

Met Office Space Weather Operations and R&D

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Met Office Motivation?





2010

- Increasing interest within UK Government
 - Max coinciding with Olympics
- Space Weather placed on NRR
- NOAA approached Met Office
 - Met Office become back-up to SWPC
- Desire to demonstrate increased value of 24x7 operations



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Identifying the risk

- Royal Academy of Engineering report
- Extreme space weather Carrington event
 - Electricity
 - ~13 transformers damaged in UK
 - 2 coastal nodes could experience disconnection
 - Possible short blackouts in urban areas
 - Impact on Satellites, Aviation, GNSS, comms also assessed
 - Need for further technical and forecast mitigation



Extreme space weather: impacts on engineered systems and infrastructure





Met Office Space Weather Operations Centre

- Embedded in Met Office Hazard Centre
 - 24x7x365 29 April' 14 🗸
 - Full capability autumn '14
 - New member of ISES
- Collaborate with academia not replicate
- Operational collaboration with NOAA SWPC & BGS
 - Daily forecast coordination
- Add UK-centric advice and impacts





NOAA impact scales

Met Office

Cat	egory	UK Effect	US and Global Effect	Physical measure	Average Frequency (1 cycle = 11 years)
Scale	Descriptor	Duration of event will inf	luence severity of effects		
		Geomagnetic Storms	5	Kp values*	Number of storm events when Kp level was met; (number of storm days)
G 5	Extreme	 Power systems: Localised voltage control and protective system problems may occur. Spacecraft operations: may experience extensive surface charging, problems with orientation, uplink/downlink and tracking satellites. Other systems: HF (high frequency) radio propagation may be impossible in many areas for one to two days, GPS satellite navigation may be degraded for days, low-frequency radio navigation can be out for hours, and aurora has been seen as low as 40° geomagnetic lat. 	 Power systems:widespread voltage control problems and protective system problems can Spacecraft operations: may experience extensive surface charging, problems with orientation, uplink/downlink and tracking satellites. Other systems: pipeline currents can reach hundreds of amps, HF (high frequency) radio propagation may be impossible in many areas for one to two days, satellite navigation may be degraded for days, low-frequency radio navigation can be out for hours, and aurora has been seen as low as Florida and southern Texas (typically 40° geomagnetic lat.)**. 	Кр = 9	4 per cycle (4 days per cycle)
G 4	Severe	 Power systems: No impact on UK power grid. Spacecraft operations: may experience surface charging and tracking problems, corrections may be needed for orientation problems. Other systems: HF radio propagation sporadic, GPS satellite navigation degraded for hours, low-frequency radio navigation disrupted, and aurora has been seen as low as 45° geomagnetic lat. 	 Power systems: possible widespread voltage control problems and some protective systems will Spacecraft operations: may experience surface charging and tracking problems, corrections may be needed for orientation problems. Other systems: induced pipeline currents affect preventive measures, HF radio propagation sporadic, satellite navigation degraded for hours, low-frequency radio navigation disrupted, and aurora has been seen as low as Alabama and northern California (typically 45° geomagnetic lat.)**. 	Kp = 8, including a 9-	100 per cycle (60 days per cycle)
		Other systems: migratory animals are affected at this and higher levisible at high latitudes (northern Michigan and Maine)**.	evels; aurora is commonly		



Space-Weather-Forecast¶ Space-Weather-Forecast¶

Space-Weather-Forecast¶

Space-Weather-Forecast¶

Issued on Tuesday, 14 January 2014 at 00:50 UTC

This guidance document provides a four day assessment of Space Weather events. This is a non-scientificdocument issued daily at 01:00 UTC but it will be updated if there are significant changes. Stated probabilities are for reaching or exceeding given levels. -For more information about space weather impacts please see http://www.metoffice.gov.uk/spaceweather

Headline: Earth is currently experiencing high speed solar wind but geomagnetic activity is low. A moderate class flare occurred late last night.

Space Weather Activity over the past 24 hours

Solar activity was moderate; the largest flare detected was a moderate flare at 21:51 UTC on the 13th January 2014, emitted from a sunspot which has now rotated onto the western limb. Due toits position it is difficult to classify but it is thought to be a complex sunspot. There are 5 othercategorised sunspots on the visible disc, all smaller and less magnetically complex. Despite highaverage solar wind speeds the magnetic polarity remains positive and geomagnetic activity has been quiet to unsettled. Solar radiation levels remain below S1 level.

