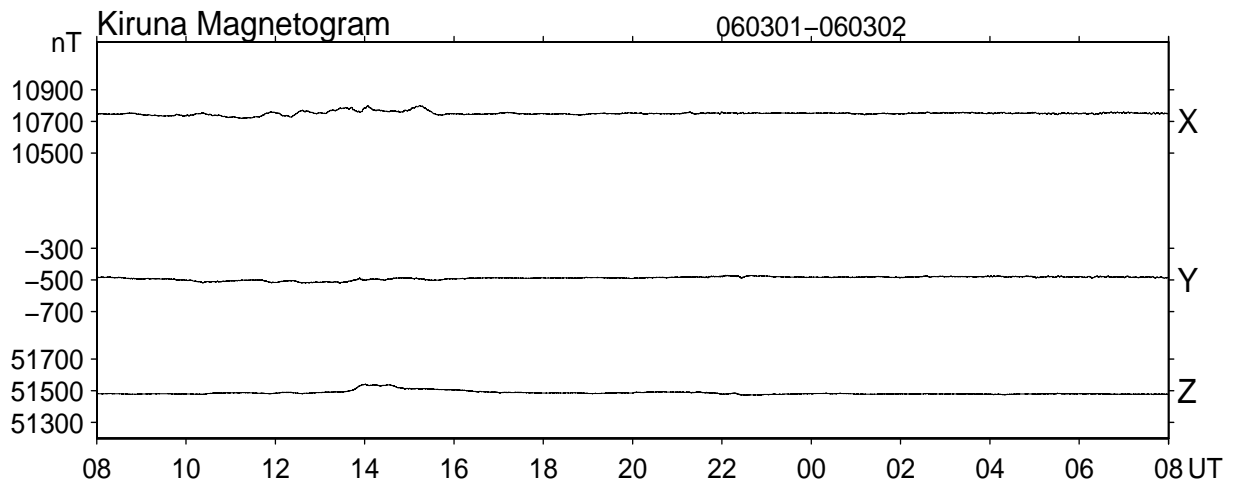




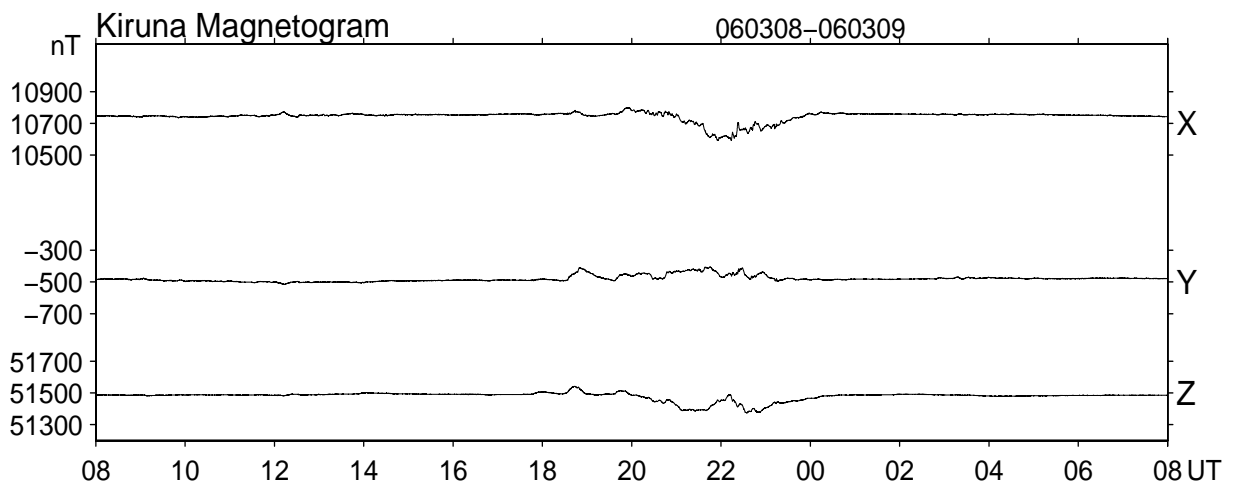
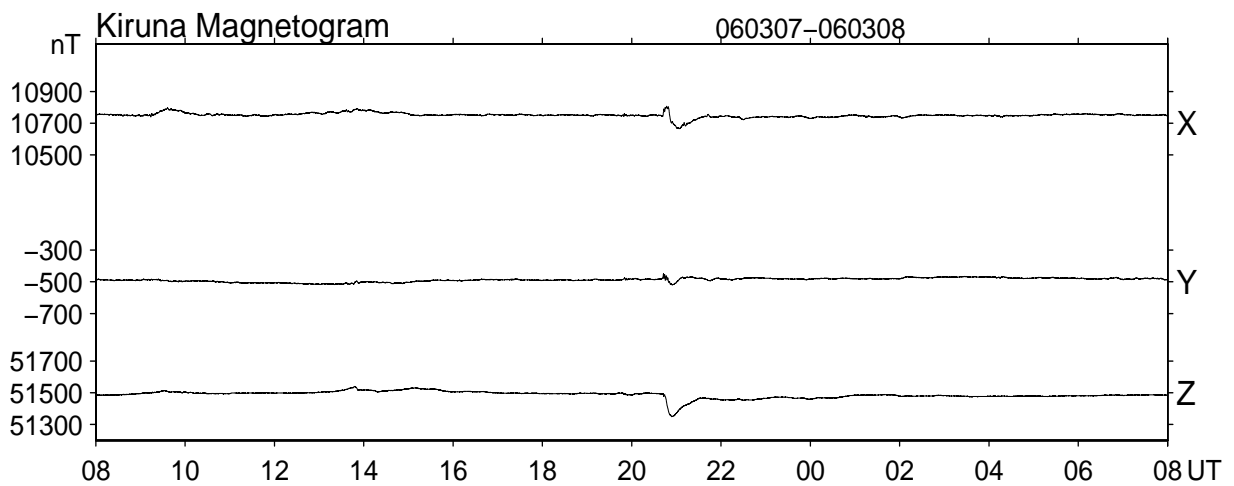
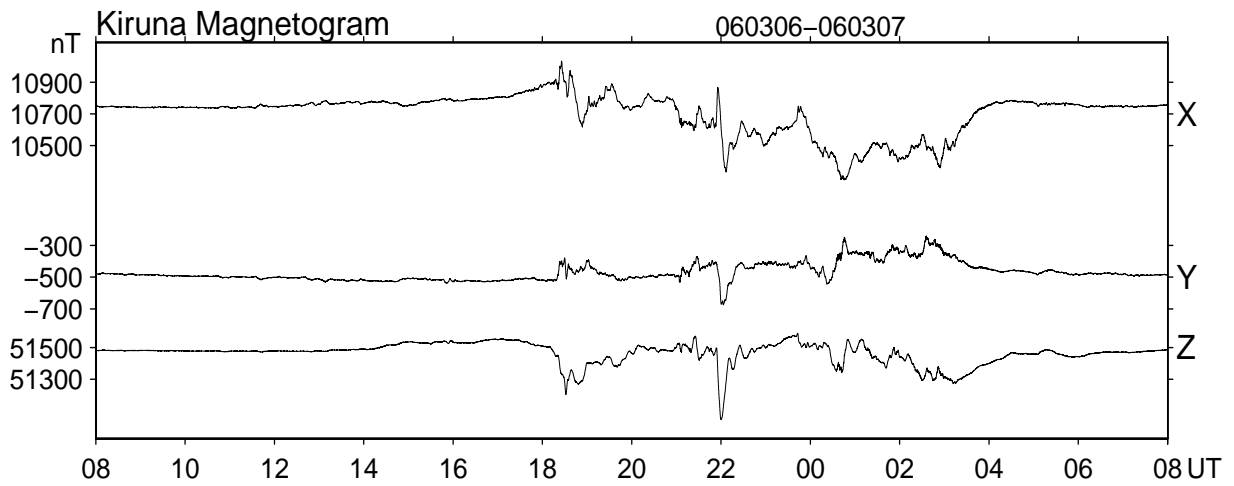
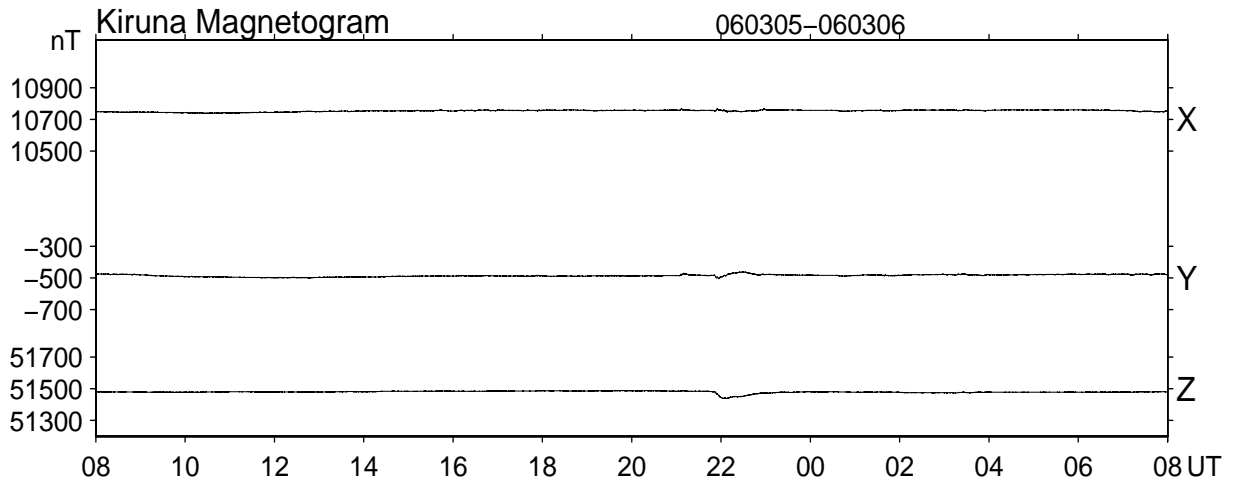


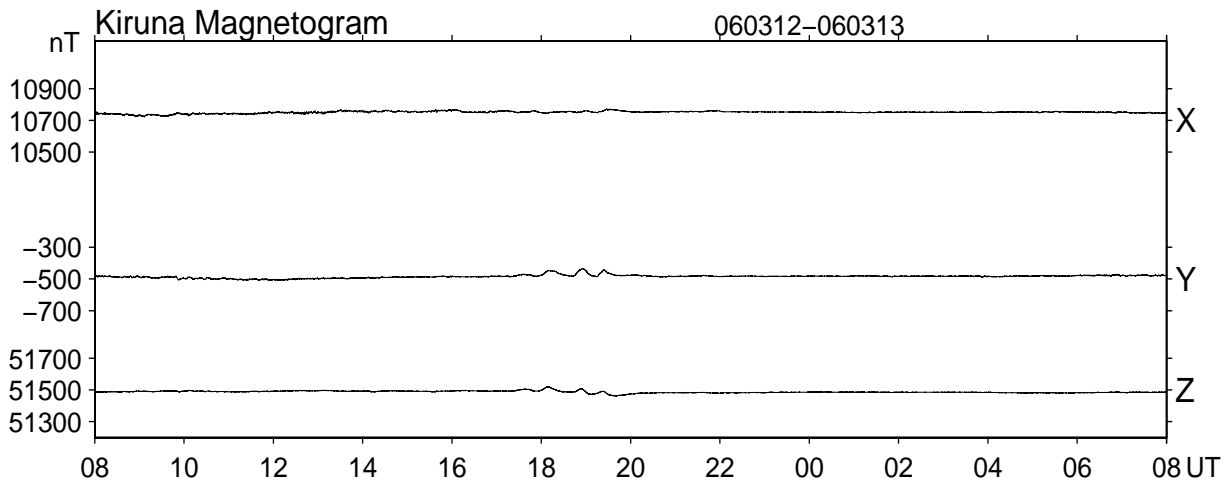
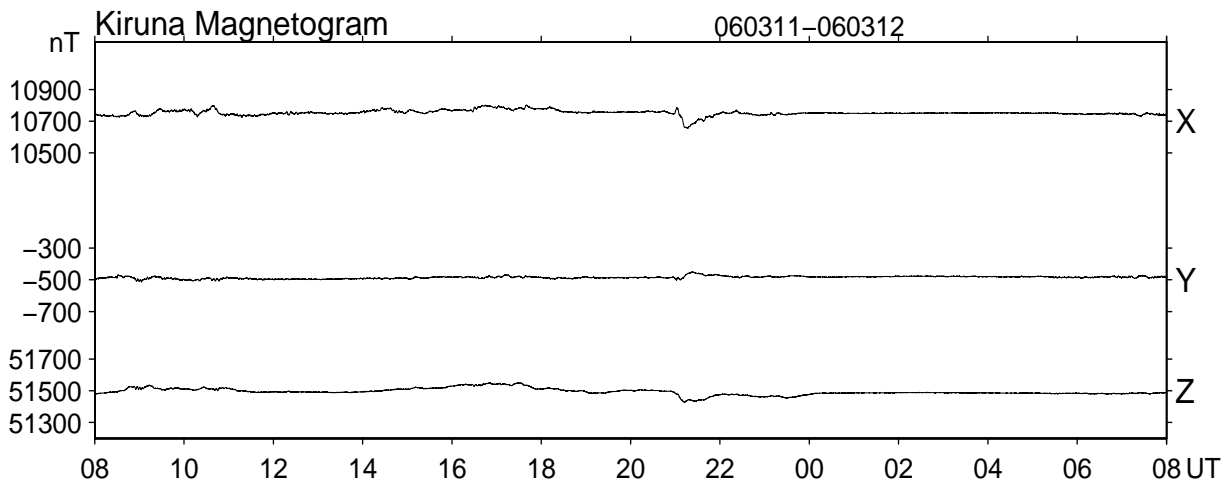
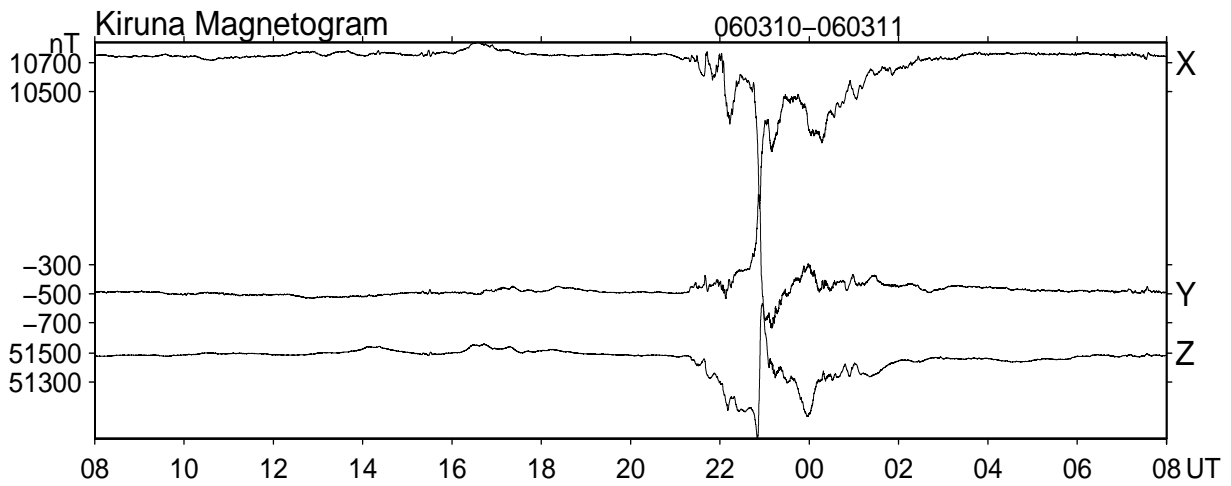
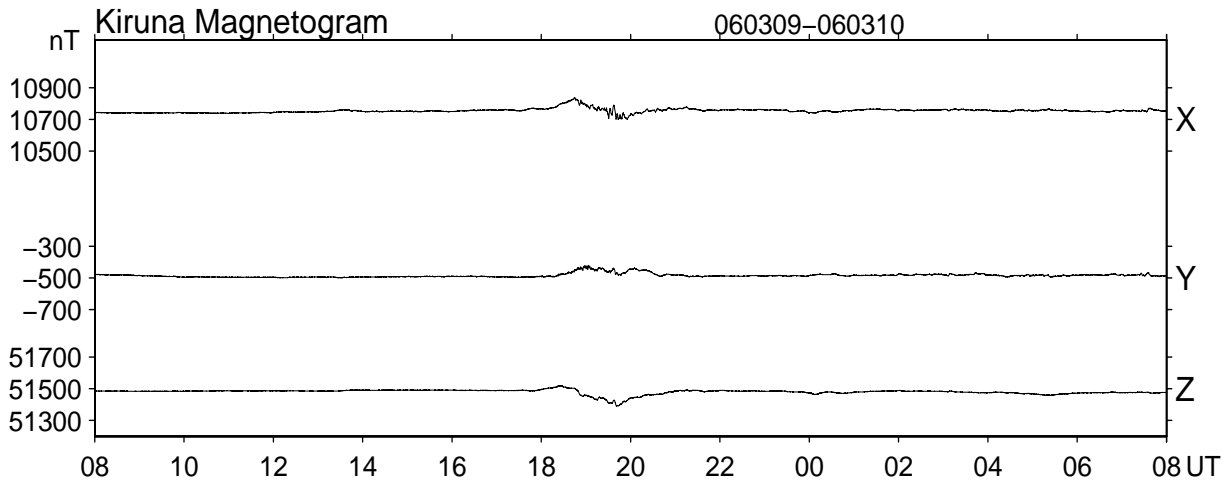