Four·Day·Space·Weather·Forecast·Summary¶

Solar activity is expected the remain low, initially with a 45% chance of further moderate class flares and only 10% chance of extreme class flares from the same complex sunspot. The risk of large flares falls further from Day 2 when this sunspot rotates off the western limb. Any further significant flares from this sunspot even after it rotates off the visible disc may lead to a solarradiation storm. The probability remains relatively high at 40% on Day 1 and 30% on Day 2, but decreases after that. Geomagnetic activity will remain unsettled on Day 1, but will be quiet from Day 2 onwards as the influence of the solar wind diminishes.

-Page Break -----...¶

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Geomagnetic Storms

Geo-Magnetic∙ Storm¤ Probability⊷ (Exceedance)¤	Level¤	Past•24₊- Hours₊- (Yes•or•No)¤	Day•1⊷ (00-24 UTC), ↔ (%)¤	Day•2⊷ (00-24 UTC)• (%)¤	Day-3⊷ (00-24 UTC)• (%)¤	Day-4₊- (00-24 UTC) ↓ (%)¤
Minor-or- Moderate¤	G1·to·G2¤	No¤	30¤	20¤	1 0¤	1¤
Strong¤	G3¤	No¤	05¤	01¤	01¤	<mark>01</mark> ¤
Severe¤	G4¤	No¤	01¤	01¤	01¤	<mark>01</mark> ¤
Extreme¤	G5¤	No¤	01¤	01¤	01¤	<mark>01</mark> ¤

Radio Blackouts - X Ray Flares

X-Ray-Flares¤ Probability+- (Exceedance)¤	Level¤	Past-24₊ Hours₊- (Yes/No)¤	Day·1⊷ (00-24· UTC)⊷ ⊷ (%)¤	Day•2⊷ (00-24• UTC)⊷ ⊷' (%)¤	Day·3₊- (00-24· UTC)₊- ←' (%)¤	Day-4⊷ (00-24∙ UTC)⊷ ↔ (%)¤
Active¤	R1-R2+-∕ M·Classo	No¤	45¤	10¤	10¤	10¤
Very -Active¤	R3-to-R5+ X-Classe	No¤	10 ¤	01¤	01¤	01¤

Solar Radiation Storms - (High Energy Protons): ¶.

Radiation∙ Storms¤ Probability₊J (Exceedance)¤	Level+- (cm-2sr-1s- 1·)¤	Past•24⊷ Hours⊷ (Yes•or•No)¤	Day·1⊷ (00-24 UTC)• ⊷ (%)¤	Day-2⊷ (00-24 UTC), ⊷ (%)¤	Day-3⊷ (00-24 UTC)• ⊷ (%)¤	Day•4⊷ (00-24 UTC)• ≁┘ (%)¤	0
Active¤	≥·\$2·*¤	No¤	40 ¤	30¤	1¤	1¤	C
VeryActive	≥·\$3.*¤	No¤	10¤	5¤	1¤	1 ¤	C
* CO > 10 MoV/	aratana > 100 r	fu C2 > 40 Ma	/> 1000 pfu or	diars E0 Ma	/ > 10 pfu / pfu		o

 $52 \ge 10 \text{ MeV}$ protons $\ge 100 \text{ ptu}.53 \ge 10 \text{ MeV} \ge 1000 \text{ ptu}$ and/or $\ge 50 \text{ MeV} \ge 10 \text{ ptu}$. (ptu = cm sr is)9 ¶



Met Office	Spac
Energy	S

FORECASTER OVERVIEV

Moderate Radio Black M-class flare likely over 1 ACE at 2130 UTC possibl

Solar activity is expected to large sunspot groups AR20[±] become more geo-effective start ACTIVE but then becor equatorial small coronal ho thought to be minimal. With view a proton event seems background values.

SOLAR TIMELAPSE

Space Weather Energy

FORECASTER OVERVIEW

Sector forecasts 🗸

Moderate Radio Blackout observed this morning. Further M-class flare likely over the next few days. Sudden impulse at ACE at 2130 UTC possible CME from 3rd May.

Solar activity is expected to remain moderate over the coming days the large sunspot groups AR2055 and AR2056 rotate around the disc and become more geo-effective. Geomagnetic activity is expected to stay start ACTUF but then become generally QUIET. There are a couple equatorial small coronal holes visible on the disc but their impact is thought to be minimal. With AR2051 now rotated completely out of view a proton event seems unlikely now and electrons should stay a background values.

VARNINGS AND AL	LERIS	
	Active alerts	▲ Warnings
Geomagnetic		-
Radio blackout		R1 21:00 20/05 03:00 21/05
Proton flux > 100 MeV	S1 11:23-now	
Proton flux > 10 MeV		
Кр	-	
Kuk	-	
Electrons		

Issued 8 May 2014 at 12:00

GEOMAGNETIC STORM FORECAST

Probabilities of geomagnetic storms Probability Past 24 h. Level Day 1 Day 2 Day 3 Day 4 Minor or 65 05 G1 to G2 No 05 05 moderate 45 Strong G3 No 01 01 01 G4 No 01 01 01 01 Severe Extreme G5 No 01 01 01 01 96 96 96 96

2014-05-07 10:00:00

No Earth directed CMEs have been observed. Maximum solar wind

Issued 8 May 2014 at 12:00

Geomagnetic commentary

The largest flare over the last 24 hours was M1.2 flare at 16:29 UTC from AR2051 which has revolved around the western limb. With one beta-gamma spot (AR2055) and one beta-gamma-delta spot (AR2056) there is still a risk of further M-class flare activity over the next few days.

Issued 8 May 2014 at 12:00



ACE MAGNETOMETER



BGS 3-HOURLY KP INDEX



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SDO/AIA 193 2014-05-20 0

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speed is 500 km/s

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Space weather models at the Met Office

- Sun to Earth: WSA Enlil
- At the Sun: CME analysis tools
- Van Allen Belts (1,000-60,000km): SPACECAST
- Geosynchronous orbit (~35,000km): REFM
- Ionosphere (50-600km): MIDAS
- Ionosphere (D-region, 50-90km): D-RAP



• **Purpose:** Models solar wind speed & density from Sun to Earth (IMF modelled but no Bz input). Predicts CME arrival times at Earth,

• Inputs:

- WSA uses solar magnetograms to model coronal magnetic field and provide inner BCs for Enlil.
- CME analysis tools (NASA StereoCAT or NOAA CAT) calculate CME parameters using fit to STEREO A&B and LASCO images.
- CME parameters input into Enlil
- Output: Run every 2hrs
- Forecasts: average error: +_ 7 hrs (Doug Biesecker); lead time: CME transit time – a few hrs









DRAP gives real-time prediction of the absorption (radio propagation) conditions in the D-region.



REFM provides daily electron fluence forecasts at GEO orbit.



FP7 project provides forecasts of particle radiation to aid satellite operation.

Ionospheric analyses of European sector produced using **MIDAS**.

Oct 30 2003, 20:00UT



- Currently planning archiving of data for verification.
- REFM is currently compared with GOES fluence data on internal webpages.
- Plan to compare Enlil density & velocity with STEREO A, STEREO B & ACE.
- Metrics tables to show that forecasters are adding value to models.



Modelling from Sun to Earth: Current R&D Status Met Office Goal: Coupled Sun to Earth models with DA for much enhanced forecast capacity Ionosphere and atmospheric Solar wind model Observed photospheric M Owens (Reading Uni) models (e.g., TING) (e.g., Enlil) magnetic field Magnetospheric mode coronal model (e.g., LFM) (e.g., MAS) Solar wind Real time regional TEC WSA Enlil persistence model (MIDAS) Ensembles for benchmarking / validation (Owens et Thermosphere: al, 2013; Kohutova) DA and Whole Improved methods of ionosphere atmosphere

coupling

model plans

tracking solar wind features (Tucker-Hood)





- Met Office Space Weather services available and being developed - 24/7 forecasting as of April.
- Associated R&D programme shall lead to pullthrough of improved services and operational forecasts.
- Large challenges call for interdisciplinary collaborations which utilise the skills of partners to maximum effect.



Questions and answers

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Met Office Space Weather Research & Development - Extra Slides



Other solar wind modelling

- solar wind persistence model WSA Enlil benchmark and quick forecast tool
- CME drag model ensemble quick heads up of CME arrival times for forecasters



RMS error of various model predictions of solar wind speed, relative to a reference model (Owens et al, 2013)



Active Region Tracking





(displayed on Udalpur H-alpha at 2014 Mar 20 0900 UTC)

Detected Filaments

- KSWC Automatic Solar Synoptic Analyser:
 - Magnetic classification,
 - Flare forecasting,
 - Coronal hole detection,
 - Filament detection.
- Further development in collaboration with Trinity College Dublin. (eg Solar Monitor Active Region Tracker (SMART; Higgins et al)

S. Murray