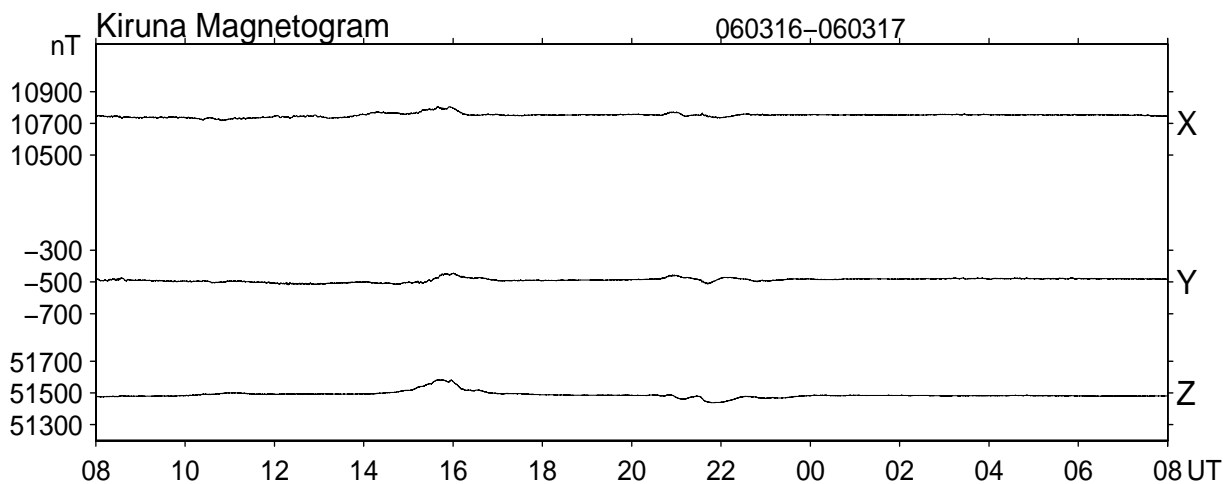
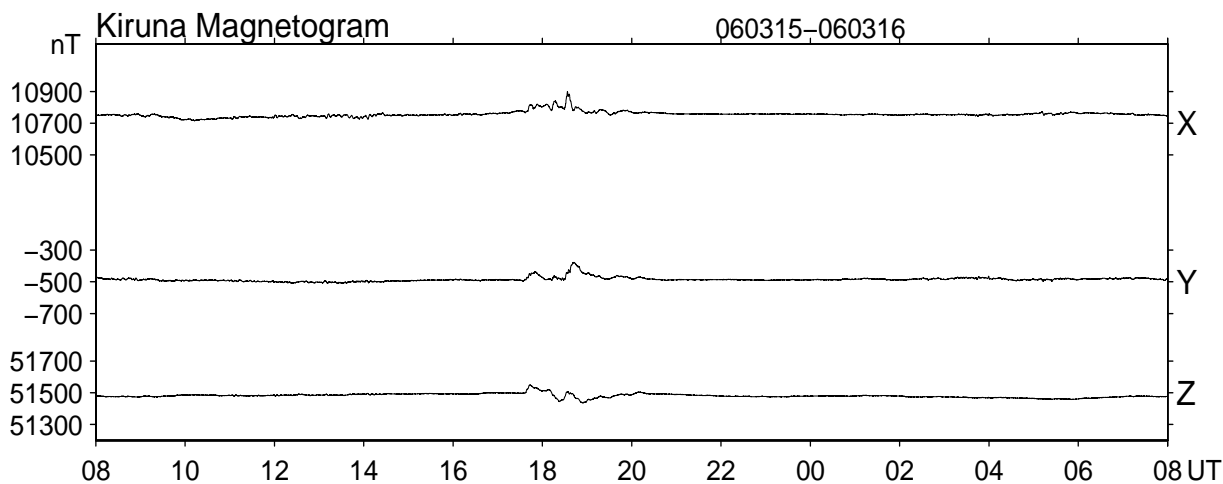
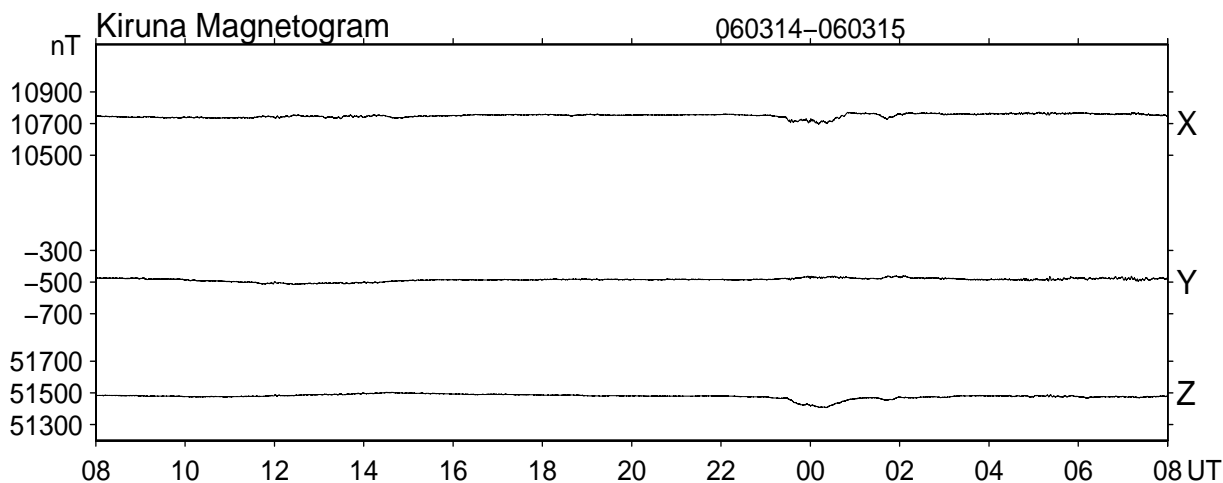
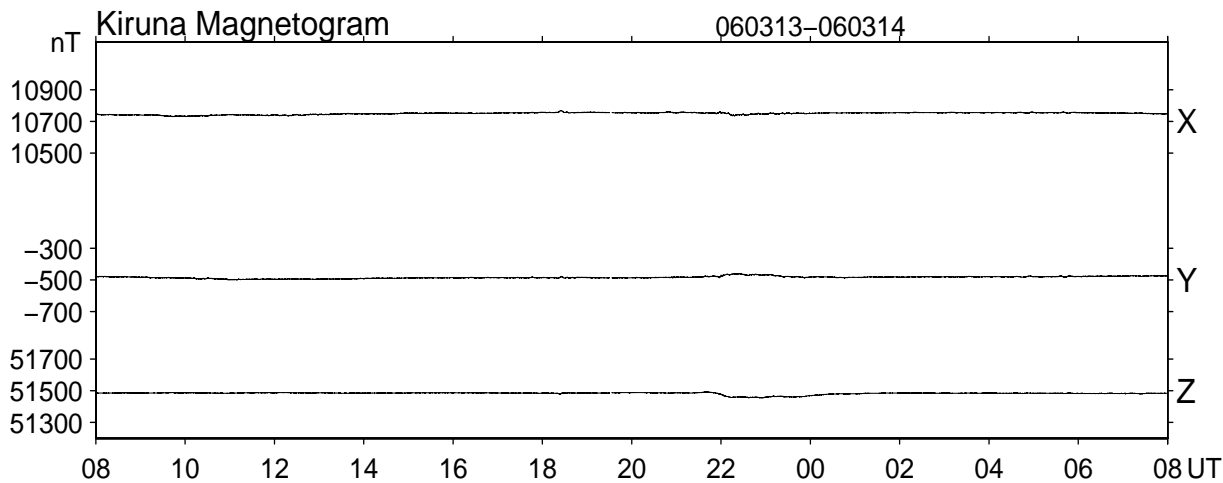


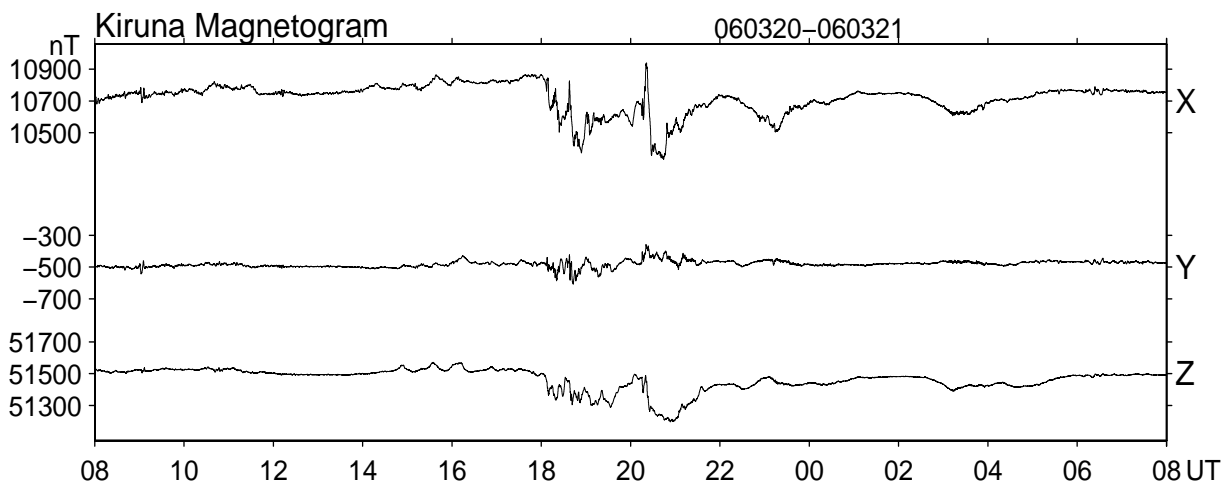
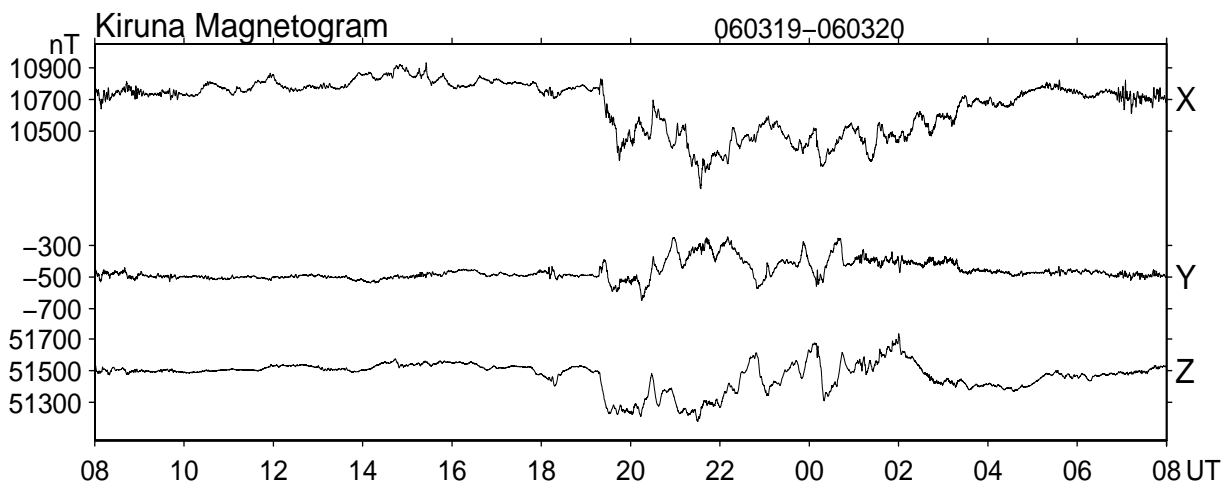
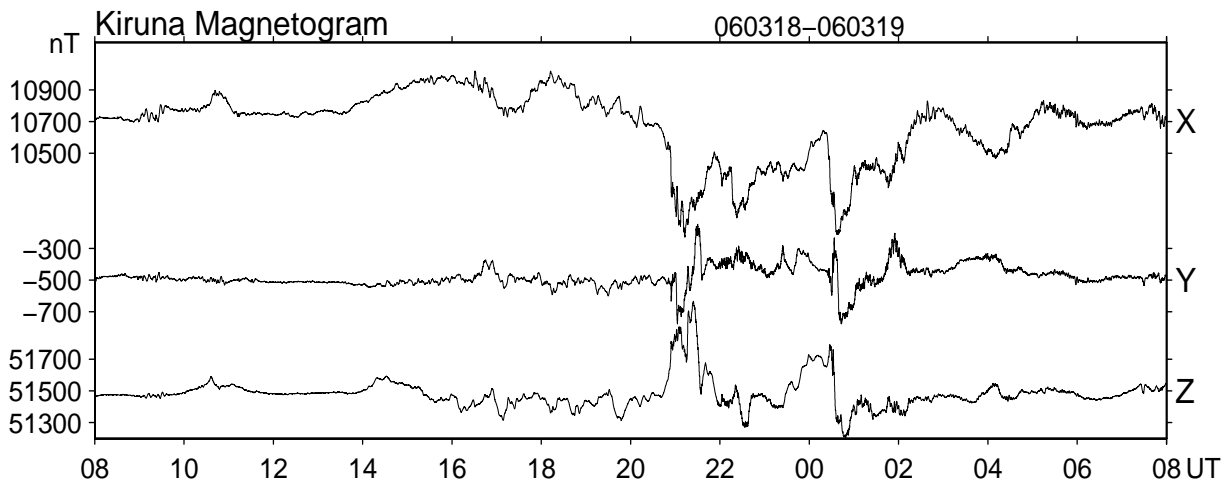
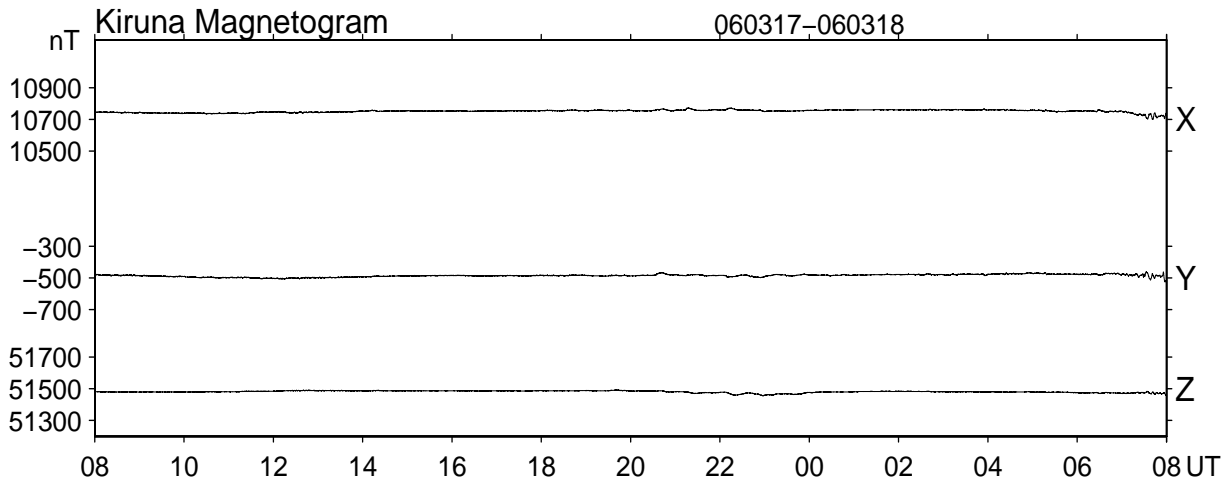


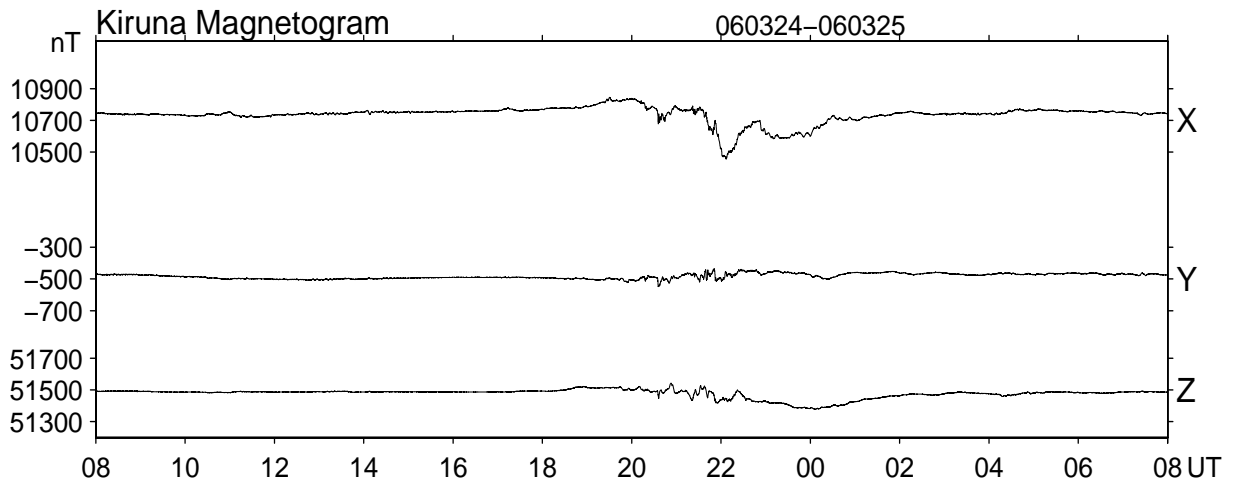
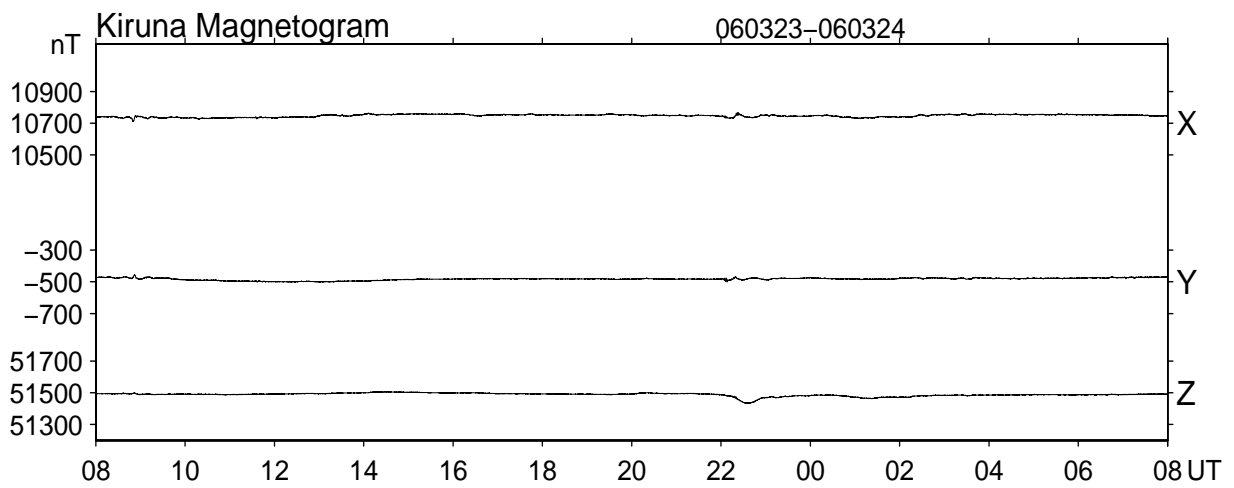
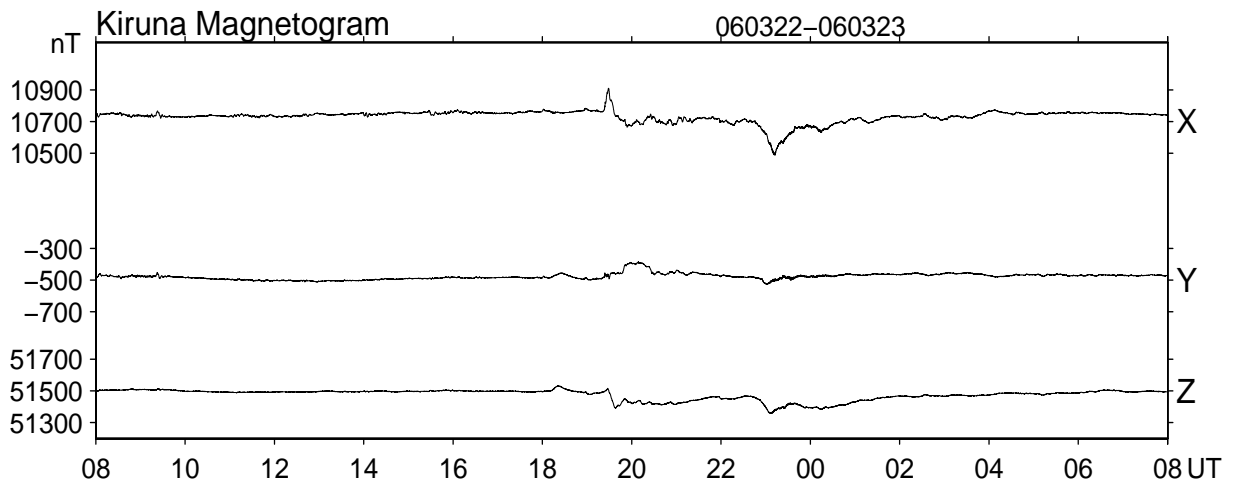
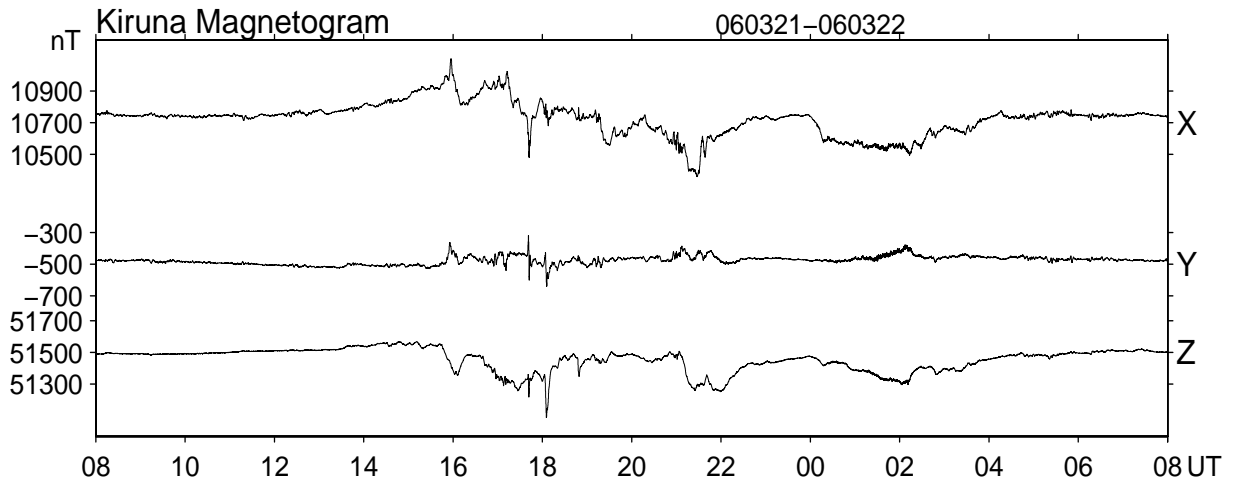


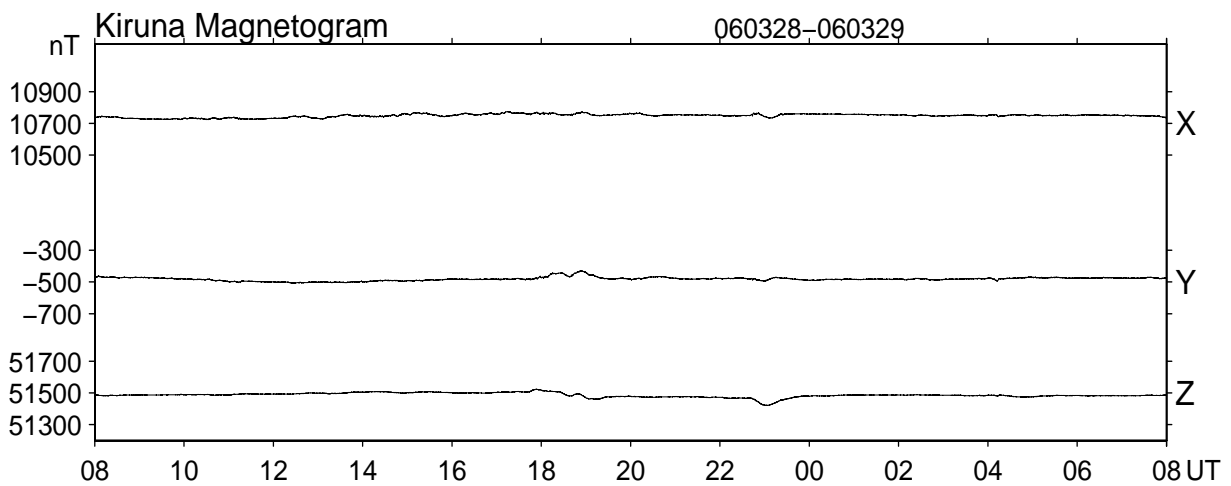
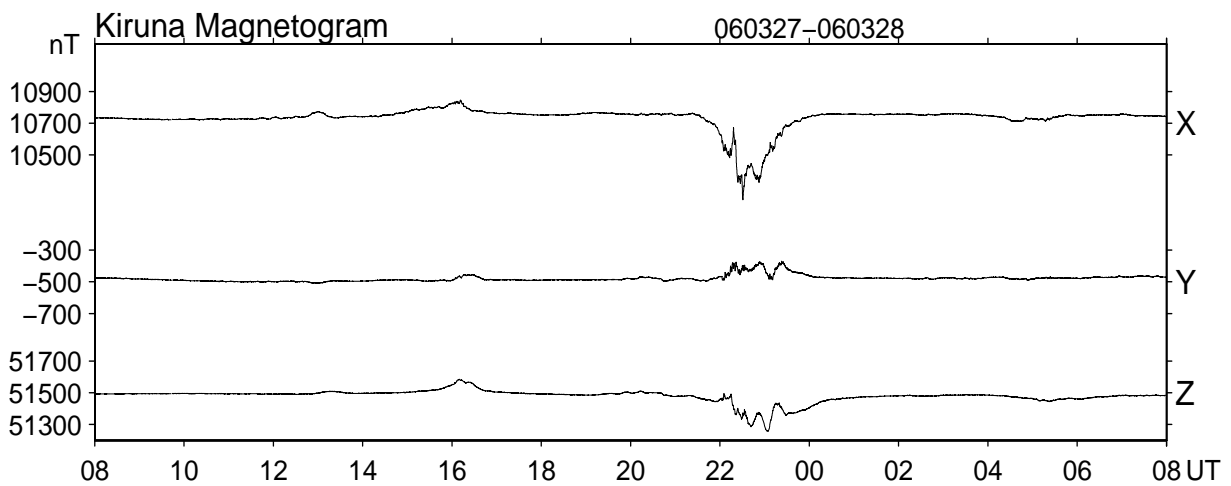
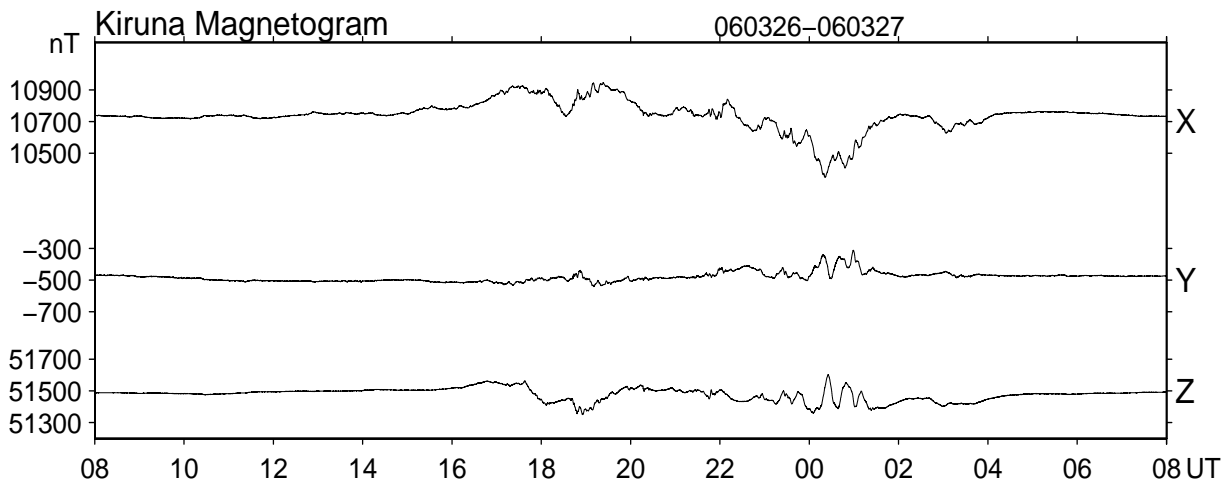
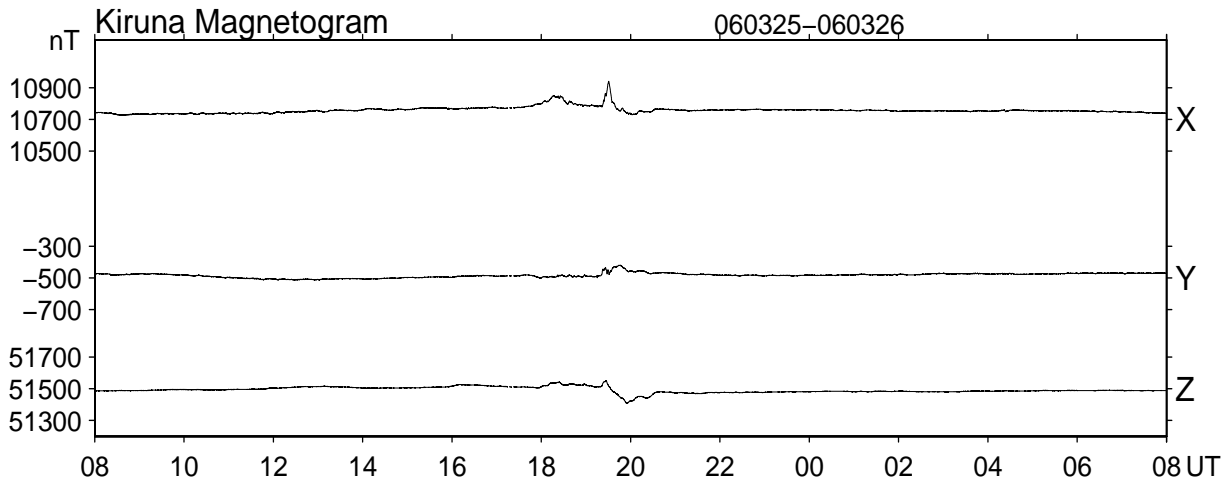


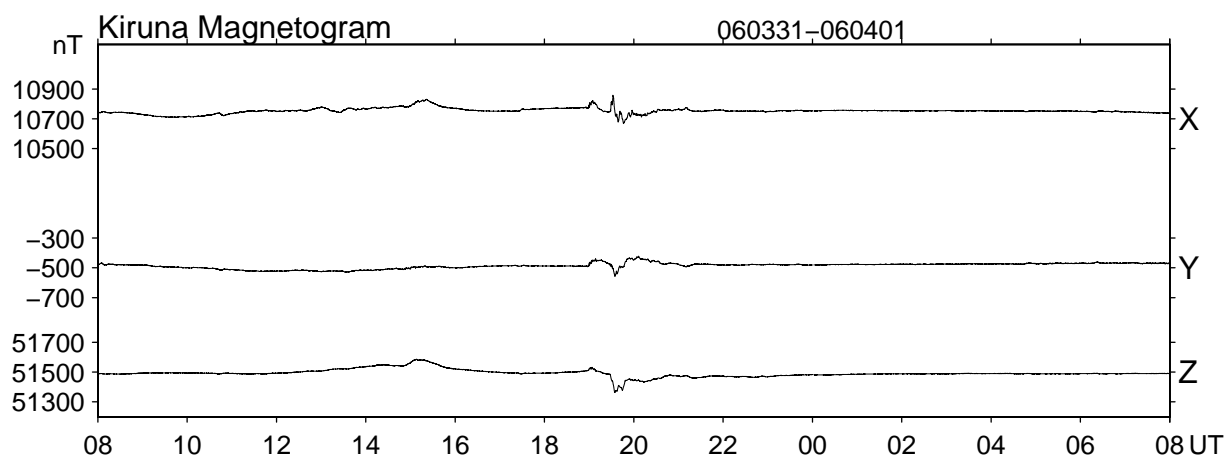
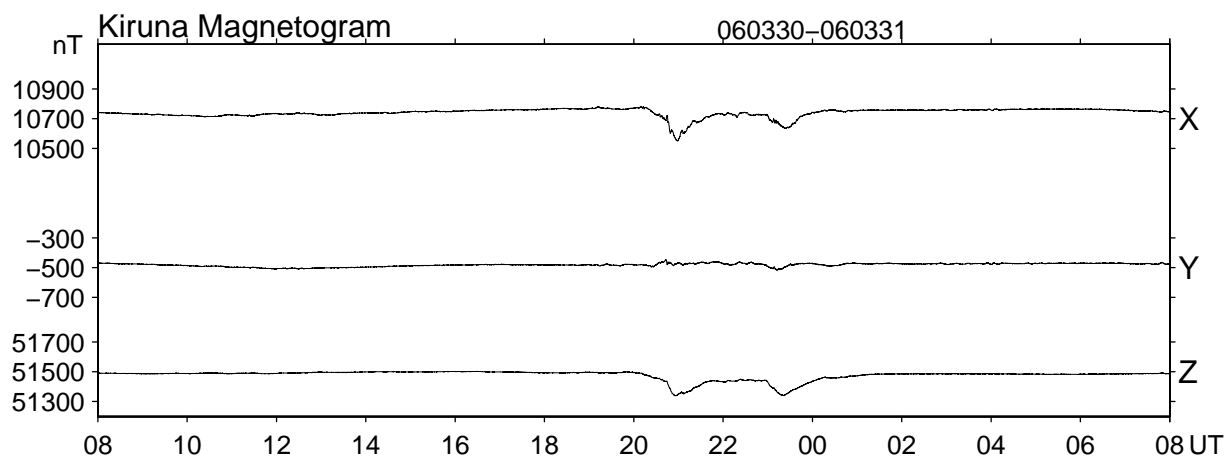
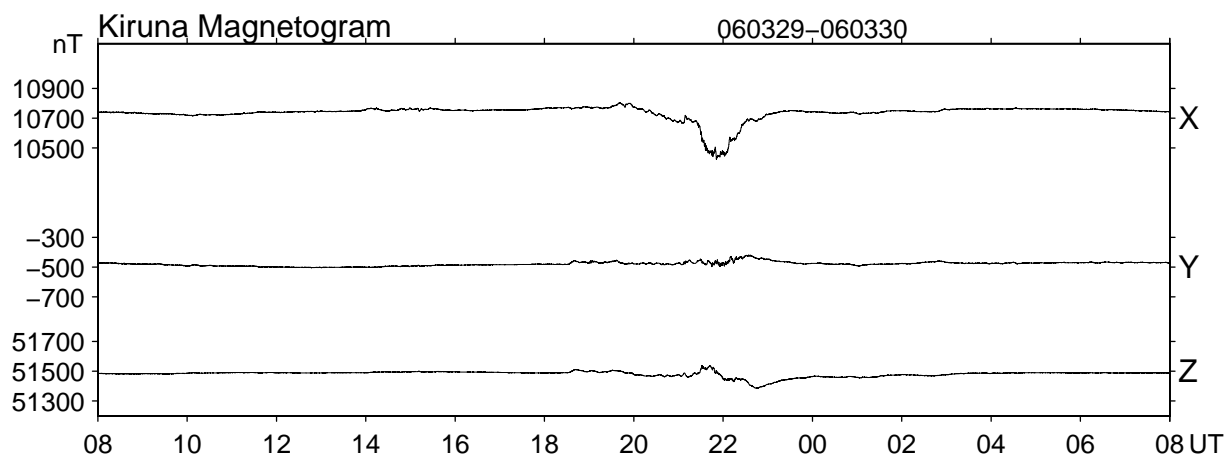
















# **KIRUNA VERTICAL RIOMETER**

January - March 2006

<http://www.irf.se/~rio>

Riometer plots are not available in this issue since the riometer had severe technical problems during the period it covers. Riometer plots will appear again from the 2010 January-March issue.







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